



# EXPOSURE



EXoPlanet Origins and Stellar Ultraviolet Regime Explorer

A UV mission to explore the origins,  
habitability, and evolution of exoplanets





# TEAM BLUE



## Science team

- Maria Aquilina - Malta
- Marina Cano Amoros - Germany
- Sara Gasparini - Norway
- Anne-Sophie Herrijgers - The Netherlands
- Vivien Simon - Hungary

## Payload team

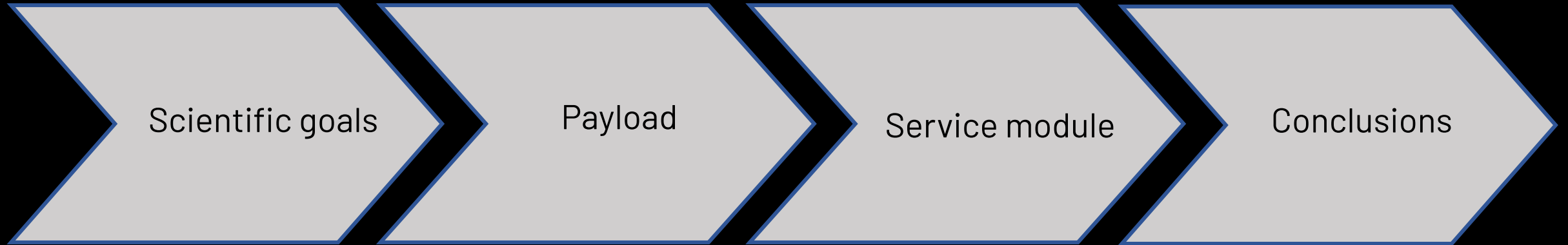
- Vincenzo Davide Cardinale - Italy
- Casper Farret Jentink - Switzerland
- Scott Green - Sweden
- Søren Truelsen - Denmark
- Stephan Zivithal - Austria

## Engineering team

- Patrick Gowran - Ireland
- Isabel Pitz - Germany
- Timo Pospisil - Austria
- Kamil Serafin - Poland
- Gonçalo Trindade - France

**Tutors:** Jerômè Loicq and Greta De Marco

# OVERVIEW



# SCIENCE



Scientific  
goals

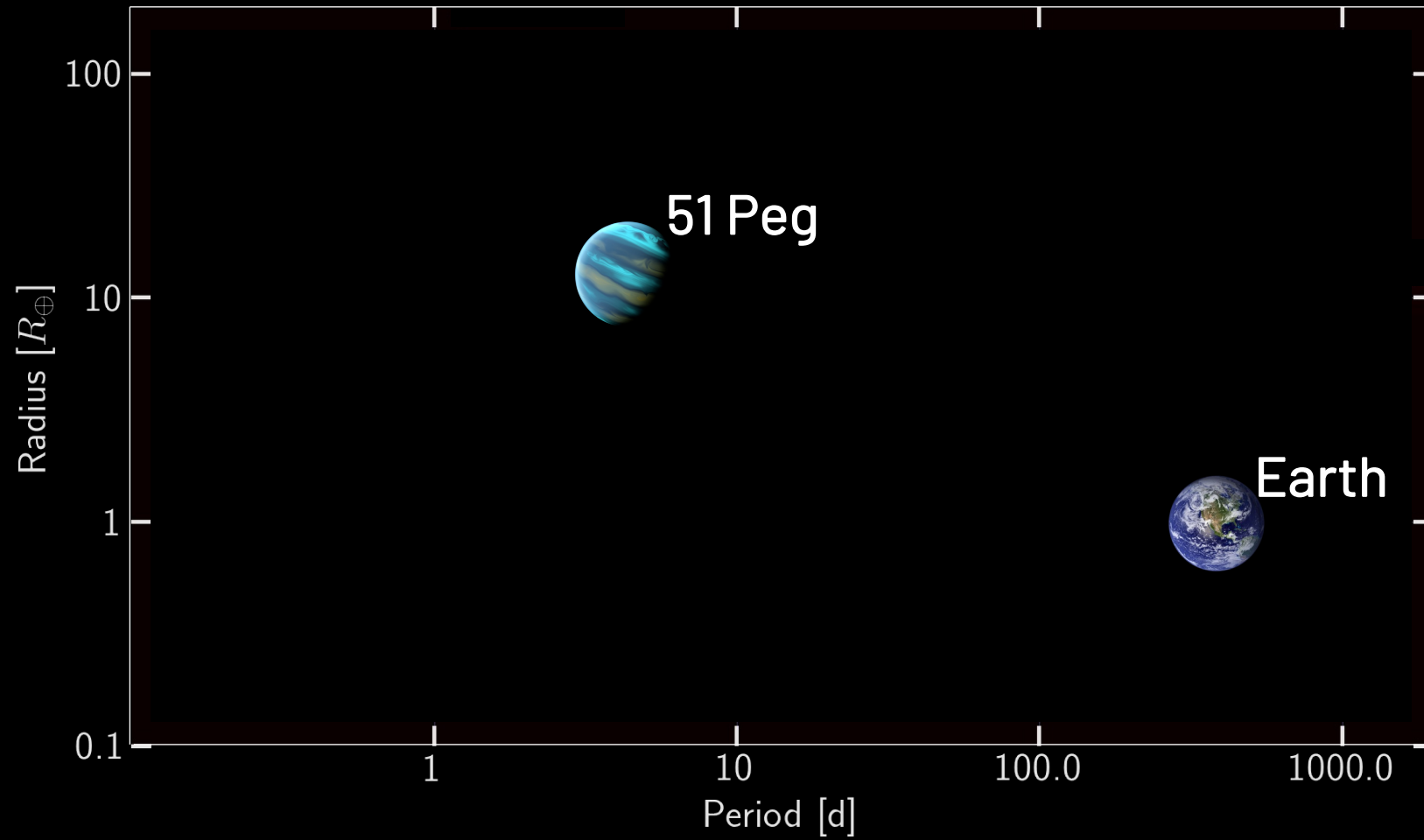
Payload

Service  
module

Conclusions

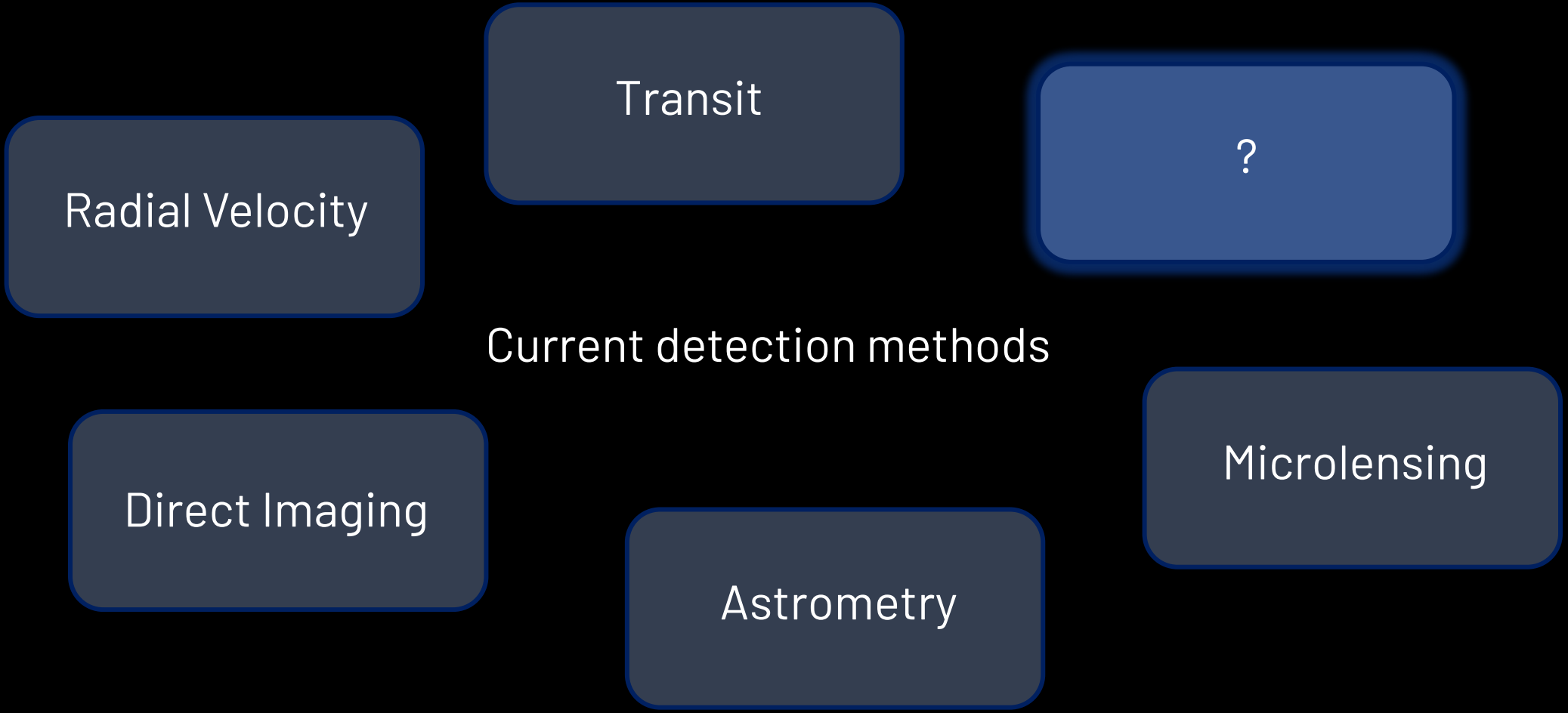


# THE SCIENCE - EXOPLANETS



- Gas giants
- Neptune-like planets
- Rocky planets

# DETECTION METHODS





# GOALS OF THE EXPOSURE MISSION

A large blue planet with dark spots and a few bright stars in the background.

MEASURE  
EXTENDED  
ATMOSPHERES

A large blue planet with dark spots and a few bright stars in the background.

MEASURE  
ATMOSPHERIC  
ESCAPE

A large blue planet with dark spots and a few bright stars in the background.

RECONSTRAIN  
HABITABLE  
ZONE

A large blue planet with dark spots and a few bright stars in the background.

DETECT NEW  
EXOPLANETS



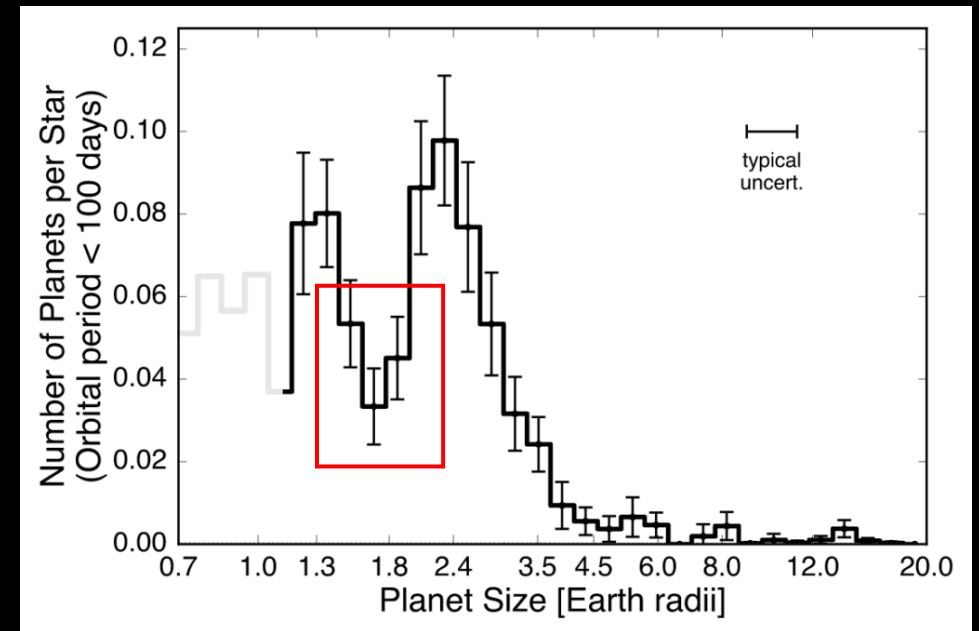
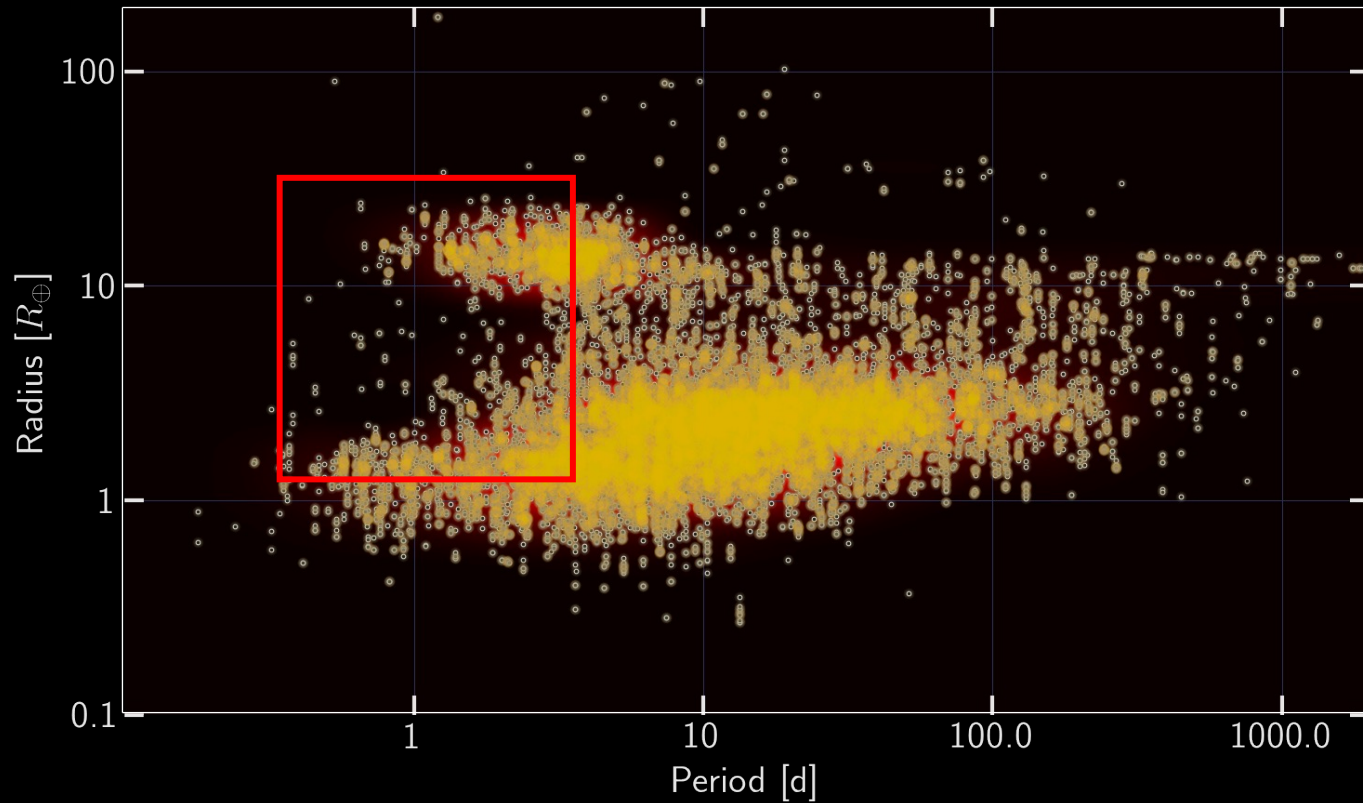
# SCIENTIFIC CASES



- MEASURE EXTENDED ATMOSPHERES
- MEASURE ATMOSPHERIC ESCAPE
- CONSTRAIN HABITABLE ZONE (UV)
- DETECT NEW EXOPLANETS



# EXOPLANET DEMOGRAPHICS



Fulton et al. 2018

# EXTENSION OF H<sub>2</sub> EXOSPHERES

Earth's geocorona captured from the Moon during Apollo 16



Escape rate: 10<sup>3</sup> g/second

**EXOSPHERE ◦  
500 – 10000KM**

THERMOSPHERE ◦  
85 – 500 KM

MESOSPHERE ◦  
50 – 85 KM

H<sub>2</sub> STRATOSPHERE ◦  
10 – 50 KM

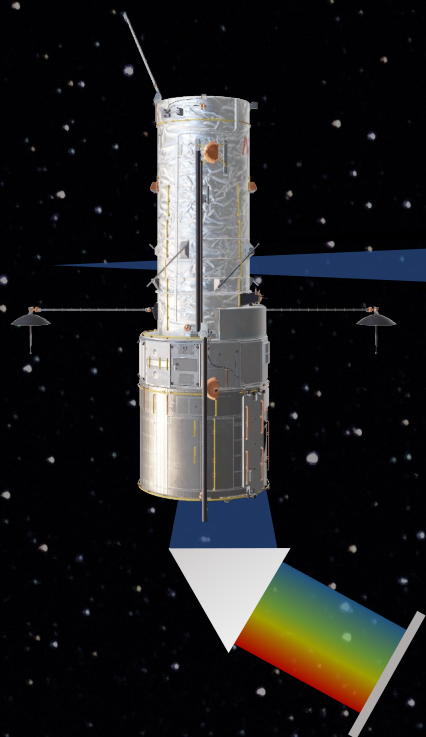
TROPOSPHERE ◦



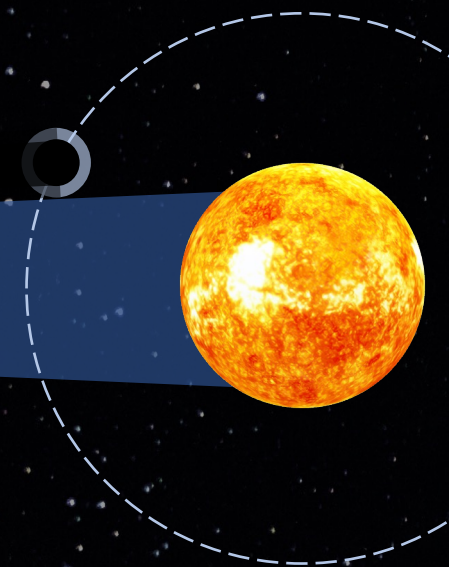


# MEASUREMENT PRINCIPLE (TRANSIT)

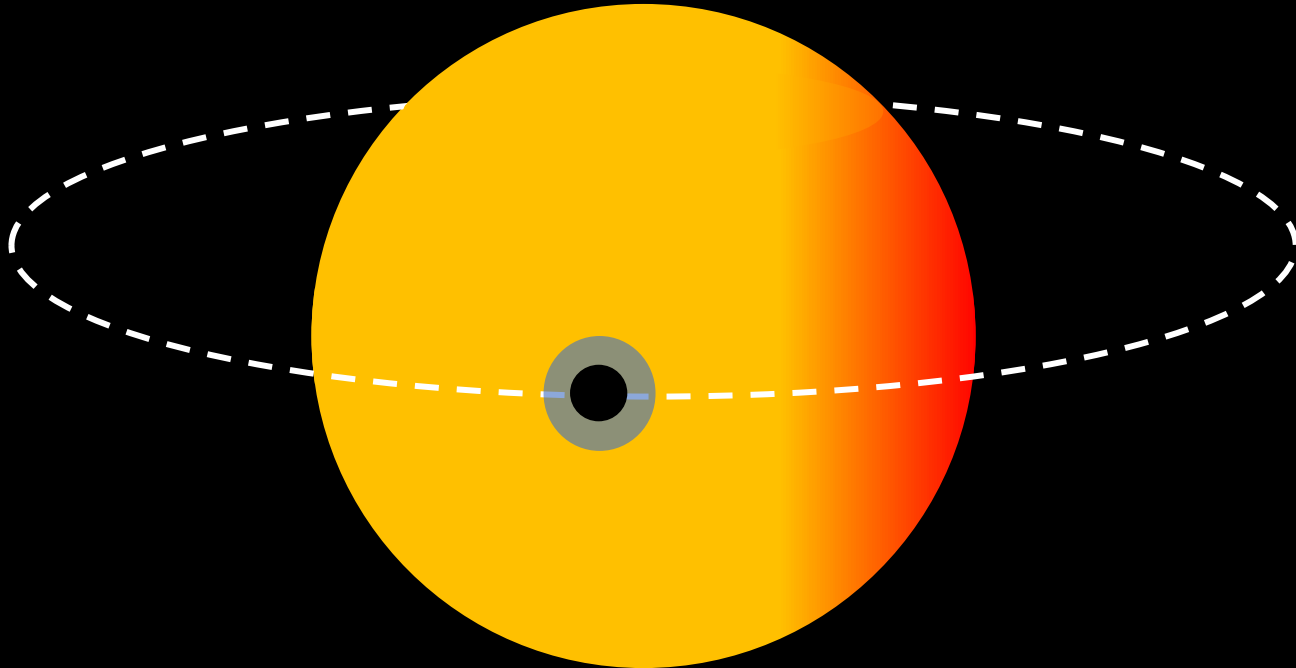
- Transmission spectroscopy in transit
- Absorption of Ly- $\alpha$  in atmosphere



DETECTOR



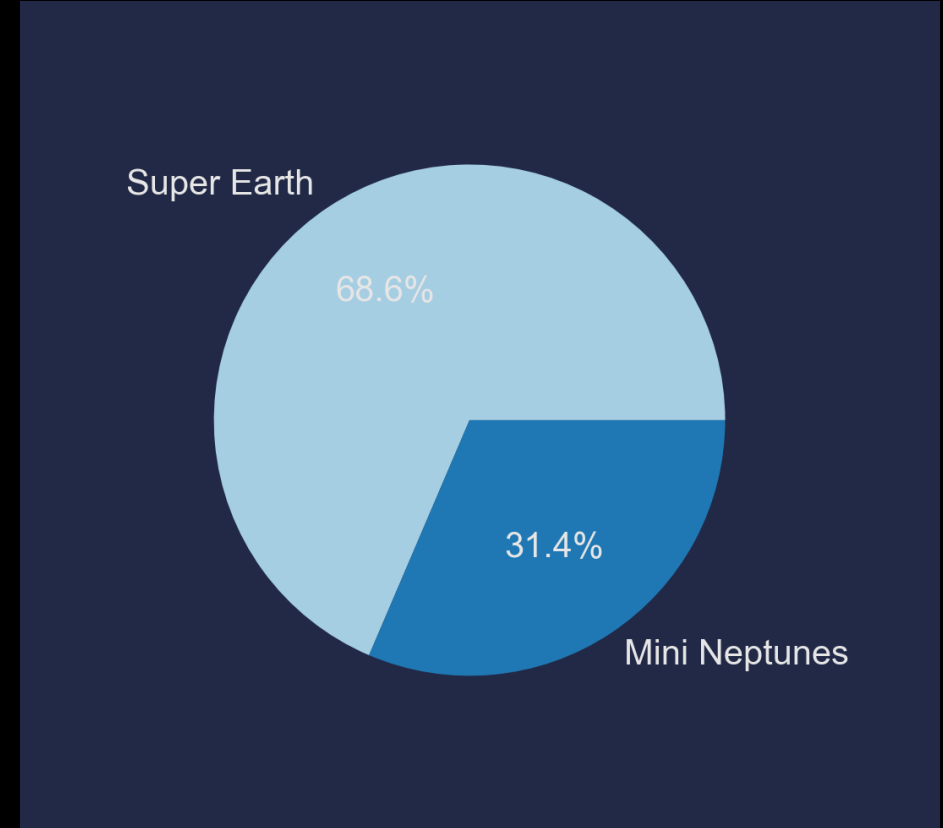
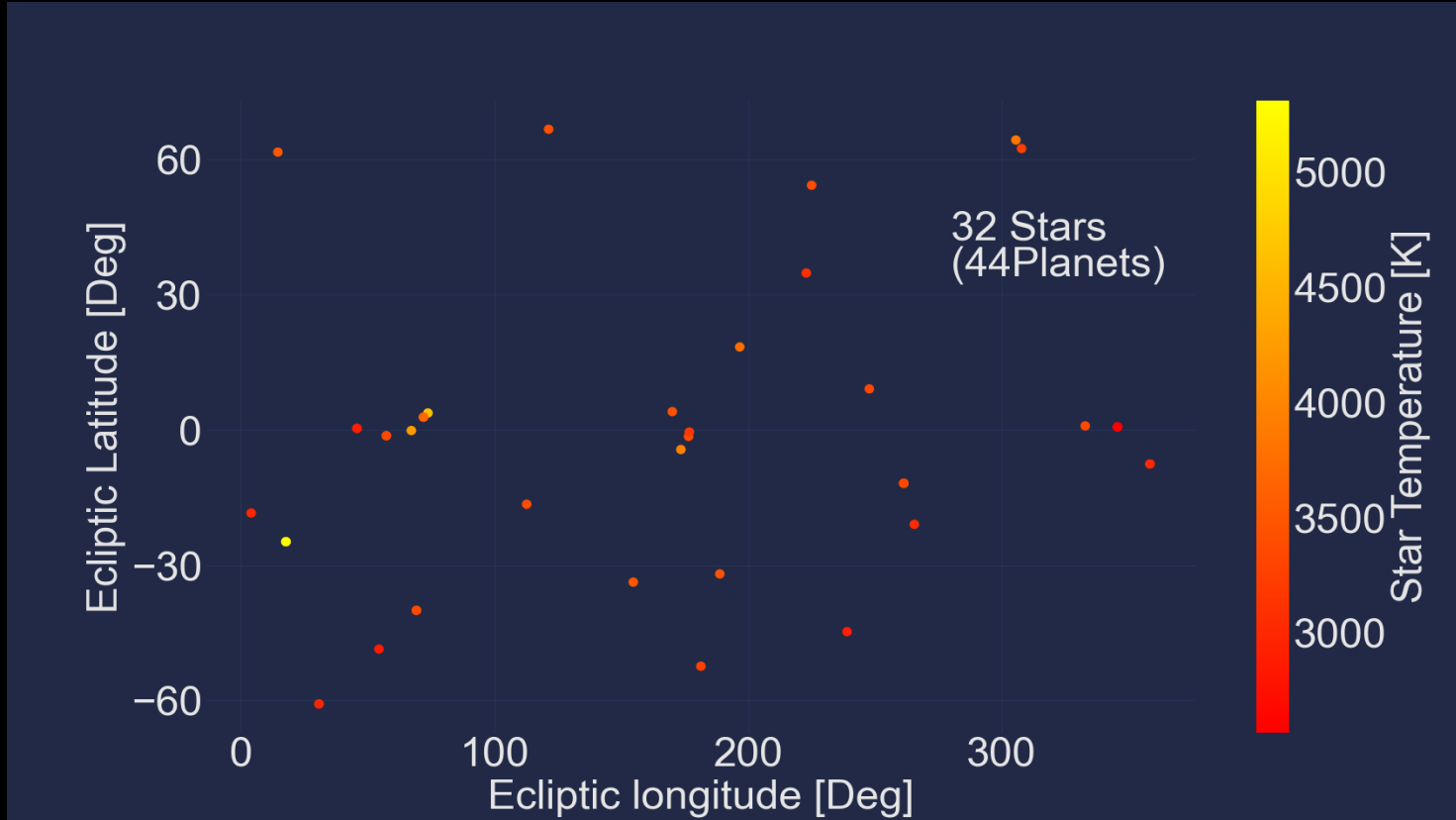
# SENSITIVITY AND REQUIREMENTS



- G, K and M-type stars
- Transit depth of 2.5%
- For a 5-sigma detection  
-> SNR of 200
- Bandwidth 115 -130 nm
- Spec. Resolution 1000



# TARGETS



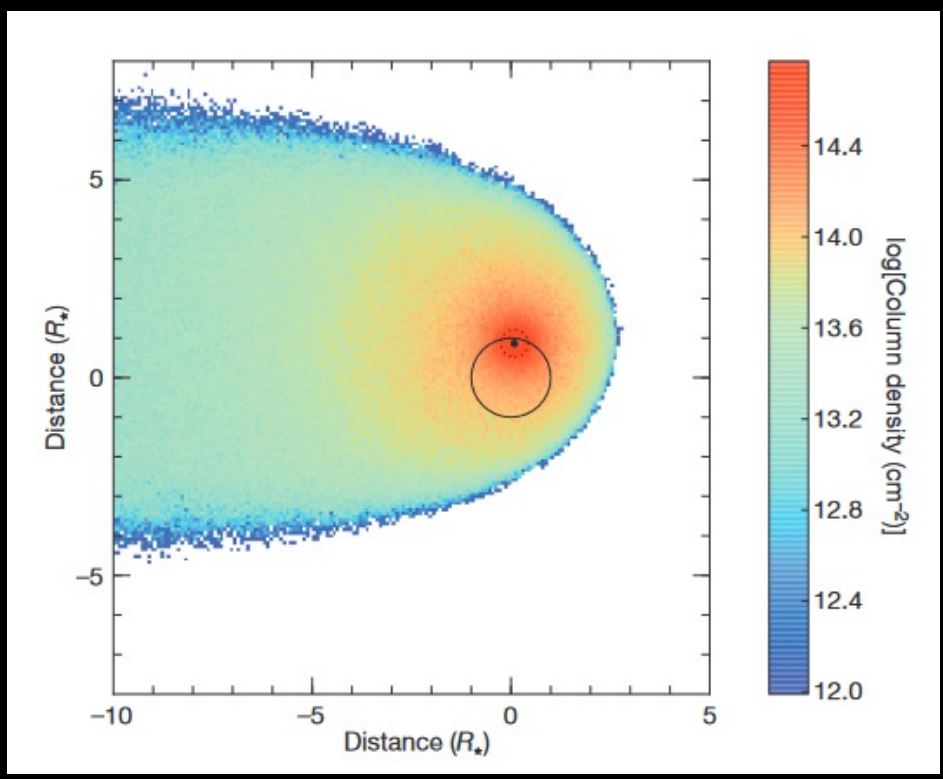
# SCIENTIFIC CASES



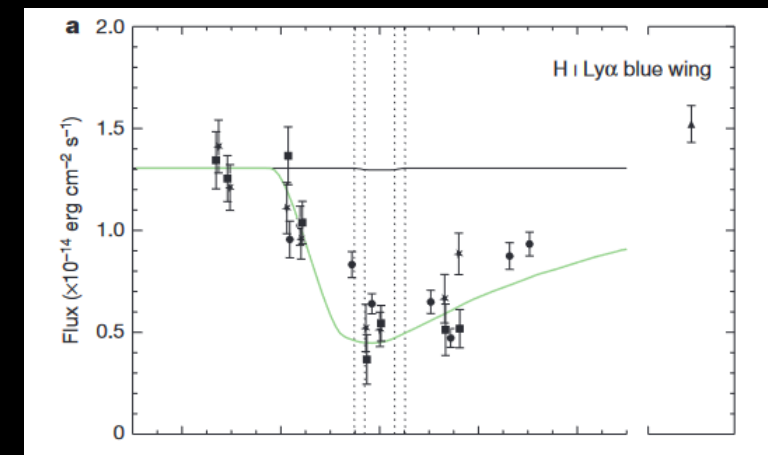
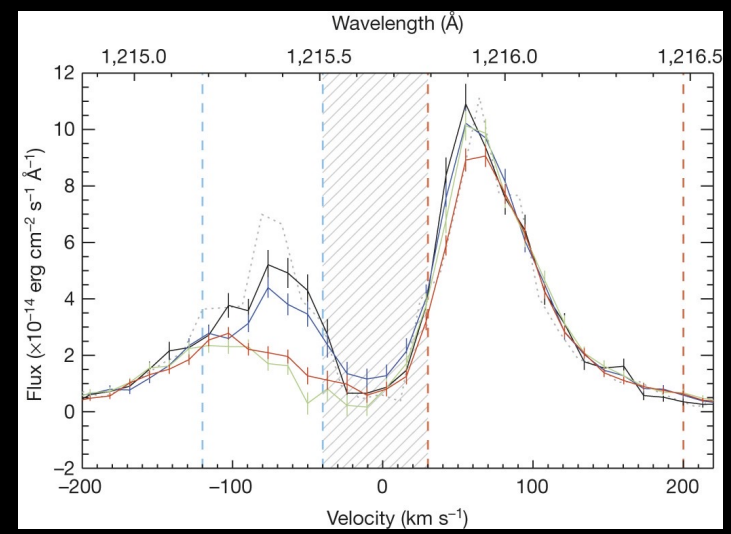
- MEASURE EXTENDED ATMOSPHERES
- MEASURE ATMOSPHERIC ESCAPE
- CONSTRAIN HABITABLE ZONE (UV)
- DETECT NEW EXOPLANETS

# ESCAPING ATMOSPHERES

GJ436b

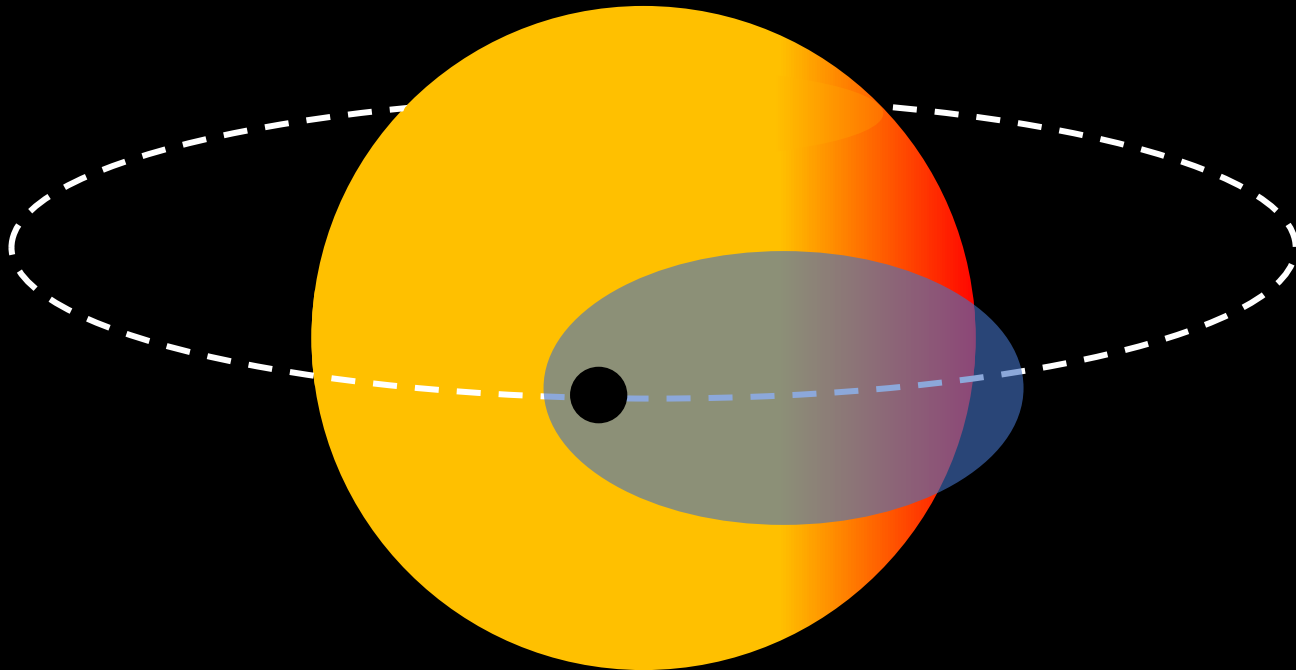


Ehrenreich et al. 2015



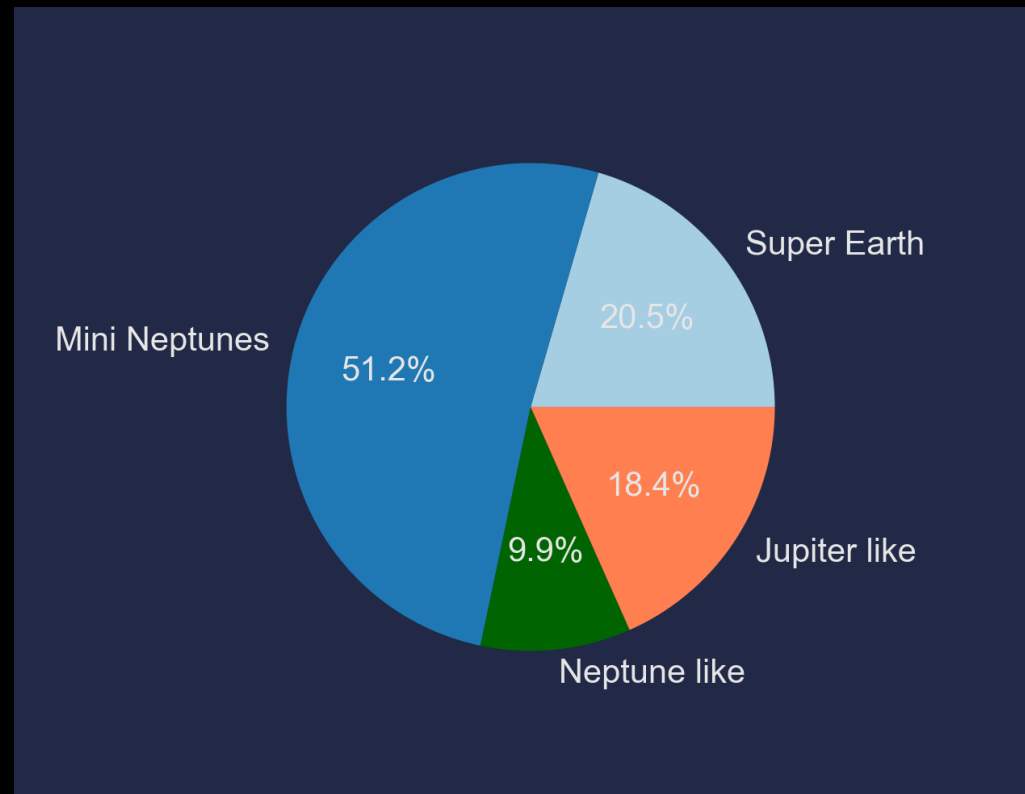
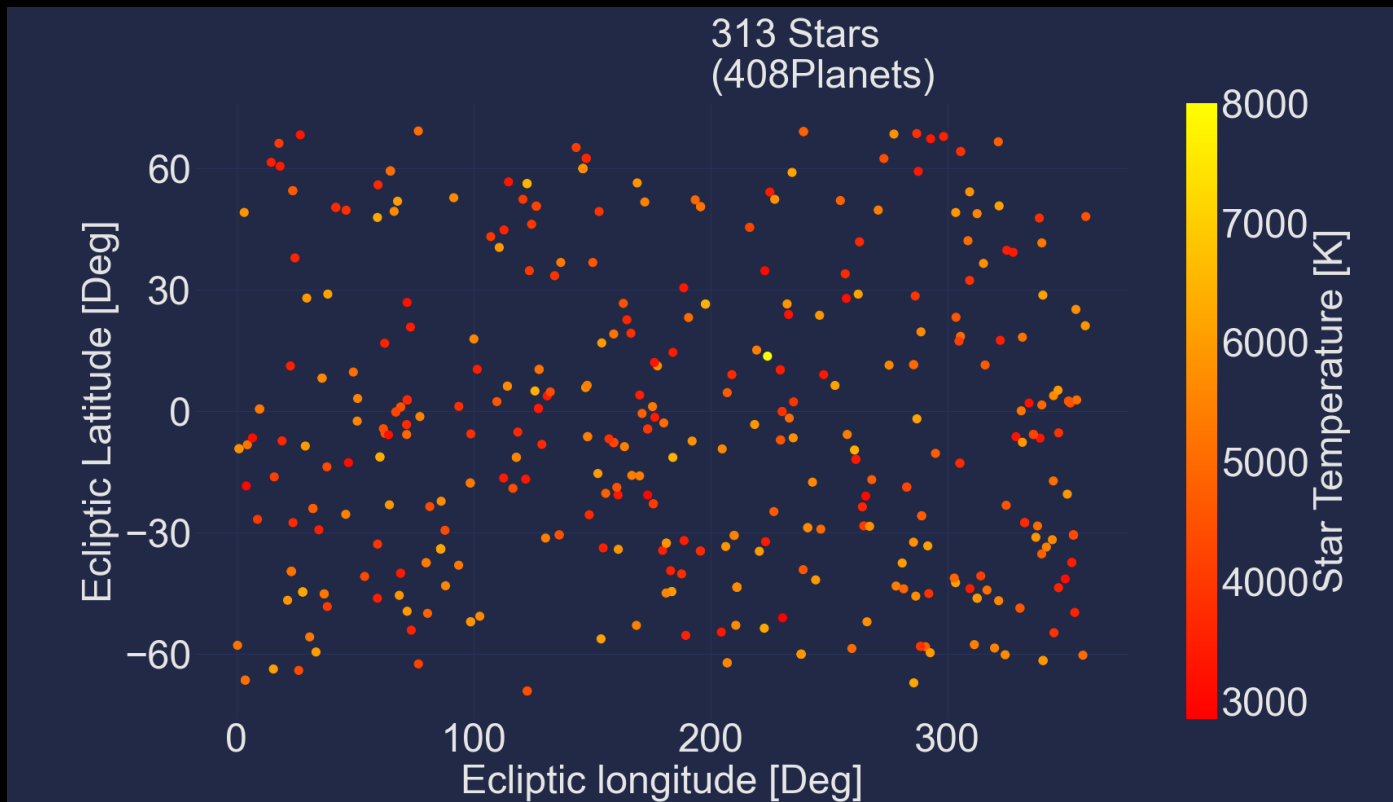


# SENSITIVITY AND REQUIREMENTS



- F, G, K and M-type stars
- Transit depth  $> 50\%$
- For a 5-sigma detection  
-> SNR of 200
- System velocities ( $\pm 100\text{km/s}$ )  
-> bandwidth 119-122 nm (Ly- $\alpha$ )
- Spec. Resolution 20 000
- RV precision: 20km/s

# TARGETS



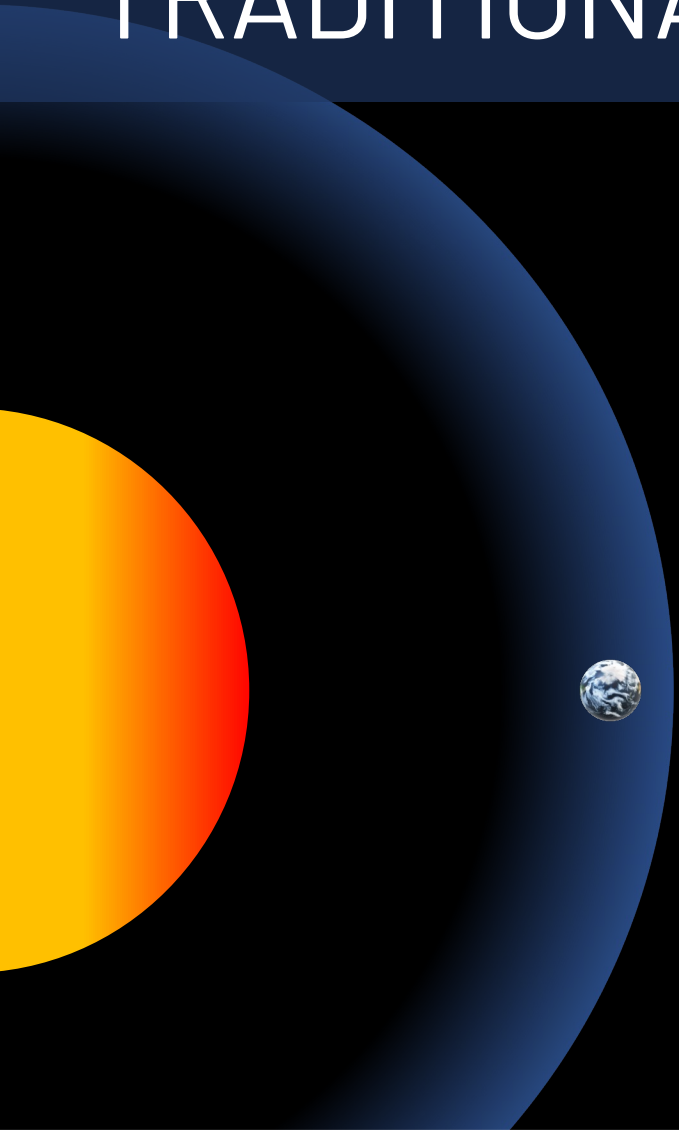


# SCIENTIFIC CASES



- MEASURE EXTENDED ATMOSPHERES
- MEASURE ATMOSPHERIC ESCAPE
- RECONSTRAIN HABITABLE ZONE
- DETECT NEW EXOPLANETS

# TRADITIONAL HABITABLE ZONE

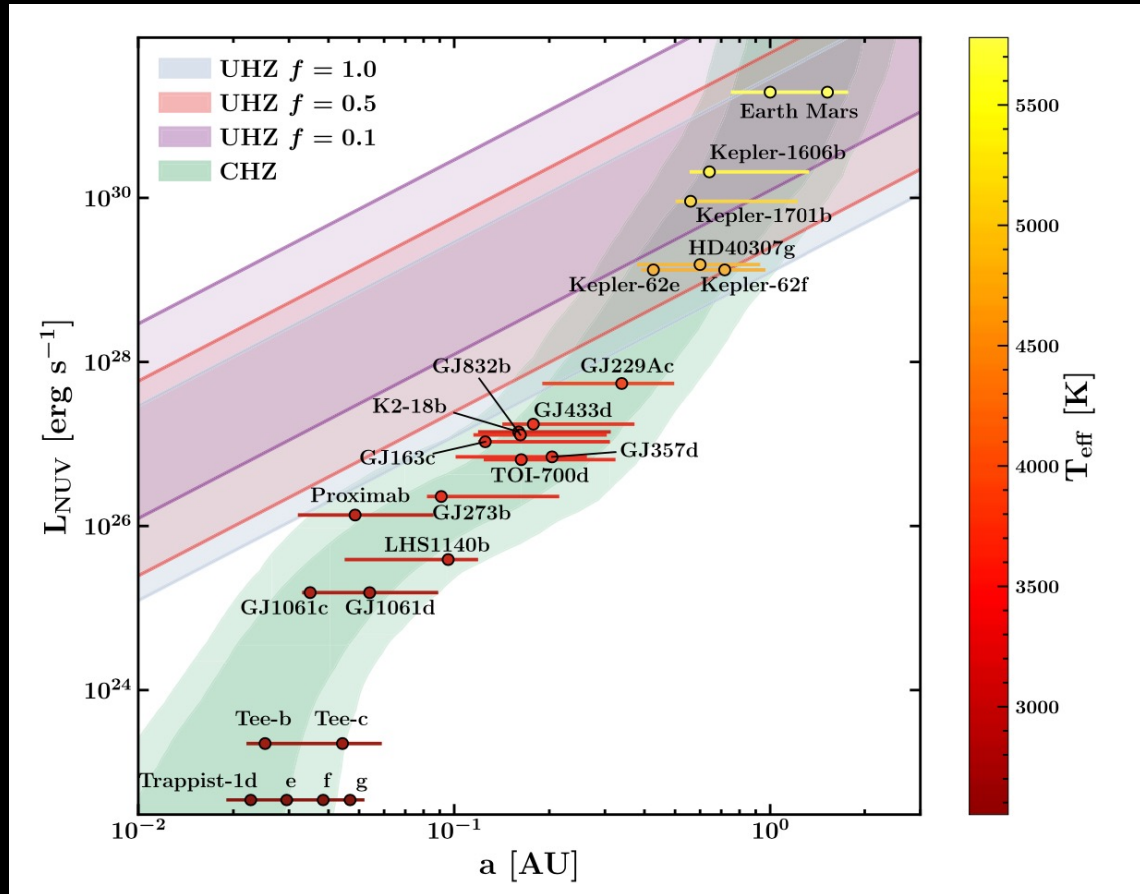
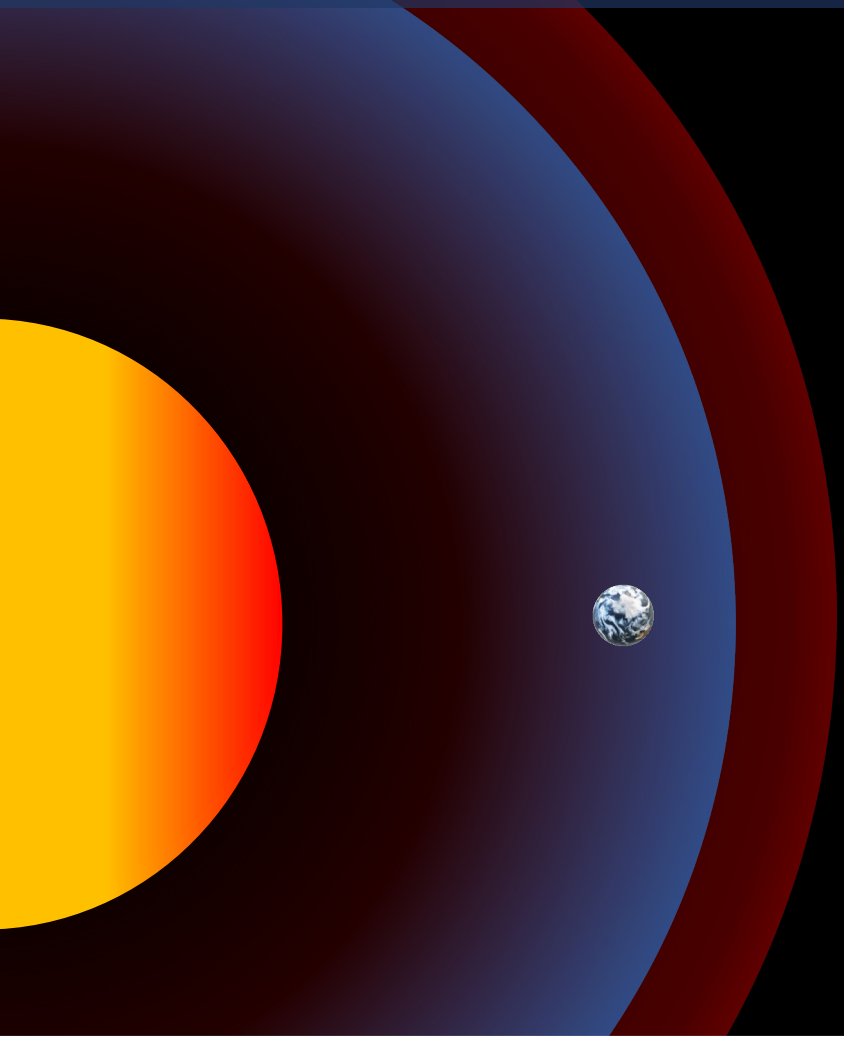


*“...region around a star where the conditions could potentially be suitable to sustain life on a planet within this region, for example allowing the presence of liquid water on its surface”*  
ESA

but...



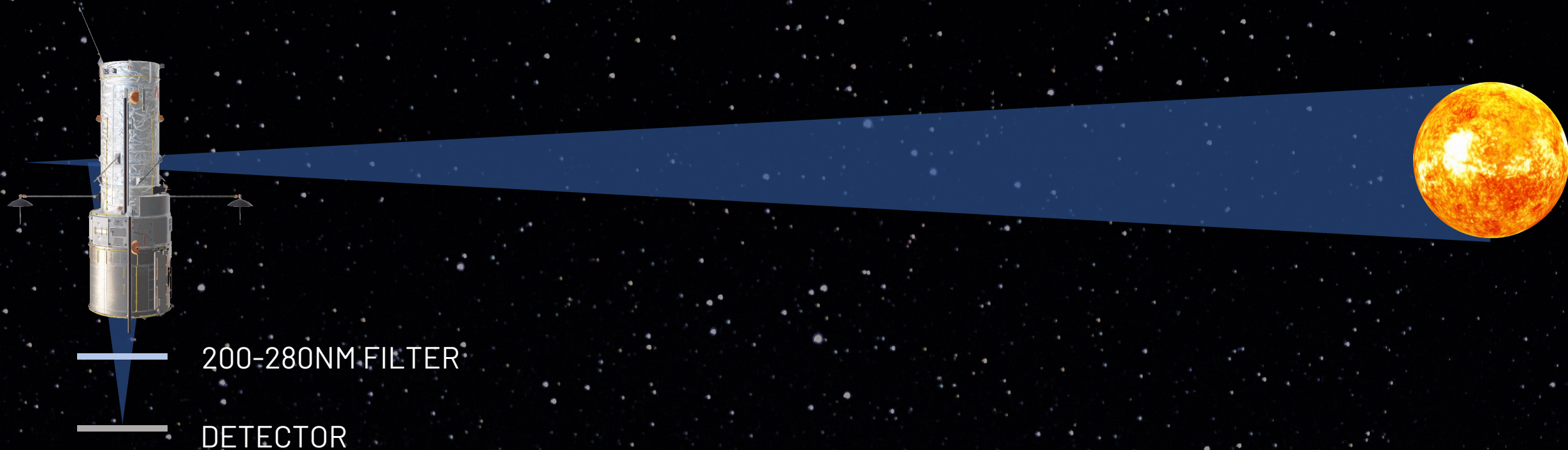
# HABITABLE ZONE RECONSTRAINMENT



Spinelli et al. 2023

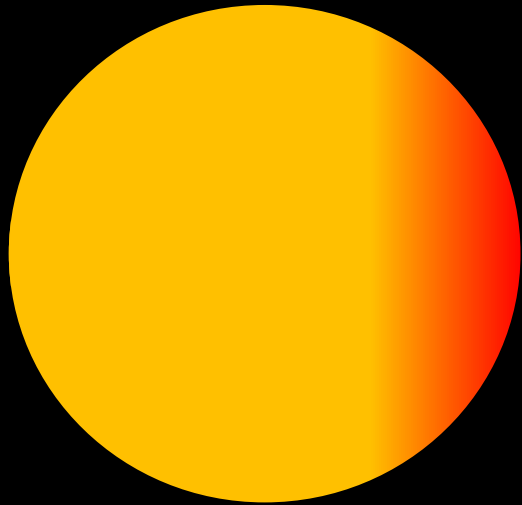
# MEASUREMENT PRINCIPLE (PHOTOMETRY)

- Measuring UV-flux from star



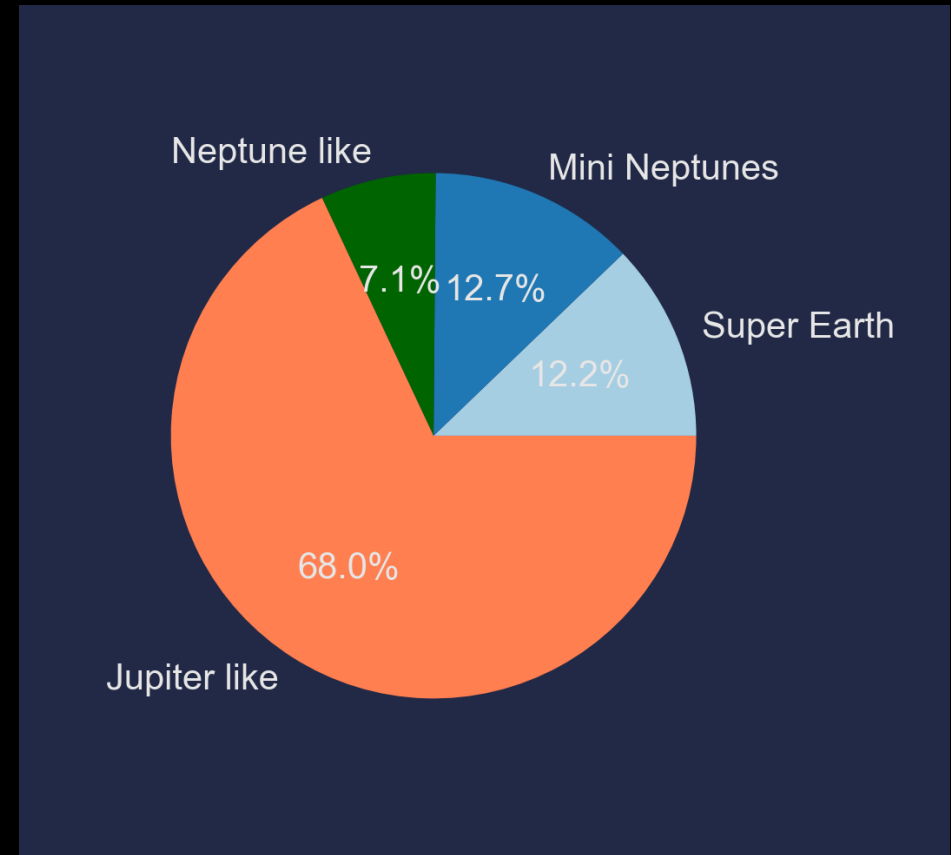
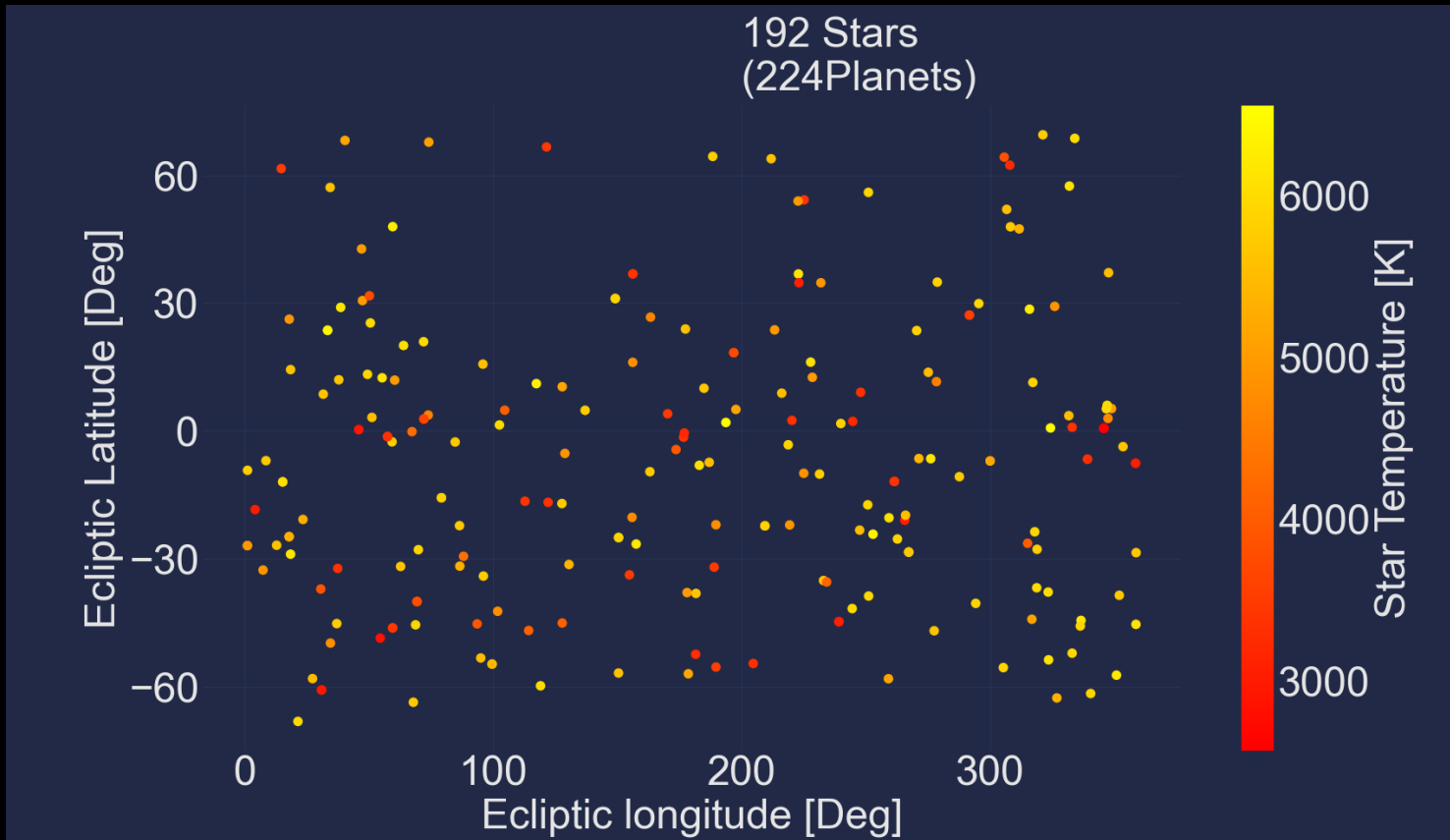


# SENSITIVITY AND REQUIREMENTS



- G, K and M-type stars
- 5% error margin (SNR: 20)
- Bandwidth 200 -280 nm (UV-C)

# TARGETS





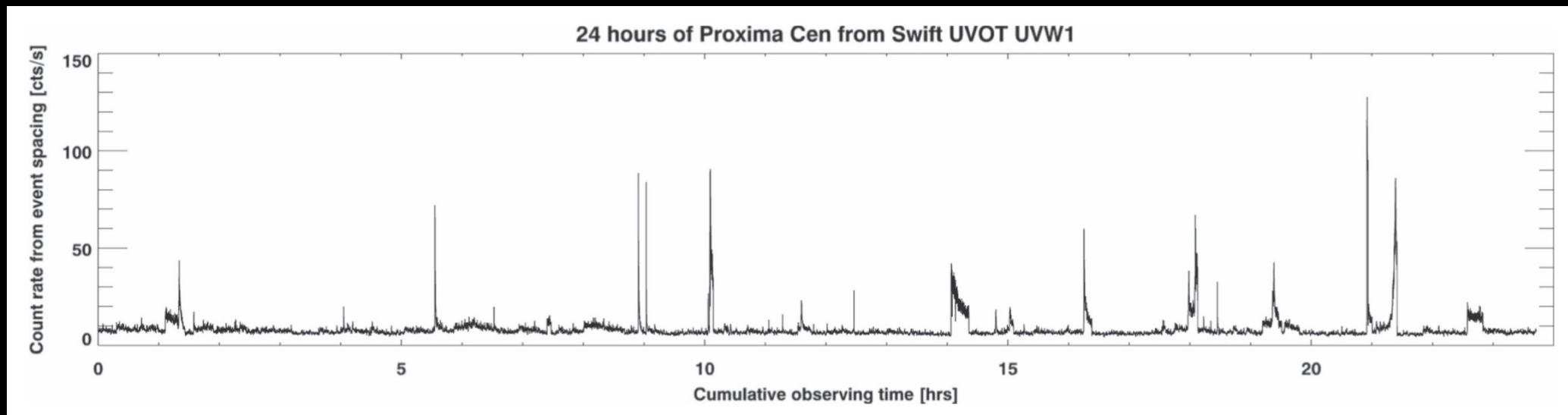
# SCIENTIFIC CASES



- MEASURE EXTENDED ATMOSPHERES
- MEASURE ATMOSPHERIC ESCAPE
- CONSTRAIN HABITABLE ZONE (UV)
- DETECT NEW EXOPLANETS

# LIGHT ECHO PLANET DETECTION

- A stellar flare can brighten a planet in orbit around its host star, producing a light curve with a faint echo (Mann et al. 2018)
- A planet's light-echo emission can potentially be discriminated from that of the host star by means of a time delay (Sparks et al. 2018)

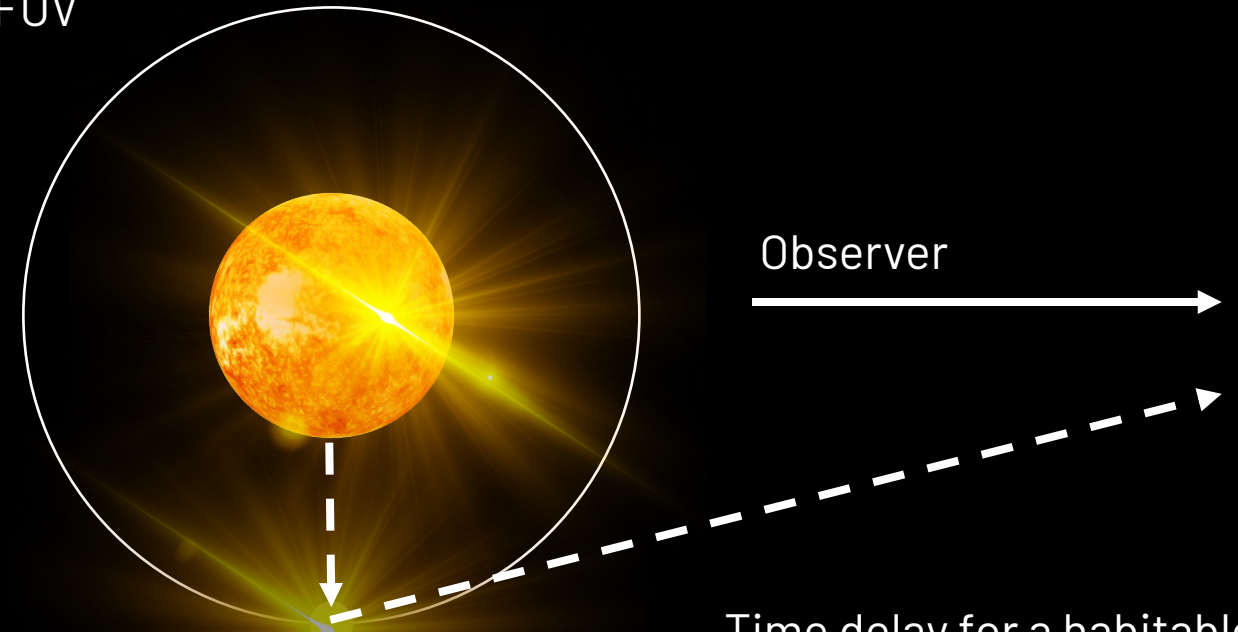
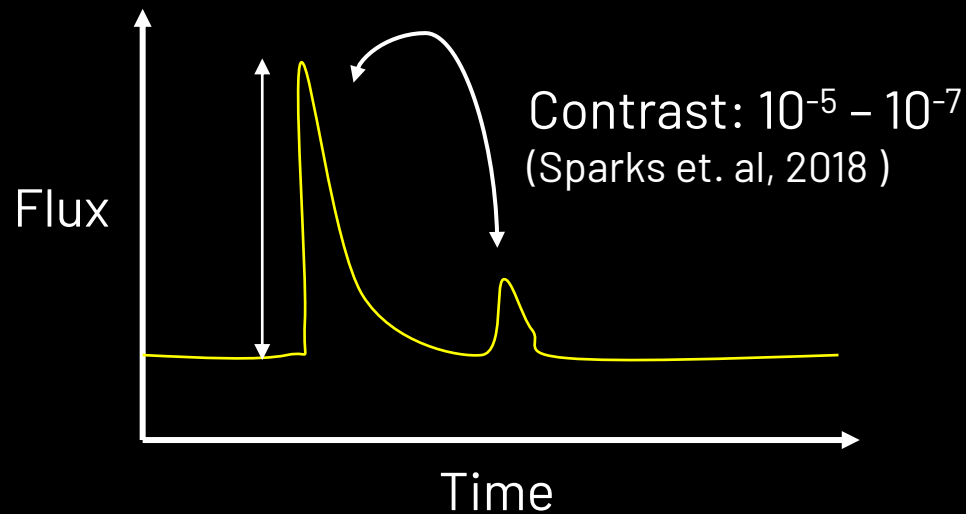




# MEASUREMENT PRINCIPLE (PHOTOMETRY)

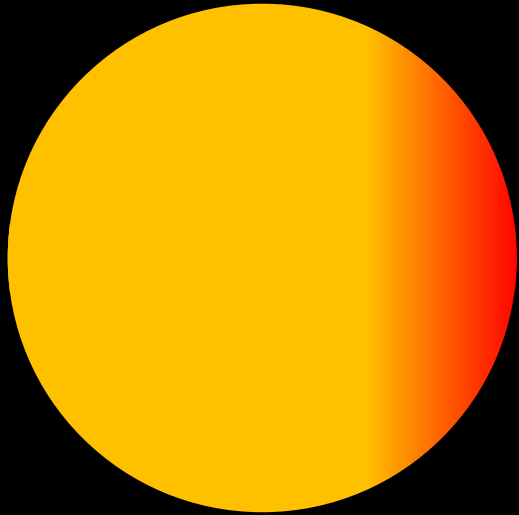
Visible in the NUV and FUV

Intensity:  $10^4 - 10^5$   
(Sparks et. al, 2018)



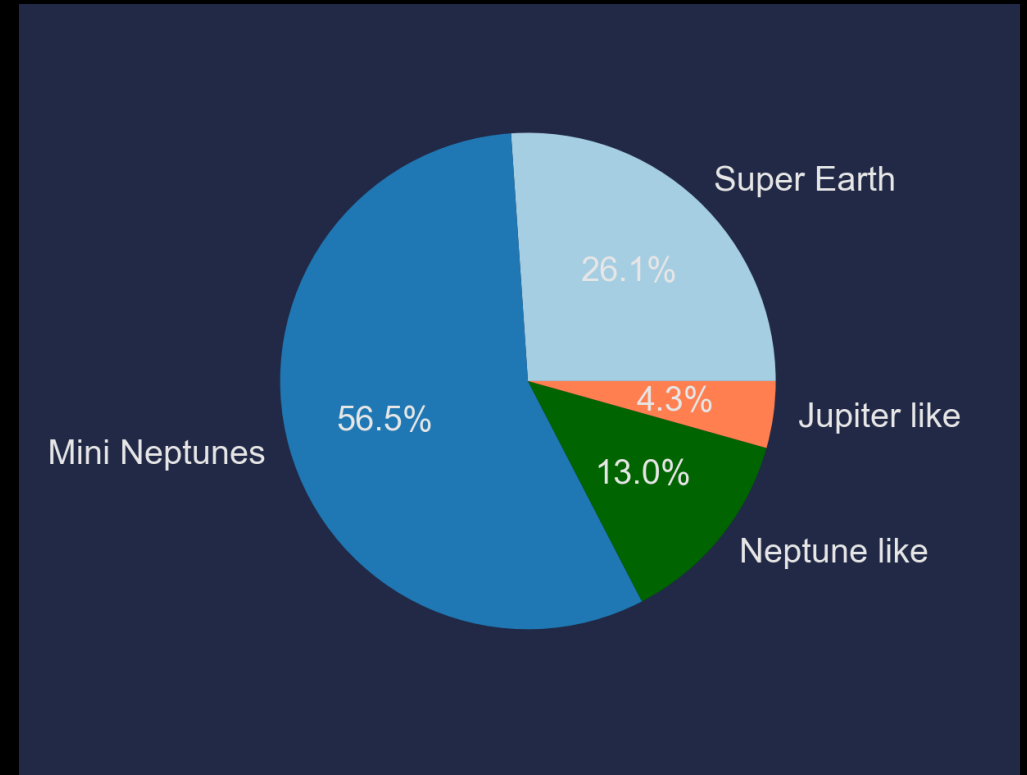
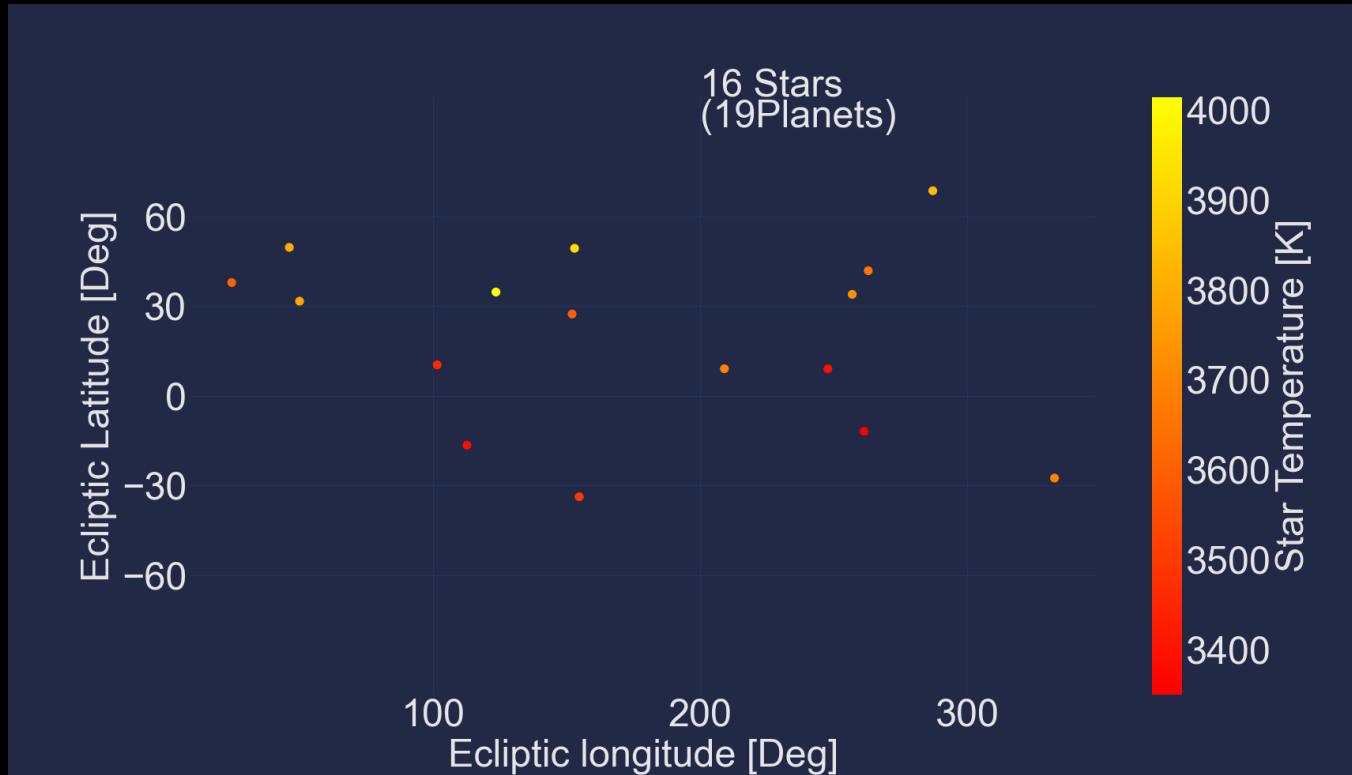
Time delay for a habitable  
zone planet around an M-  
dwarf:  
 $12.01 - 97.91 [s]$

# SENSITIVITY AND REQUIREMENTS



- M-type stars
- For a 2-sigma detection  
-> SNR of 2000
- Access planets  $>0.06$  AU  
-> maximum integration time: 30 seconds
- Bandwidth 100 - 400 nm (variable)

# TARGETS



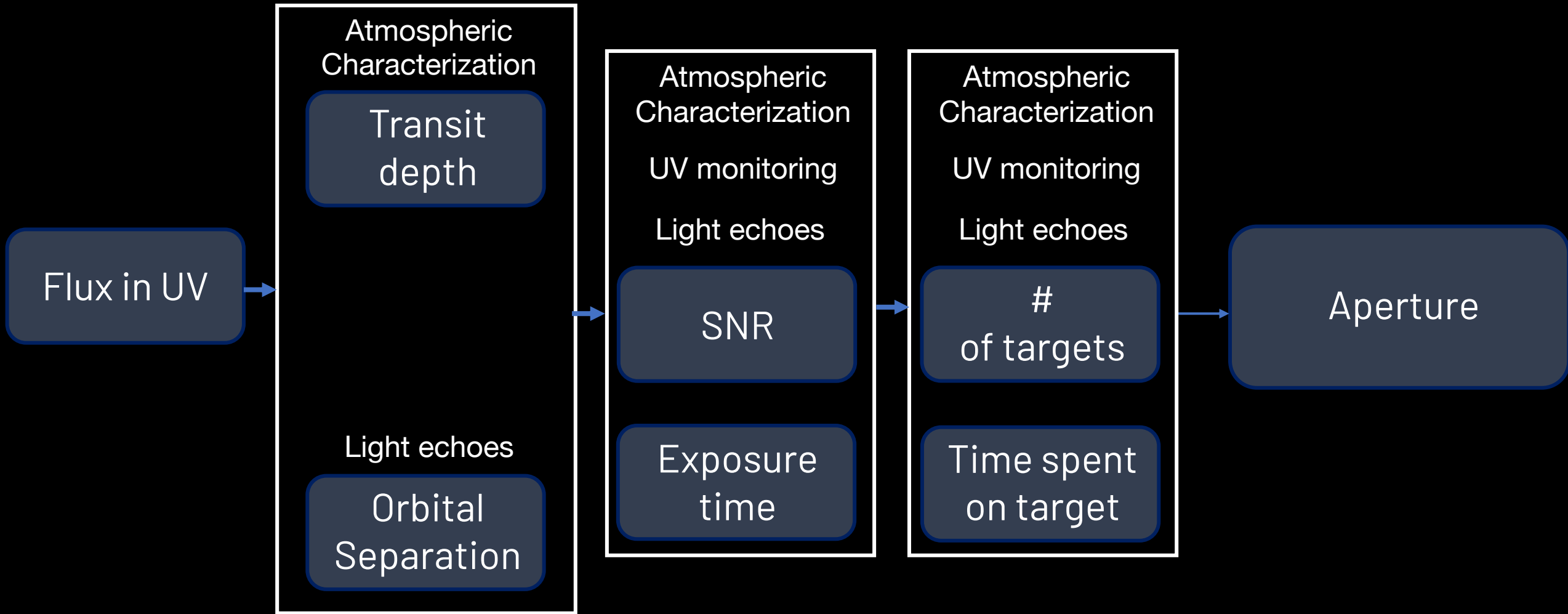


# MISSION OBJECTIVES

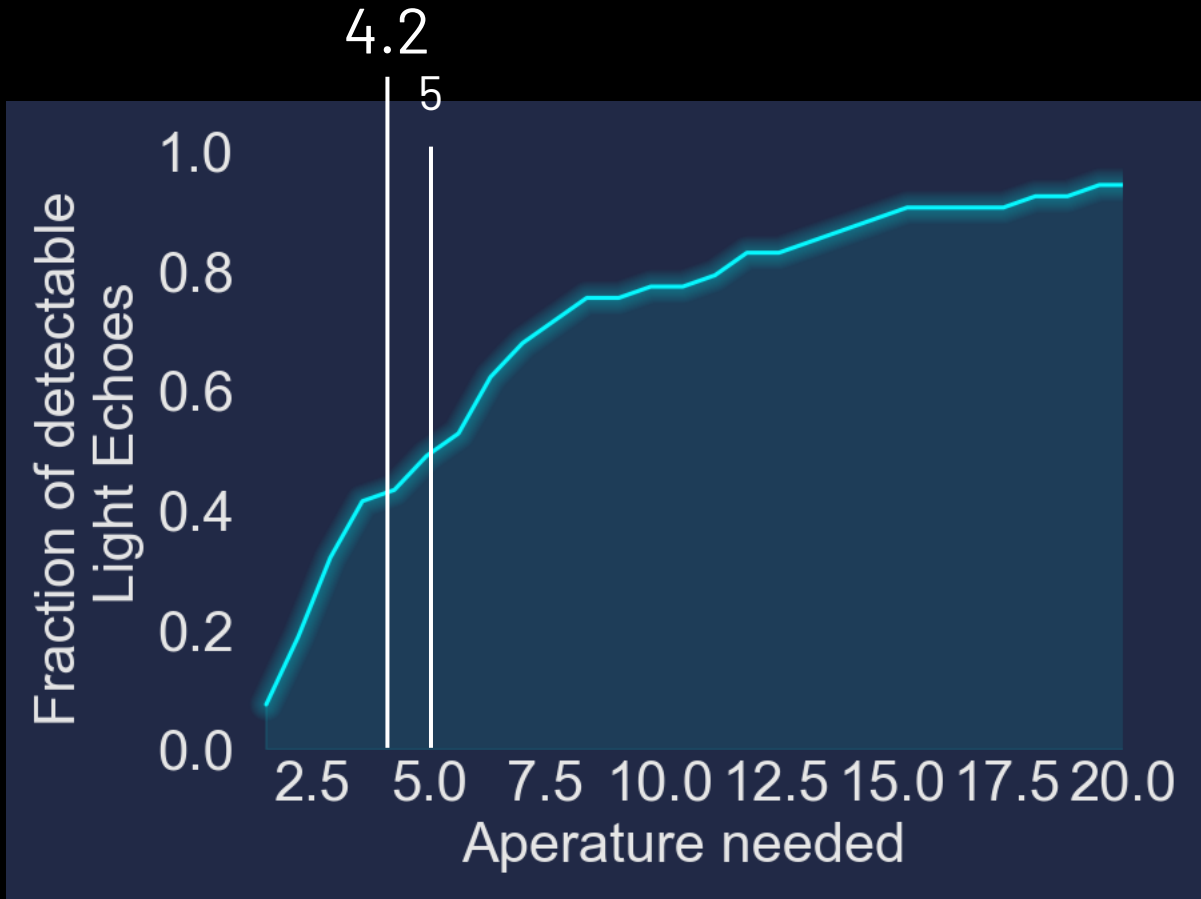


- Atmospheric Escape: Constrain evolutionary models of exoplanets
- Exospheres of habitable planets: Study formation processes of exospheres
- UV flux monitoring: Reconstrain the habitable zone for known systems
- Light echoes: Detect new habitable zone exoplanets

# REQUIREMENT ON APERTURE SIZE



# APERTURE SIZE – LIGHT ECHOES



- Initially: 50% of targets (34)
- Trade-off on fairing diameter:  
-> Down to 19 targets





# KEY MISSION DRIVERS

- Diameter of telescope: 4.2 meters.
- Pointing: 5 mas (absolute pointing + fine guiding).
- Throughput >4%.
- Orbit outside Earth's exosphere (Ly-alpha absorption).  
All within 100pc:
  - > 40 planets for exospheres of habitable planets
  - > 400 planets for atmospheric escape
  - > 150 stars for UV monitoring (16 days)
  - ~20 planets for light echo detections (16 days monitoring)

# OBSERVATIONAL STRATEGY



- **Flexibility** is key to **assist** with stellar activity monitoring
- **Observe transits** of known exoplanets, baseline of 3 transits lengths
- **Photometry** for  $0.5 \times$  rotational period of the star
- When monitoring for light echoes, if no signal is found then change target

# MISSION DURATION



Variable	Time	Total time
Average time of transit for each planet (Ext. atm.)	2.2 hrs	6 hrs per target
Average time of transit for each planet (Esc. atm.)	2.2 hrs	6 hrs per target
Number of transits required for each target (Esc. atm. and Ext. atm.)	10 x 2	120 hrs
Exposure time of each star (HZ)	Max. 30 s	15 days
Exposure time of each star (Light echoes)	Max. 30 s	15 days
Repointing	1 per target	300 hours
TOTAL	-	6.5 years



# PAYLOAD



Scientific  
goals

Payload

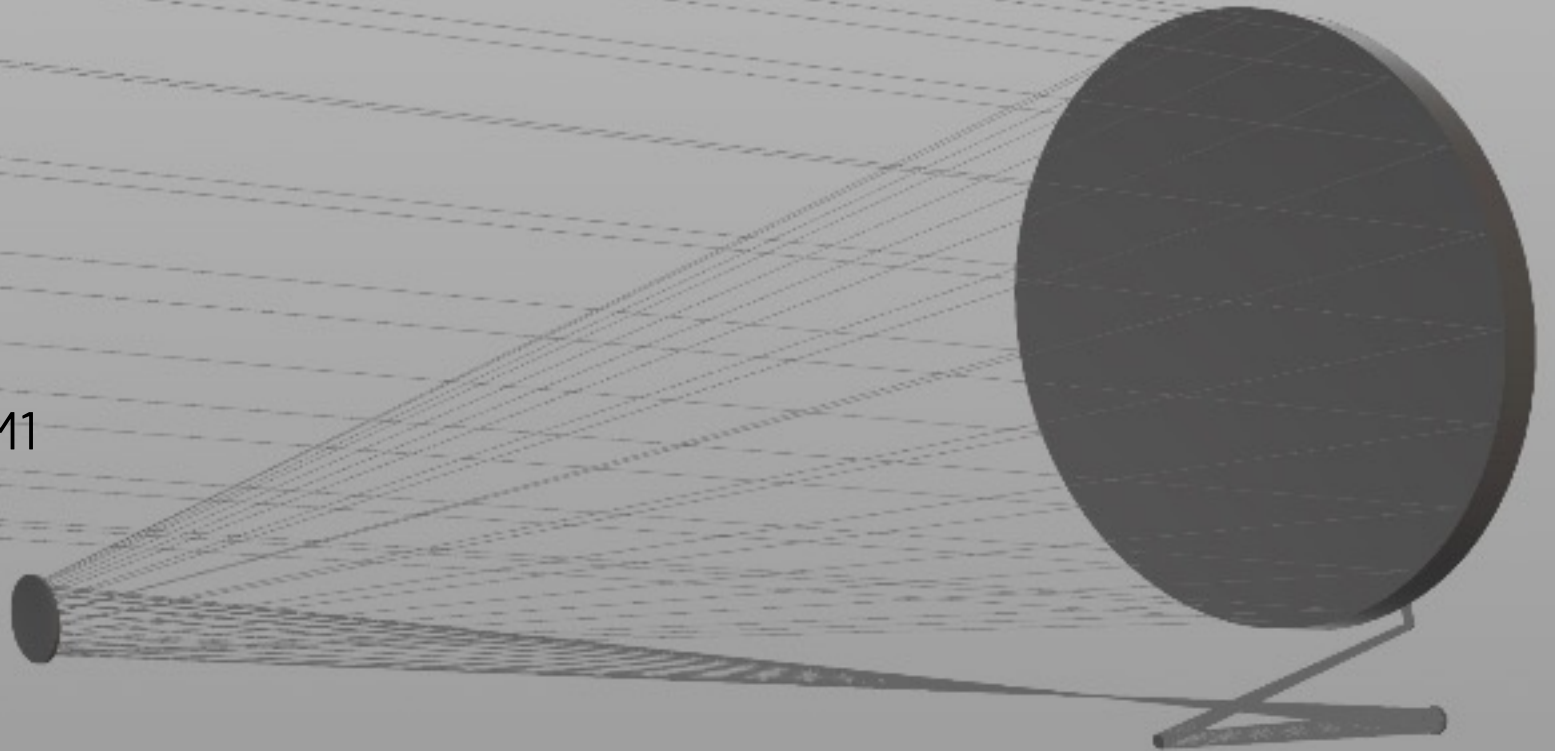
Service  
module

Conclusions

# TELESCOPE DESIGN



- **Off-axis** three-mirror anastigmat
- **4.2-meter** aperture
- Fine-steering mirror
- Additional relay mirrors to slow beam to **F/100** and fold below M1
- **High efficiency**



# INSTRUMENT: UV-IMSPEC



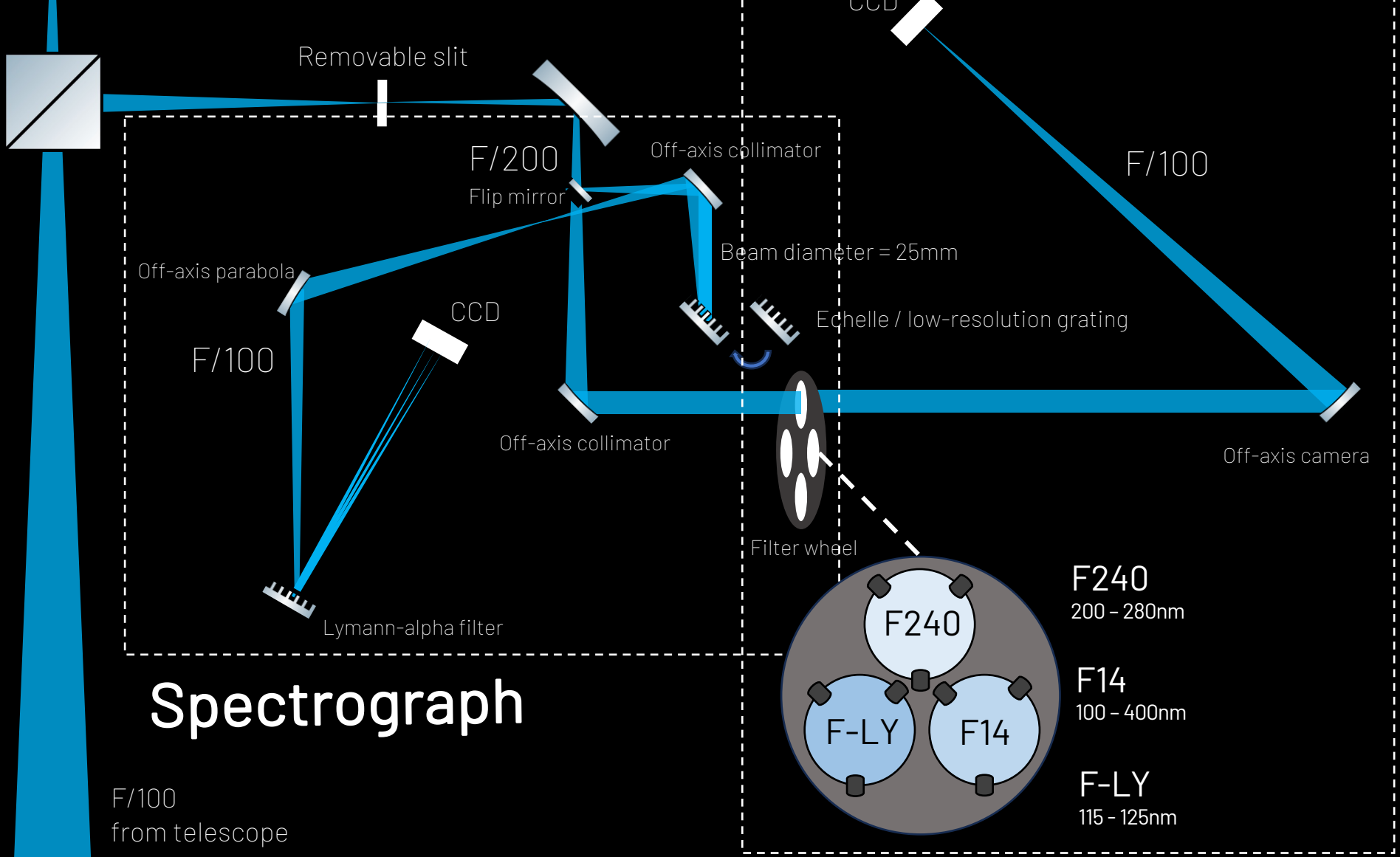
- Spectroscopy mode (R: 1000 – 20.000)
- Imaging mode for photometry





Fine guidance sensor (FGS)

# The UV-IMSPEC instrument



## Spectrograph

F/100 from telescope

## Imager

### Challenges:

- **Slow beams** to reach sampling requirements of **2 pixels/FWHM**.
- **Minimal amount of surfaces** to increase throughput.



# DETECTORS (SPECIFICATIONS)

	IMAGER	SPECTROGRAPH
Op. wavelength [nm]	100-400	115-125
MgF2 (op. for ^1)[nm]	200-250	120-150
Quant. Eff. (SOTA)	30-60%	20-50%
Target develop.	60 % (100-400 nm)	60 % (115-125 nm)
Op. temp. [K]	< 163	< 163

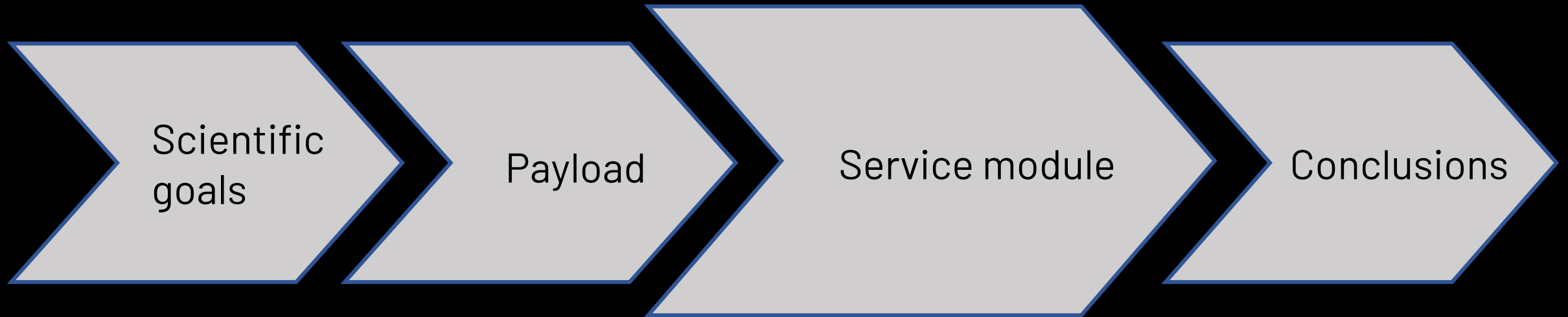
Read-out Noise<sup>2</sup>:  
< 1 e<sup>-</sup>

Dark Current Noise<sup>2</sup>:  
< 0.025/hr/pix

Pixel Size (SOTA): 10μm  
Target Development: 7μm

TRL: 5  
(Demonstrated at 39km altitude  
during FIREBall-2 mission)

# SERVICE MODULE







# KEY SYSTEM DRIVERS

DRIVER	VALUE
PRIMARY MIRROR SIZE	4.2 METERS
DISTANCE BETWEEN MAIN MIRROR AND SECONDARY MIRROR	7 METERS
POINTING ACCURACY FOR SPACECRAFT	0.1 ARCSEC
DATA AMOUNT	240 GB/DAY
MISSION DURATION	6.5 YEAR
TYPE OF ORBIT	L2
STRAYLIGHT	SHADED MIRROR

# PLATFORM REQUIREMENTS



Req-ID	Traceability	Type	Description
AOCS-01	S-21	AOCS	The ADCS shall determine and control the attitude of the satellite with an accuracy of at least 0.1arcsec.
AOCS-03	S-23	AOCS	The attitude control system shall be able to maintain a stable pointing of 0.1 arcsec for at least 16 hours.
AOCS-04		AOCS	The satellite shall be capable of executing collision avoidance maneuvers.
AOCS-05		AOCS	The ADCS shall be able to withstand singlepoint failure, without any loss of capability
COM-01	S-24	COM	The operating frequency used for telecommand and housekeeping telemetry shall be 32GHz downlink, 34.6GHz uplink (Ka-band).
COM-02	S-24	COM	The data transmission frequency used for payload and experiment data telemetry transmission shall be 31.8GHz downlink, 34.3 GHz uplink (Ka-band).
COM-03	S-24	COM	The data-rate shall be at least 130Mbps
C-01	S-23, S-10, S-11, S-12, S-13, S-14	CONST	The nominal mission duration shall be 6.5 years.
C-02		CONST	The end-of-life scenario shall comply to the ECSS standards for the space debris mitigation.
C-03		CONST	The spacecraft design shall comply to the ECSS standards.
E-01	O-02	ENV	The satellite shall withstand the accumulated radiation loads over the entire lifetime in orbit without loss of functionality.
E-03	O-02	ENV	The satellite shall withstand the thermal loads for the predicted environments during in-orbit operation.
E-04	O-02	ENV	The satellite shall withstand the mechanical loads for the predicted environments during storage, launch and on-orbit operation.
EPS-02		EPS	The EPS shall provide 100 W of electrical power to the payload.

# PLATFORM REQUIREMENTS

Req-ID	Traceability	Type	Description
EPS-03		EPS	The EPS shall have power capacity to provide the spacecraft for 8 hours without power generation.
GS-01	S-24	GS	The ground station contact time shall be sufficient to download all scientific data
GS-02	COM-01, COM-02	GS	The ground station selected shall be able to receive data in the S-band for safe mode and in Ka-band for science data transmission
GS-03	COM-01, COM-02	GS	The ground station shall be able to transmit in S-band for safe mode case and in Ka-band for telecommand
O-02	S-14, I-12	MIS	The target orbit shall be around L2 and avoid eclipses.
O-03	S-14	MIS	The orbit shall be reached and maintained for the required nominal and extended mission lifetime.
O-04		MIS	At the end of mission lifetime, the S/C shall be disposed.
OBC-01	S-24	OBC	The OBC shall be able to store 240Gbyte/day amount of data
OBC-03		OBC	The OBC shall be able to distribute commands to the various subsystems.
SMS-01		SMS	The spacecraft shall fit the payload envelope of the launcher system.
SMS-02	S-20, I-10	SMS	The secondary mirror shall be at a distance of 7m from the main mirror
SMS-03	S-21, T-03, T-05, T-06, T-09	SMS	The stability of the secondary mirror shall be sufficient to achieve a pointing accuracy of 0.005arcsec
SMS-04	S-25	SMS	The sunshield shall shade the telescope from all incoming straylight
SMS-05	S-21, T-03, T-05, T-06, T-09	SMS	The thermal stability of the primary mirror backplane shall be sufficient to achieve a pointing accuracy of 0.005arcsec
TCS-01		TCS	The TCS shall monitor and keep the temperature of each component within the required operational limits



# TOP REQUIREMENTS

Req-ID	Traceability	Type	Description
AOCS-01	S-21	AOCS	The ADCS shall determine and control the attitude of the satellite with an accuracy of at least 0.1arcsec.
AOCS-03	S-23	AOCS	The attitude control system shall be able to maintain a stable pointing of 0.1 arcsec for at least 16 hours.
C-01	S-23, S-10, S-11, S-12, S-13, S-14	CONST	The nominal mission duration shall be 6.5 years.
O-02	S-14, I-12	MIS	The target orbit shall be around L2 and avoid eclipses.
GS-01	S-24	GS	The ground station contact time shall be sufficient to download all scientific data
SMS-01		SMS	The spacecraft shall fit the payload envelope of the launcher system.
SMS-02	S-20, I-10	SMS	The secondary mirror shall be at a distance of 7m from the main mirror

# MISSION ANALYSIS



## TARGET

- Halo orbit around Sun-Earth L2 point
- Apoapsis altitude:  $8.52 \times 10^5$  km
- Periapsis altitude:  $2.5 \times 10^5$  km
- Small orbit corrections necessary

Getting there  $\longrightarrow$

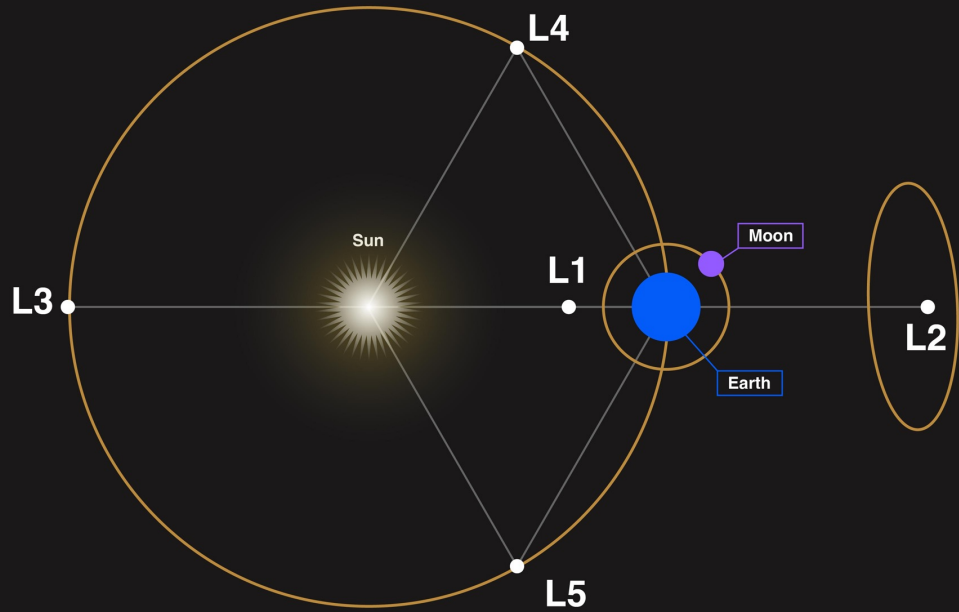
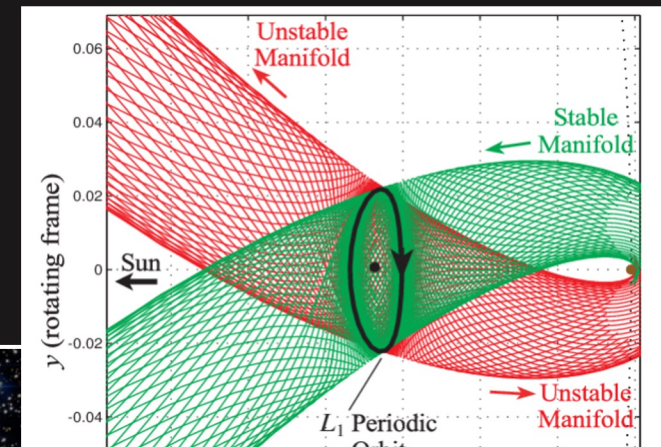


Image: NASA

## Launch Window considerations

- Earth/Moon eclipses
- Instruments orbiting L2
- Stable Manifold Transfer
- Target Orbit Insertion

Wang et al., 2000

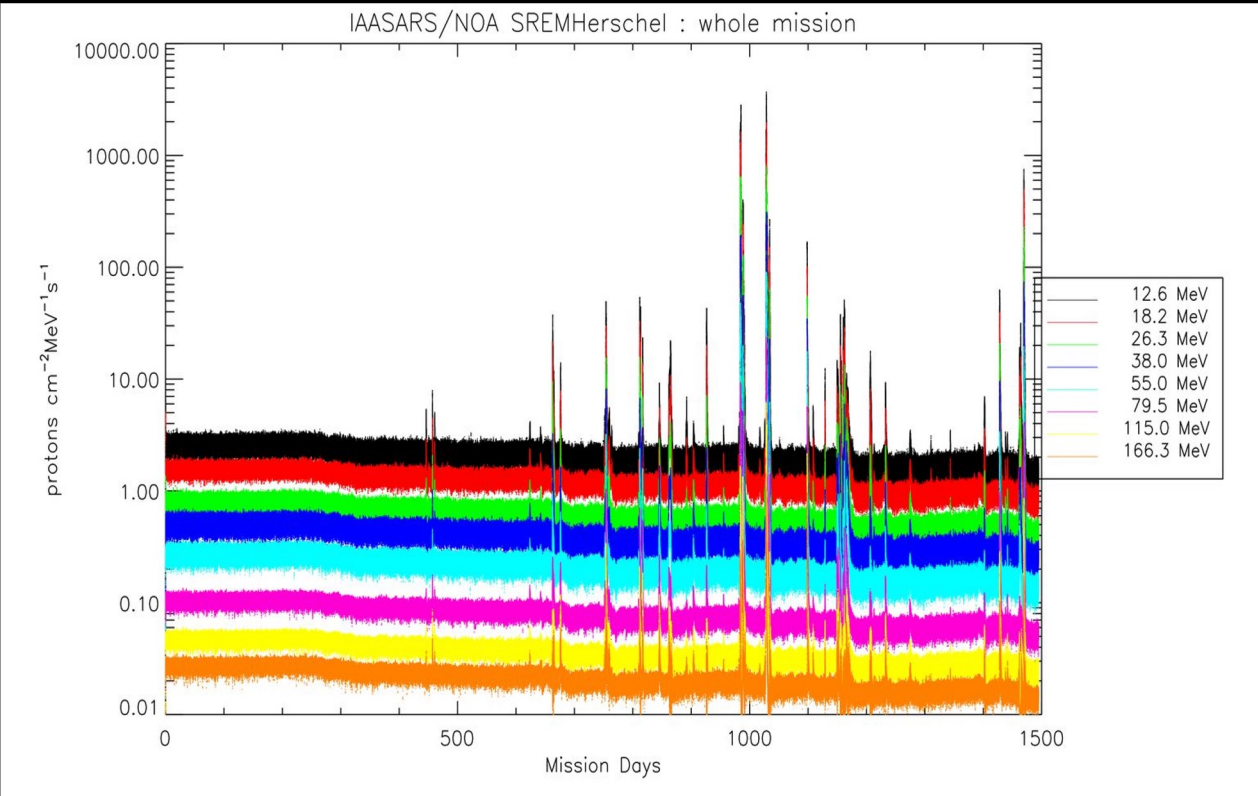




# RADIATION ENVIRONMENT



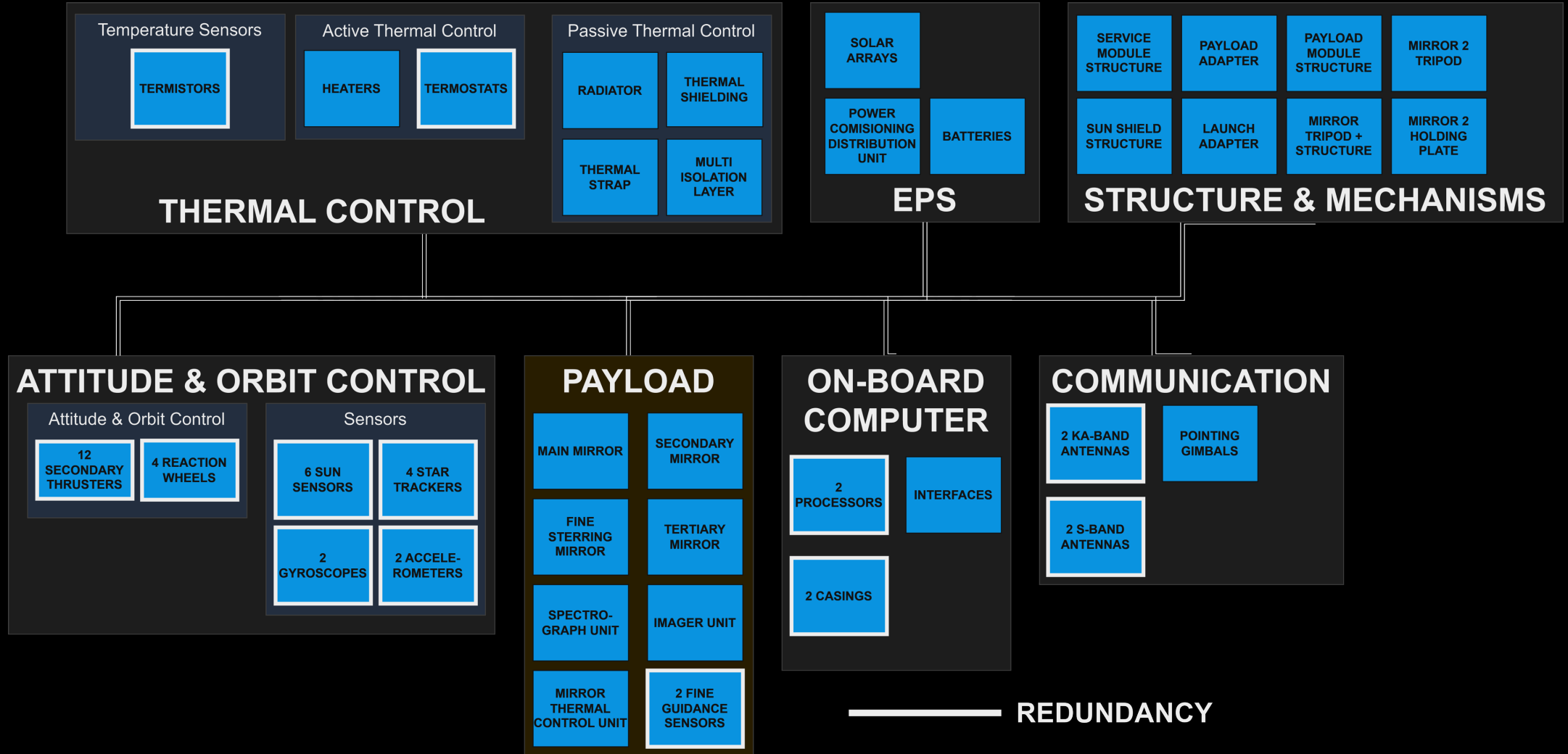
- The radiation environment at L2 is less harmful than LEO or GEO
- Sources are the Sun and cosmic ray background
- 1- 10 particles/cm<sup>3</sup> (95% protons)
- Herschel mission ~1 SEU/day (“bit flip”)
- Need for Failure Detection Isolation and Recovery (FDIR) in the OBC.



Ingmar Sandberg (ESTEC)



# SYSTEM OVERVIEW





# STRUCTURE & MECHANISMS REQUIREMENTS

- **SMS-01:** The spacecraft shall fit the payload envelope of the launcher system
- **SMS-05:** The thermal stability of the primary mirror backplane shall be sufficient to achieve a pointing accuracy of  $0.005\text{arcsec}$

# Structures and Mechanisms System

**SERVICE  
MODULE  
STRUCTURE**

**PAYLOAD  
ADAPTER**

**PAYLOAD  
MODULE  
STRUCTURE**

**MIRROR 2  
TRIPOD**

**SUN SHIELD  
STRUCTURE**

**LAUNCH  
ADAPTER**

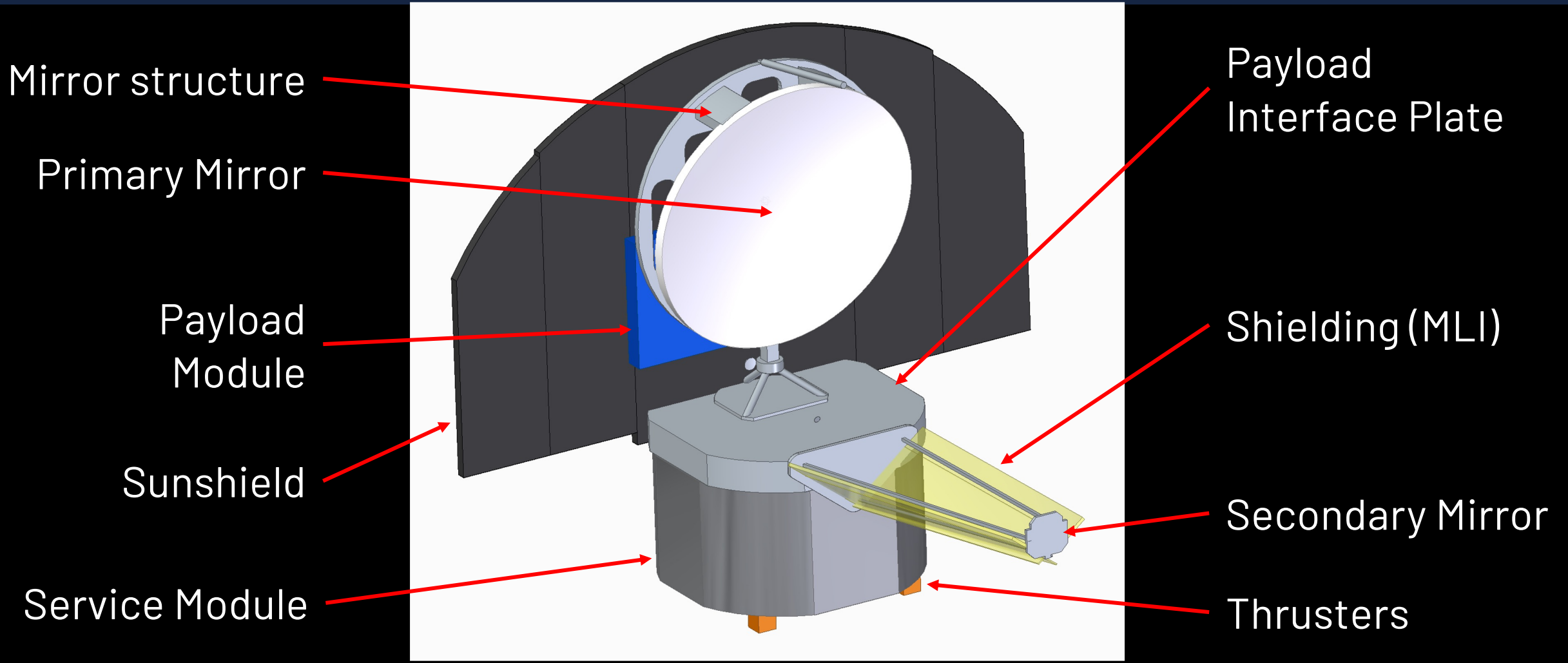
**MIRROR  
TRIPOD +  
STRUCTURE**

**MIRROR 2  
HOLDING  
PLATE**

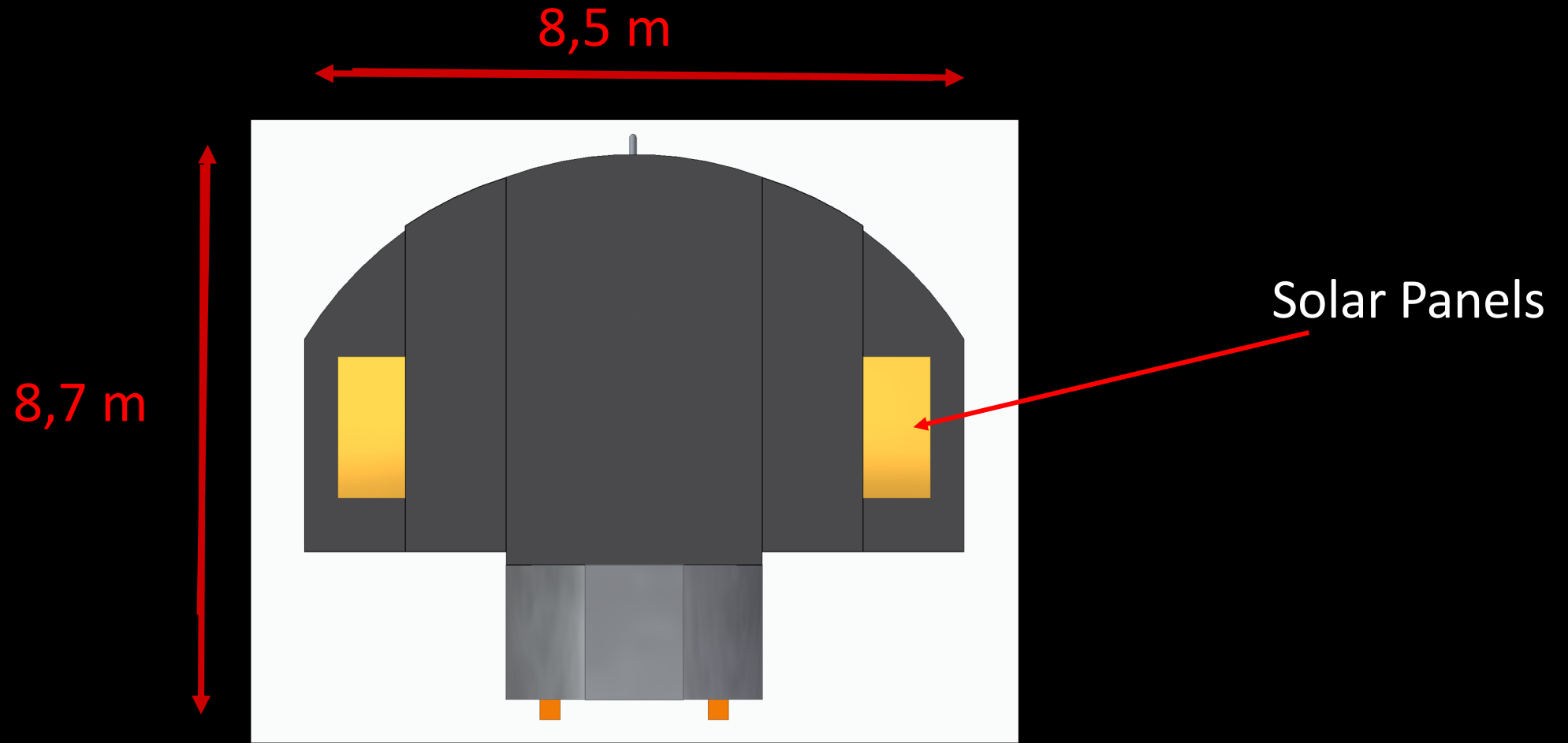
## **STRUCTURE & MECHANISMS**



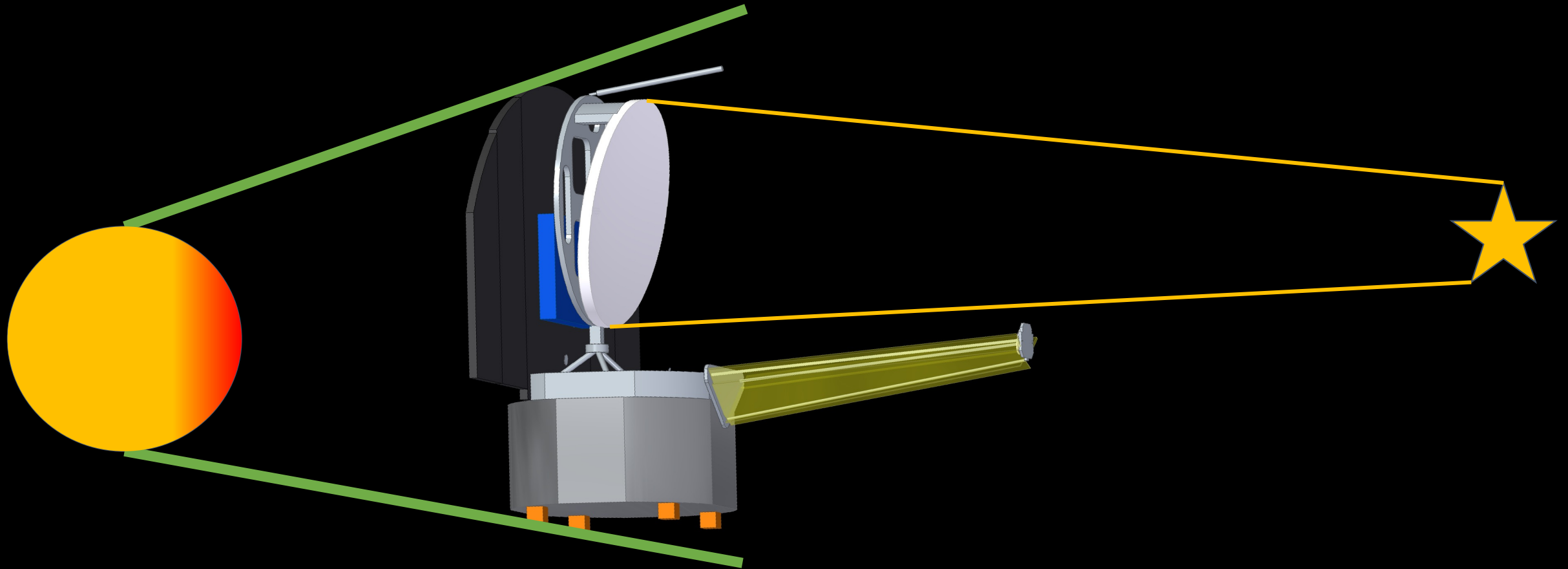
# SATELLITE DESIGN



# SATTELITE DESIGN

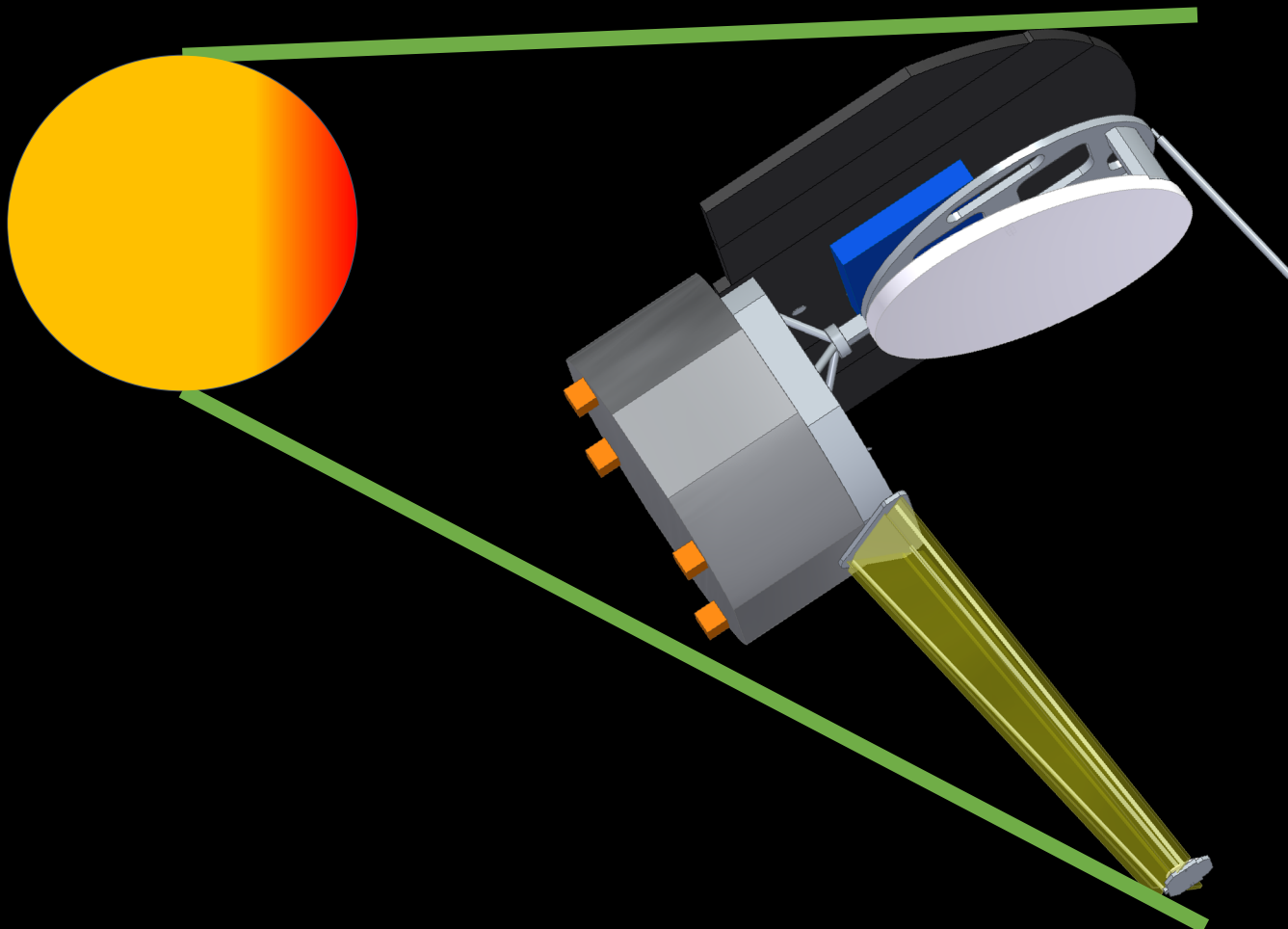


# SUBSYSTEMS - STRUCTURES AND MECHANISMS



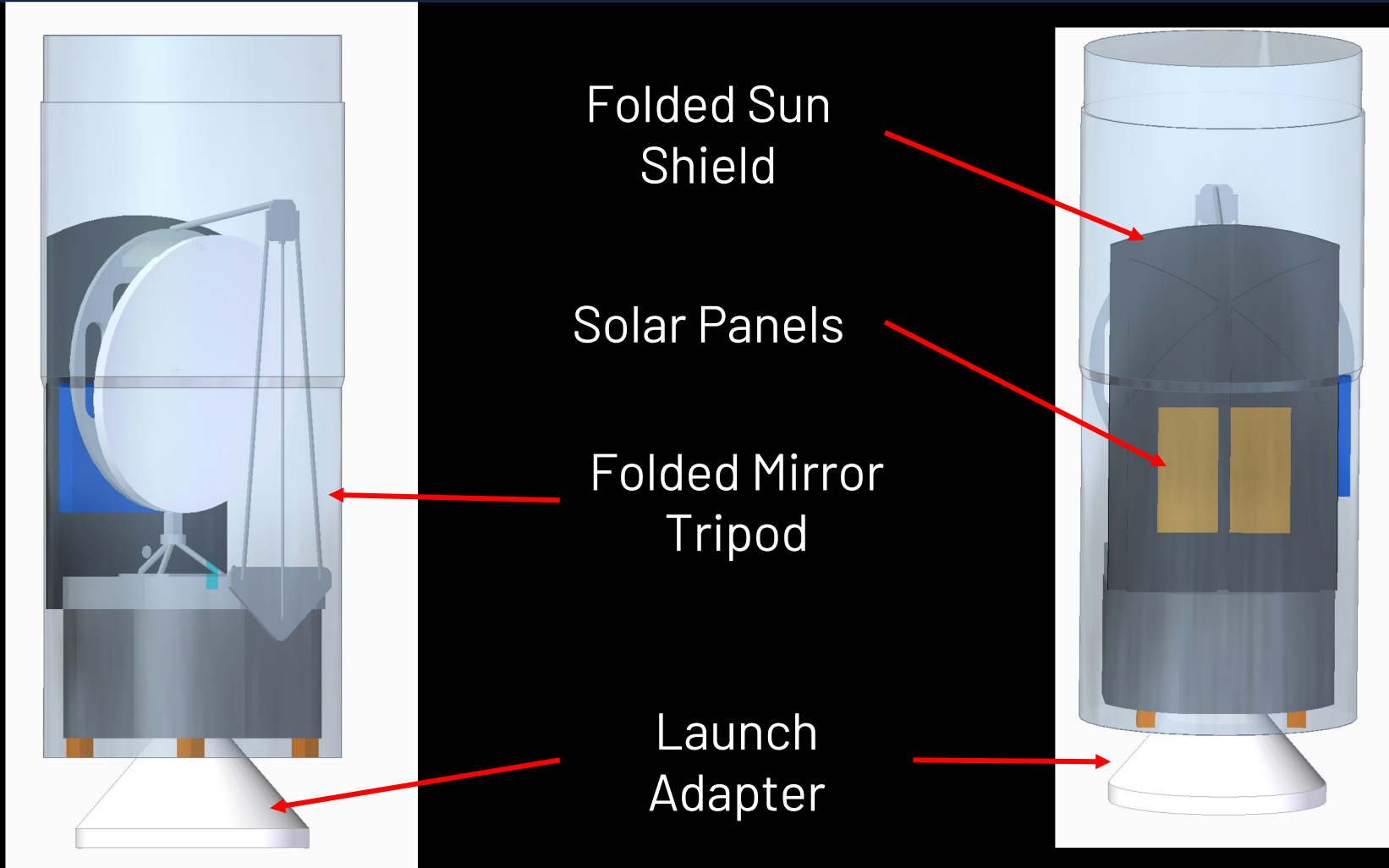


# SUBSYSTEMS - STRUCTURES AND MECHANISMS



+/- 70°

# SPACECRAFT IN LAUNCH CONFIGURATION

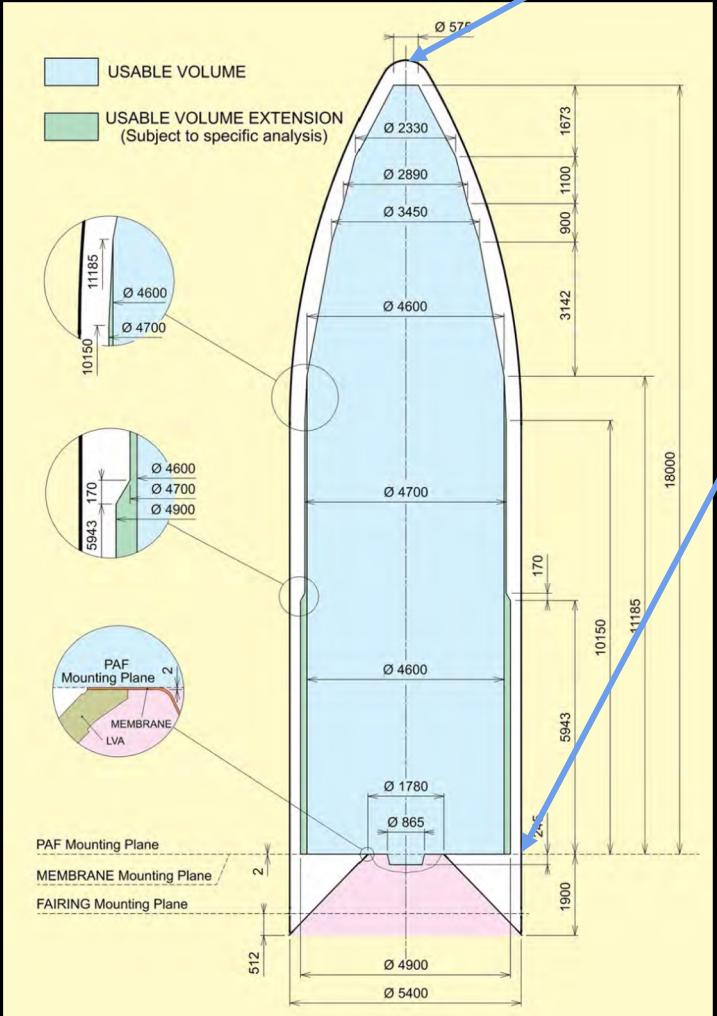


# LAUNCHER – ARIANE 6.2



Fairing Diameter (internal)	4.6m
Fairing Height (internal)	18m
Max. Payload Mass	3300kg

Payload Height	11m (incl. 20% margin)
Payload Mass	3000kg (incl. 20% margin)





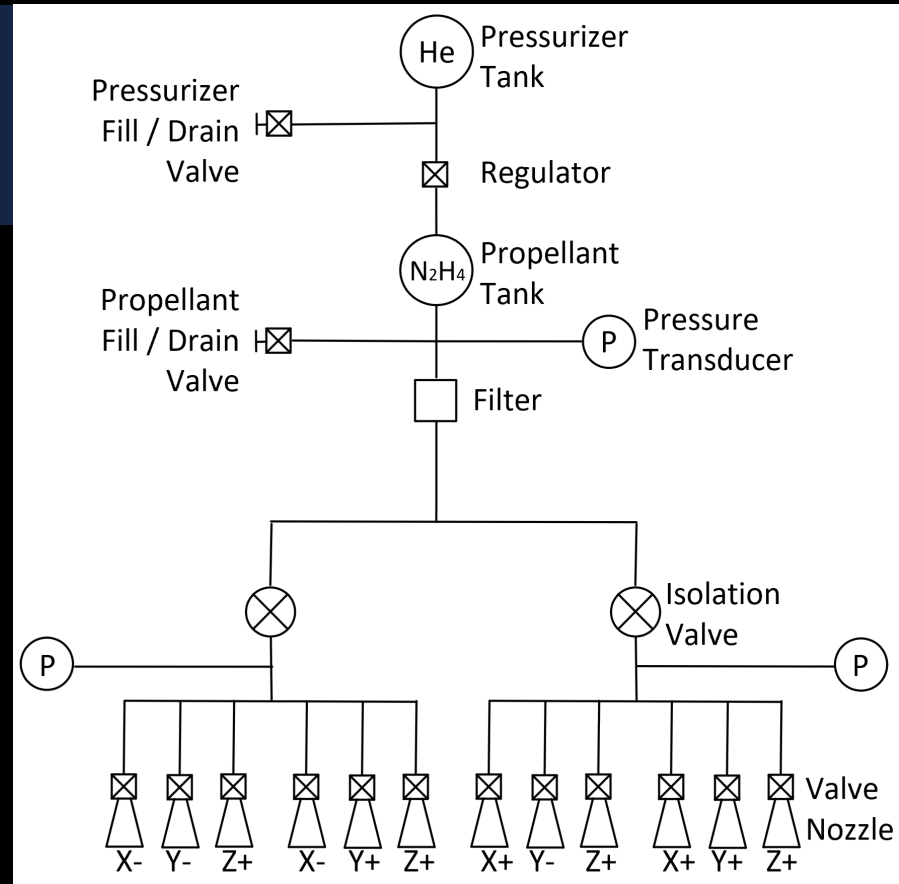
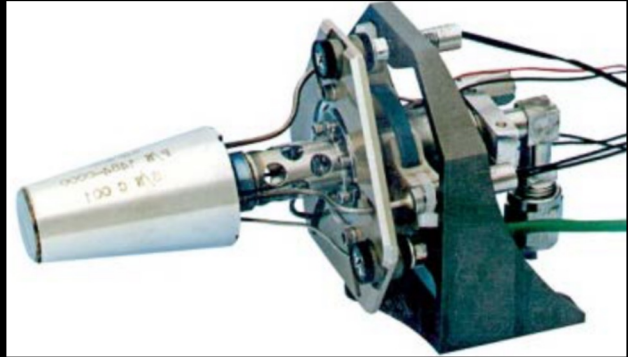
# DELTA-V



Category	Description	Delta-V	Margin	Total Delta-V
Orbital Maneuver	Removal of launcher dispersion	50 m/s	5 %	90 m/s
	Mid Coarse correction	5 m/s	5 %	
	Orbit maintenance	1 m/s / year	5 %	
	Deorbit maneuver	20 m/s	5 %	
Attitude Change	Wheel desaturation	1 m/s / year	100 %	25 m/s
	Safe mode reserve	5 m/s	0 %	
<b>Total</b>				<b>115 m/s</b>

# PROPULSION

5N Monopropellant Thruster - Rafael Space	
Propellant	N <sub>2</sub> H <sub>4</sub>
Steady Thrust	6.1 - 1.8 N
Total Lifetime	15 years
TRL	9



Maneuver Duration	
Removal of launcher dispersion (50 m/s)	~ 3 min
Wheel desaturation	< 10 s



# COMMUNICATIONS REQUIREMENTS

- COM-01: The operating frequency used for telecommand and house-keeping telemetry shall be 32GHz downlink, 34.6GHz uplink (Ka-band).

- 

COM-02: The data transmission frequency used for payload and experiment data telemetry transmission shall be 32.8GHz downlink, 34.3 GHz uplink (Ka-band).



# Communications System

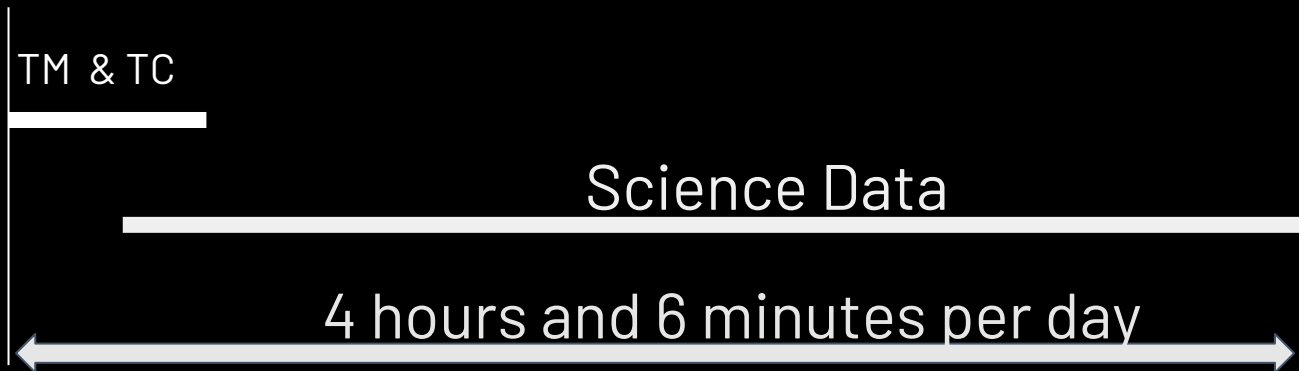
## COMMUNICATION

**2 KA-BAND  
ANTENNAS**

**POINTING  
GIMBALS**

**2 S-BAND  
ANTENNAS**

# COMMUNICATIONS



Acquisition of Signal (AOS)



**Data Rate:** 130Mbps (Ka-band)  
**Data Amount:** 240 Gbytes/day  
**Transmission Power:** 10 Watts  
**Stations:** Cebreros & Malargüe (35-meter antenna)

Loss of Signal (LOS)



- 2 Ka-band high-gain antennas (equipped with pointing gimbals, and with a 0.26-meter diameter) and 2 S-band medium-gain antennas in the telescope.
- **Telemetry and Telecommands:** frequency: Ka-band, 32.0 GHz downlink, 34.6 GHz uplink
- **Science Data:** frequency: Ka-band, 31.8 GHz downlink, 34.3 GHz uplink
- **Emergency Situations:** New Norcia antenna, using S-band, with frequencies of 2200MHz downlink, and 2050MHz uplink. Antenna size of 0.2m.

# Attitude and Orbit Control System

## ATTITUDE & ORBIT CONTROL

### Attitude & Orbit Control

12  
SECONDARY  
THRUSTERS

4 REACTION  
WHEELS

### Sensors

6 SUN  
SENSORS

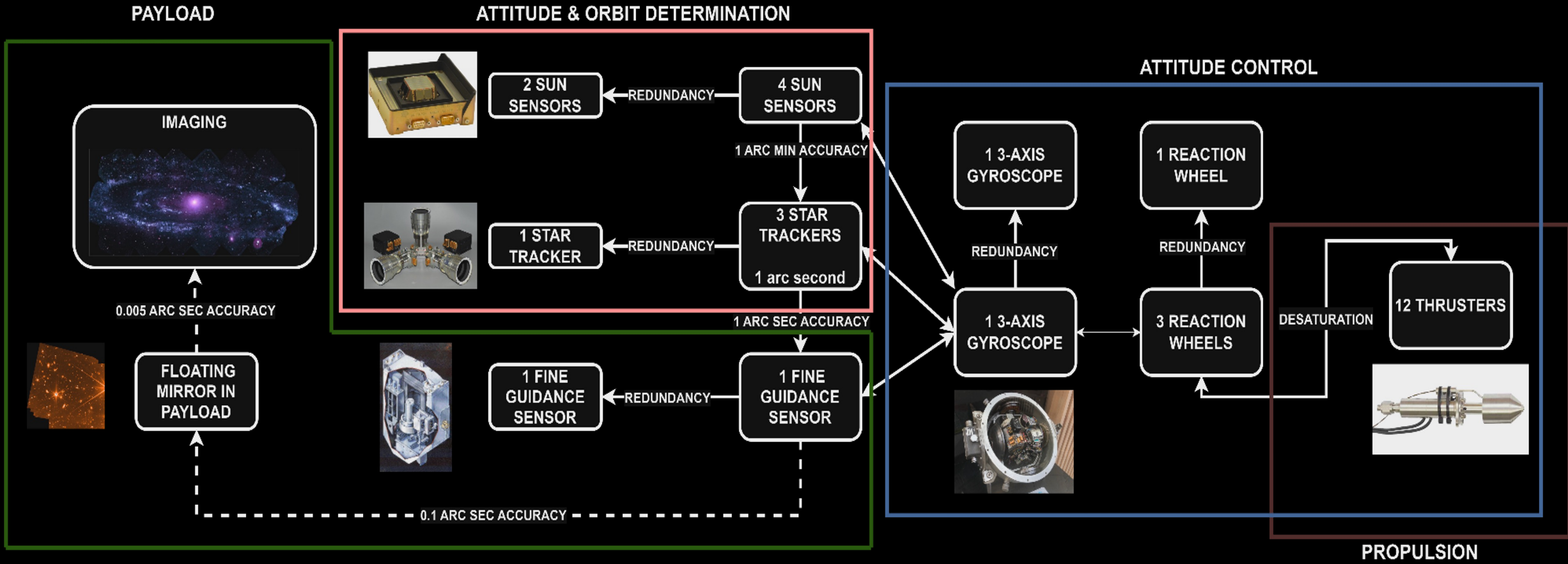
4 STAR  
TRACKERS

2  
GYROSCOPES

2 ACCELE-  
ROMETERS

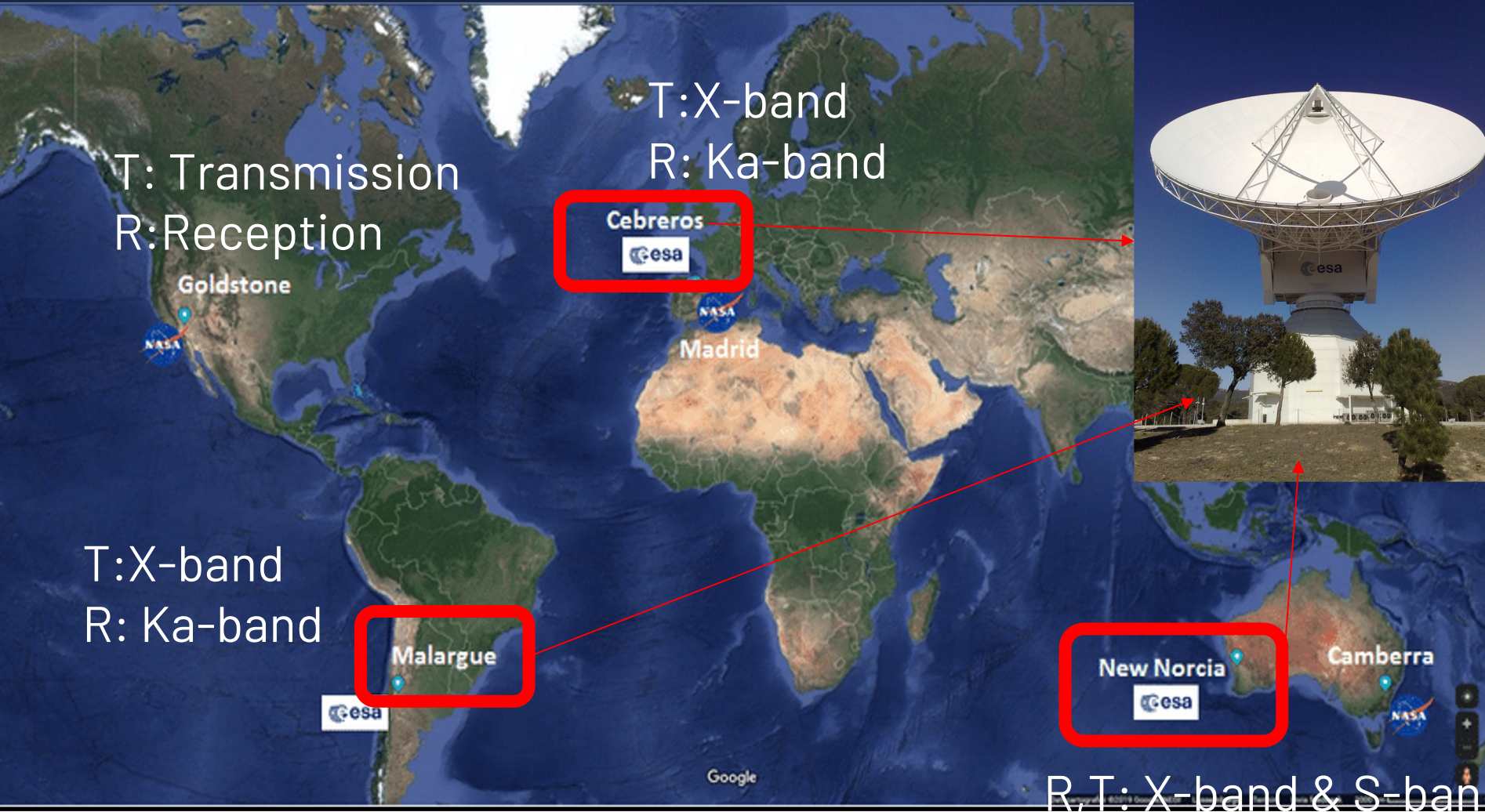


# ATTITUDE & ORBIT CONTROL SYSTEM





# GROUND STATIONS: ESTRACK Deep Space Network



- All 3 antennas with a 35-meter diameter
- Deep-space probe tracking

# ON BOARD COMPUTER REQUIREMENTS



- OBC-01: OBC-01: The OBC shall be able to store 240Gbyte/day amount of data.



# On-board Computer

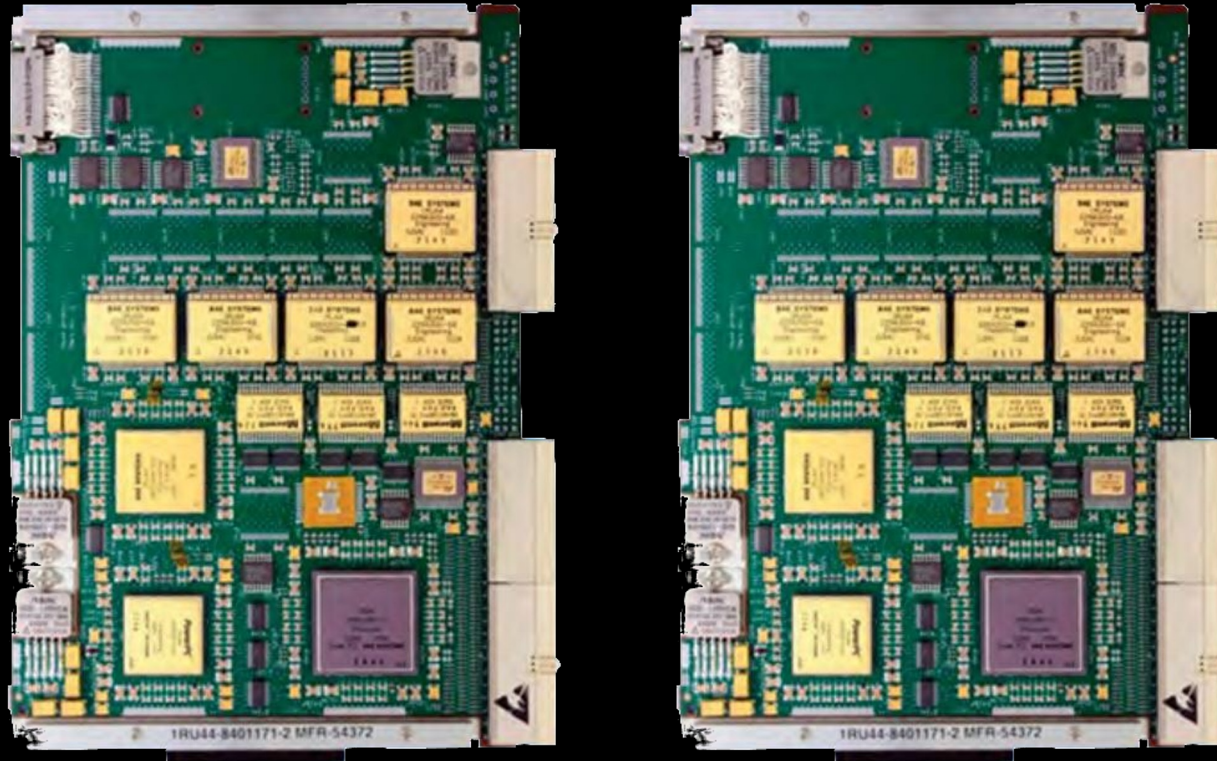
## ON-BOARD COMPUTER

2  
PROCESSORS

INTERFACES

2 CASINGS

# ON BOARD COMPUTER (OBC)



MASS MEMORY: 1 TB

↑  
REDUNDANCY

NAME	<b>RAD750 SBC</b>
SUPPLIER	<b>BAE SYSTEMS</b>
POWER CONSUMPTION	<b>10W</b>
PRODUCED	<b>2001</b>
FREQUENCY	<b>200 MHz</b>
TRL	<b>9</b>
FLIGHT HERITAGE	<b>JWST</b>

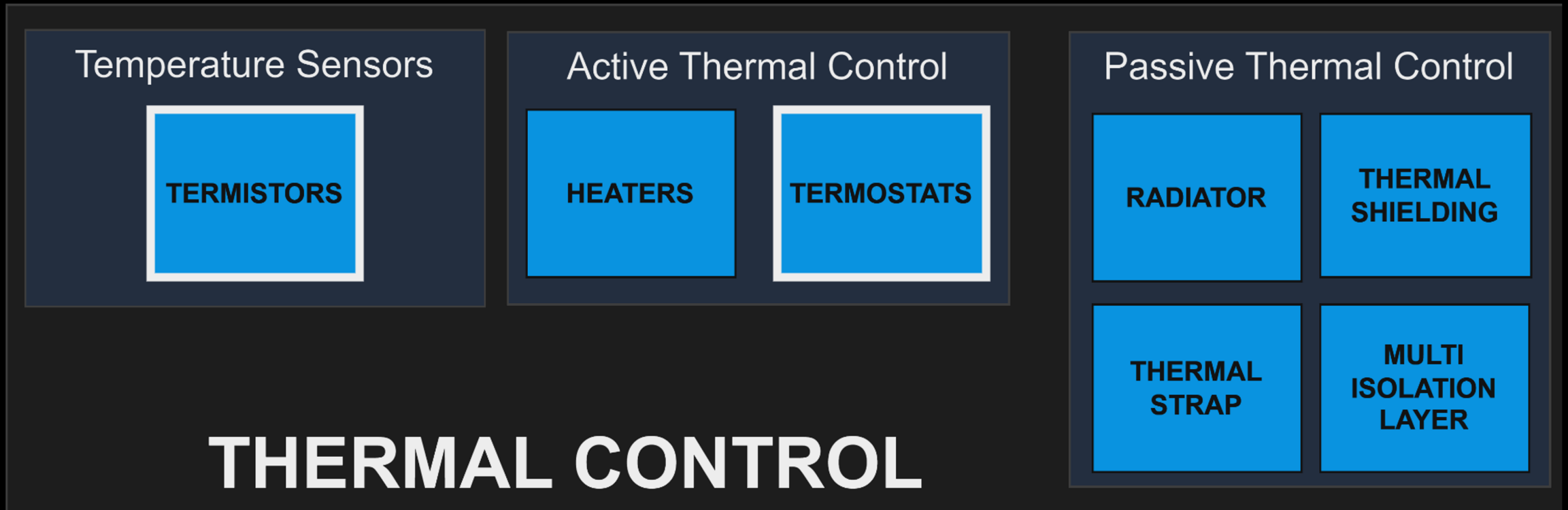


# THERMAL CONTROL SYSTEM REQUIREMENTS

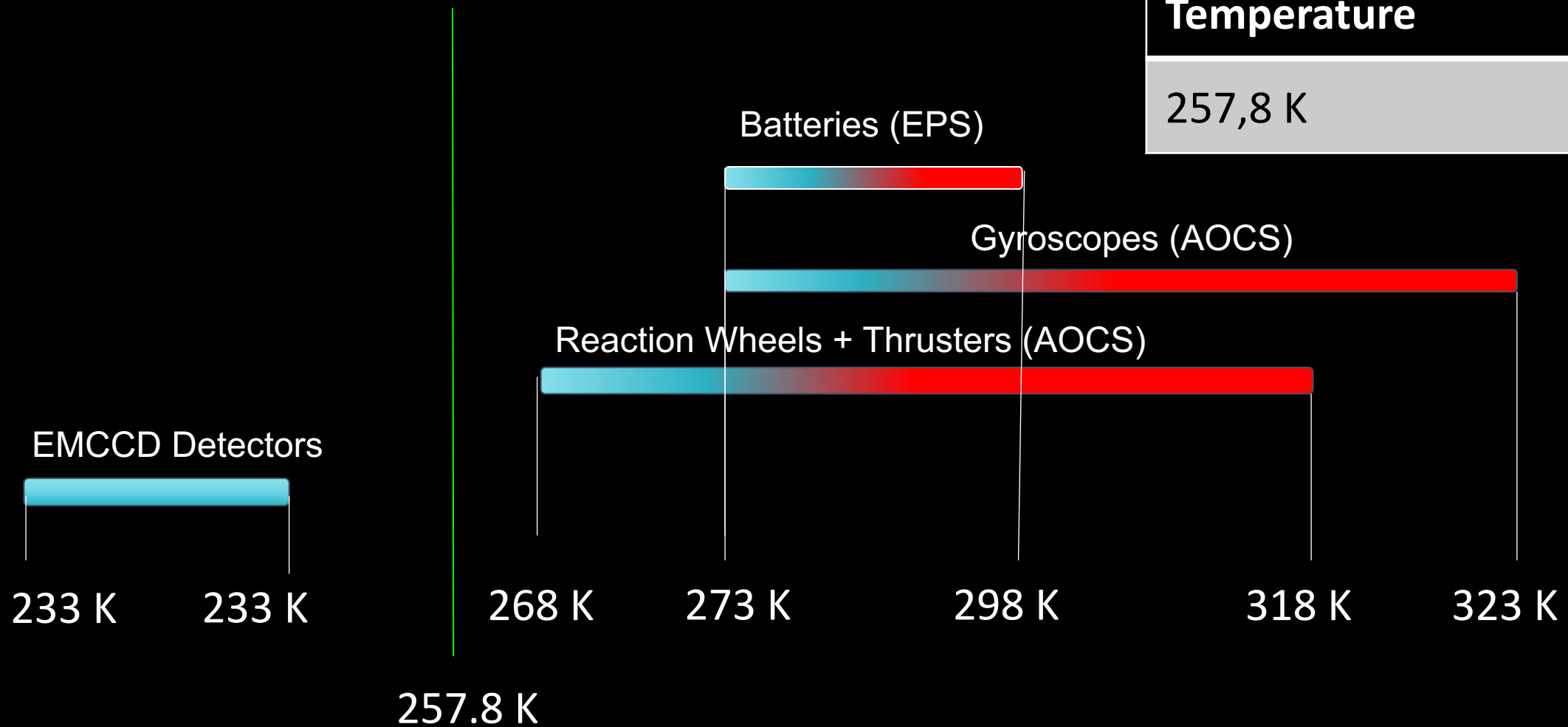
- **TCS-01:** The TCS shall monitor and keep the temperature of each component within the required operational limits.



# Thermal Control System



# THERMAL CONTROL SYSTEM





# THERMAL CONTROL SYSTEM

## PASSIVE HEATING

Thermal/sun Shield

MLI layer

Radiator on the shadow side of S/C

Thermal strap linked to radiators

## ACTIVE HEATING

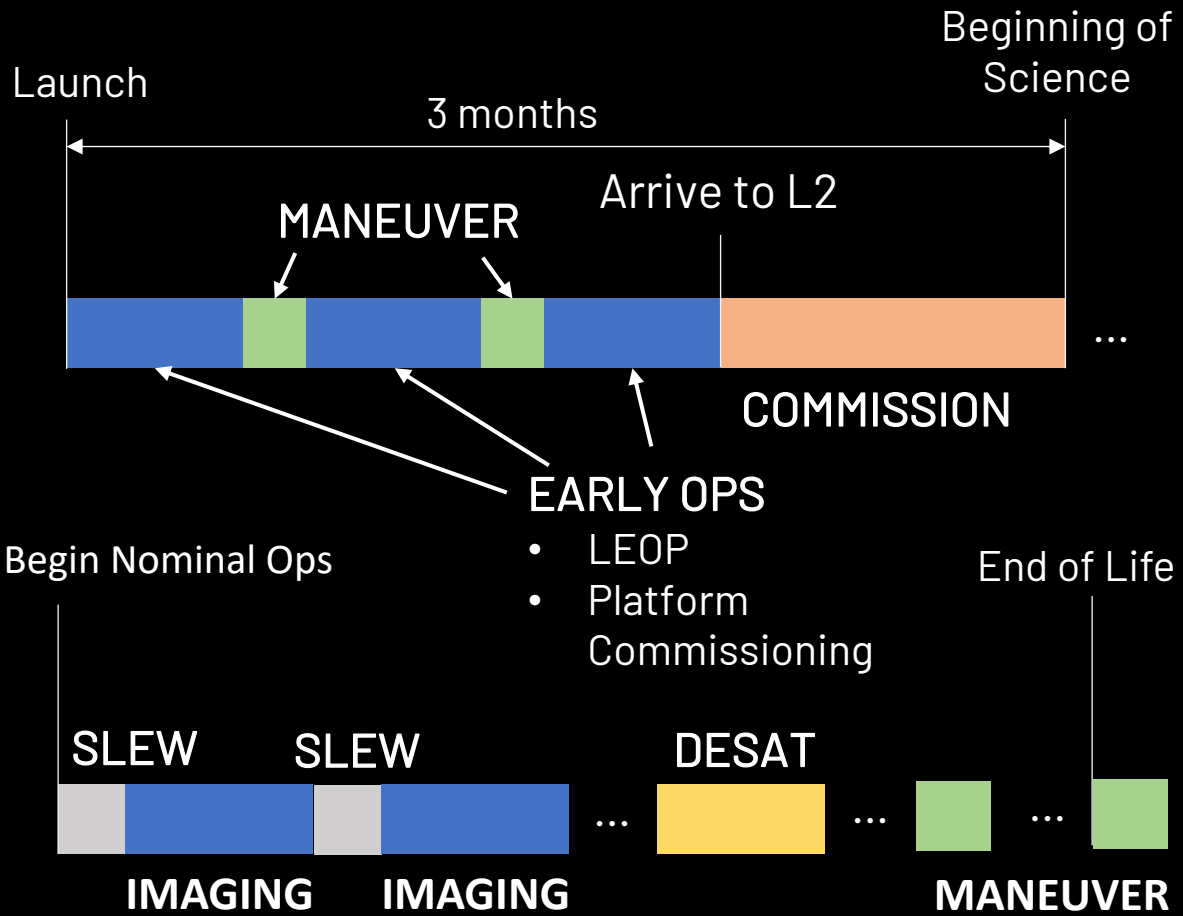
Heaters with thermostat

Max Power Dissipation	500 W
Radiator Area	1,23 m <sup>2</sup>

	Thermal Stability ( $\Delta T$ )
EMCCD Detectors	5 K



# SYSTEM MODES



System/Subsystem	USED							
	NOT USED	EARLY OPS	COMMISSIONING	MANEUVER	SLEW	DESATURATION	IMAGING	SAFE
DHS (Data-Handling)								
OBC								
PWR (Power)								
PCDU + Battery Packs								
AOGNC (Attitude and Orbit Guidance, Navigation and Control)								
Reaction Wheel								
Sun Sensors								
Star Trackers								
Gyros								
Accelerometers								
COM (Communications)								
Ka-band Transceiver								
S-band Transceiver								
INS (Instruments)								
M2 Correction Mechanism								
M4 Fine Steering Mirror								
Fine Guidance Sensors								
UV-IMSPEC								
PRO (Propulsion)								
FCV (Valve)								
Thruster Heaters								
SMS (Structure & Mechanisms)								
Heat Shield Mechanism								
M2 Folding Mechanism								



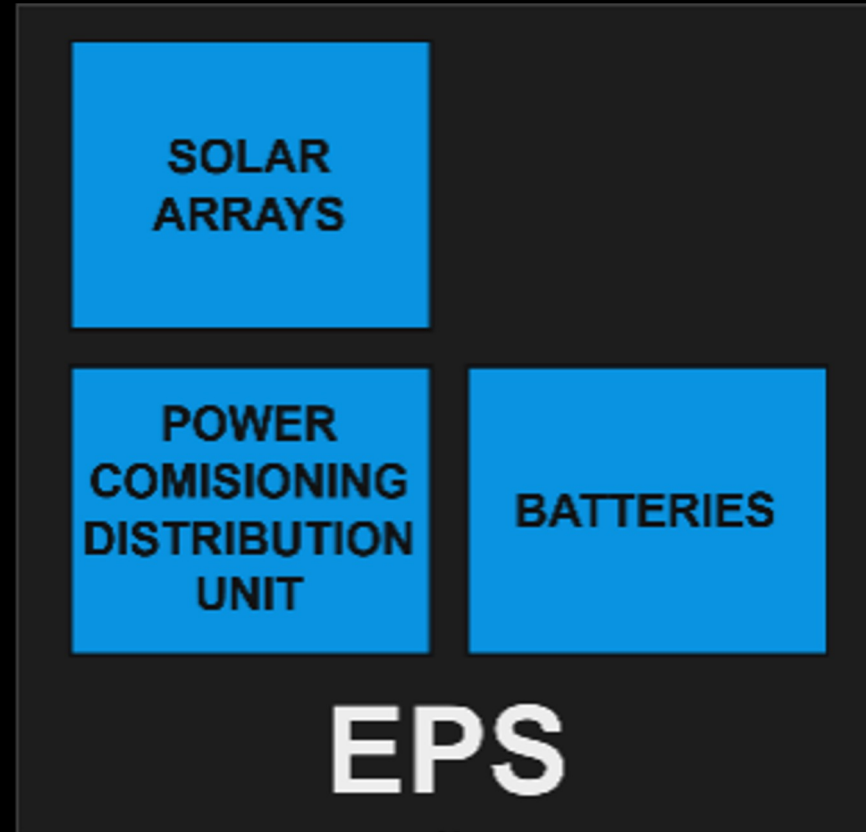
# ELECTRICAL POWER SYSTEM REQUIREMENTS

- **EPS-02:** The EPS shall provide 100 W of electrical power to the payload.

- 

**EPS-03:** The EPS shall have power capacity to provide the spacecraft for 8 hours without power generation

# Electrical Power System

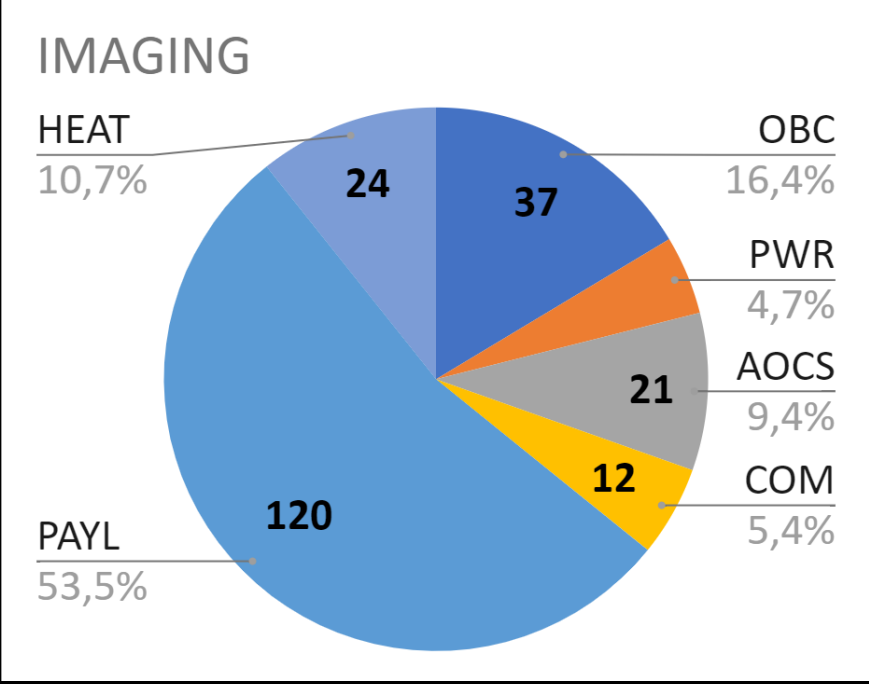
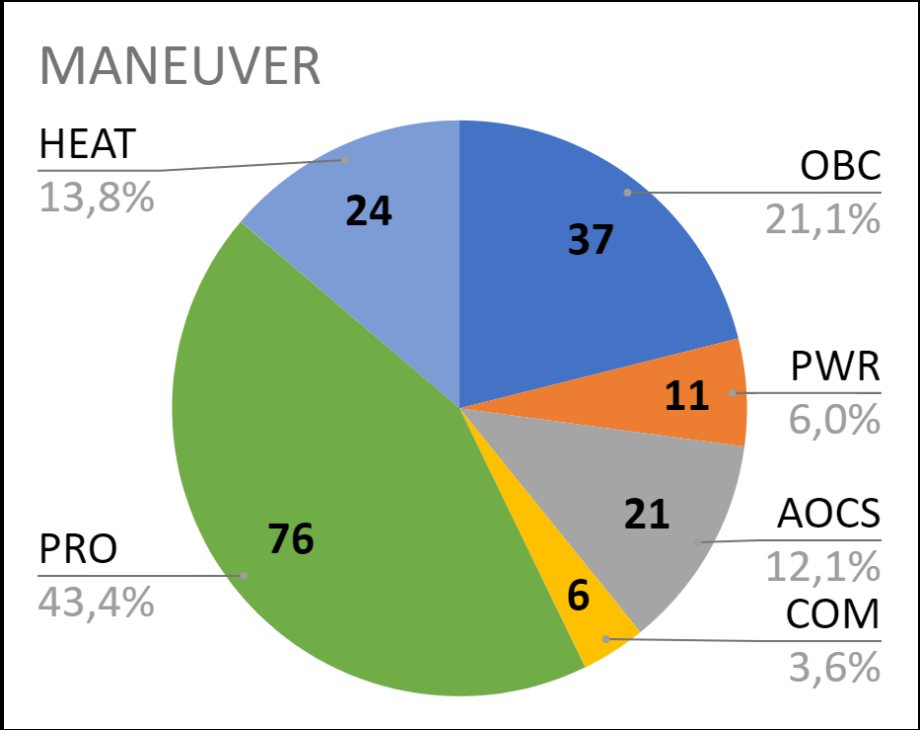




# POWER BUDGET



MODE	POWER REQ. [W]
EARLY OPS	203
COMMISSIONING	133
MANEUVRE	209
<b>IMAGING</b>	<b>269</b>
SLEW	131
<b>DESATURATION</b>	<b>349</b>
IDLE	118
SAFE	344

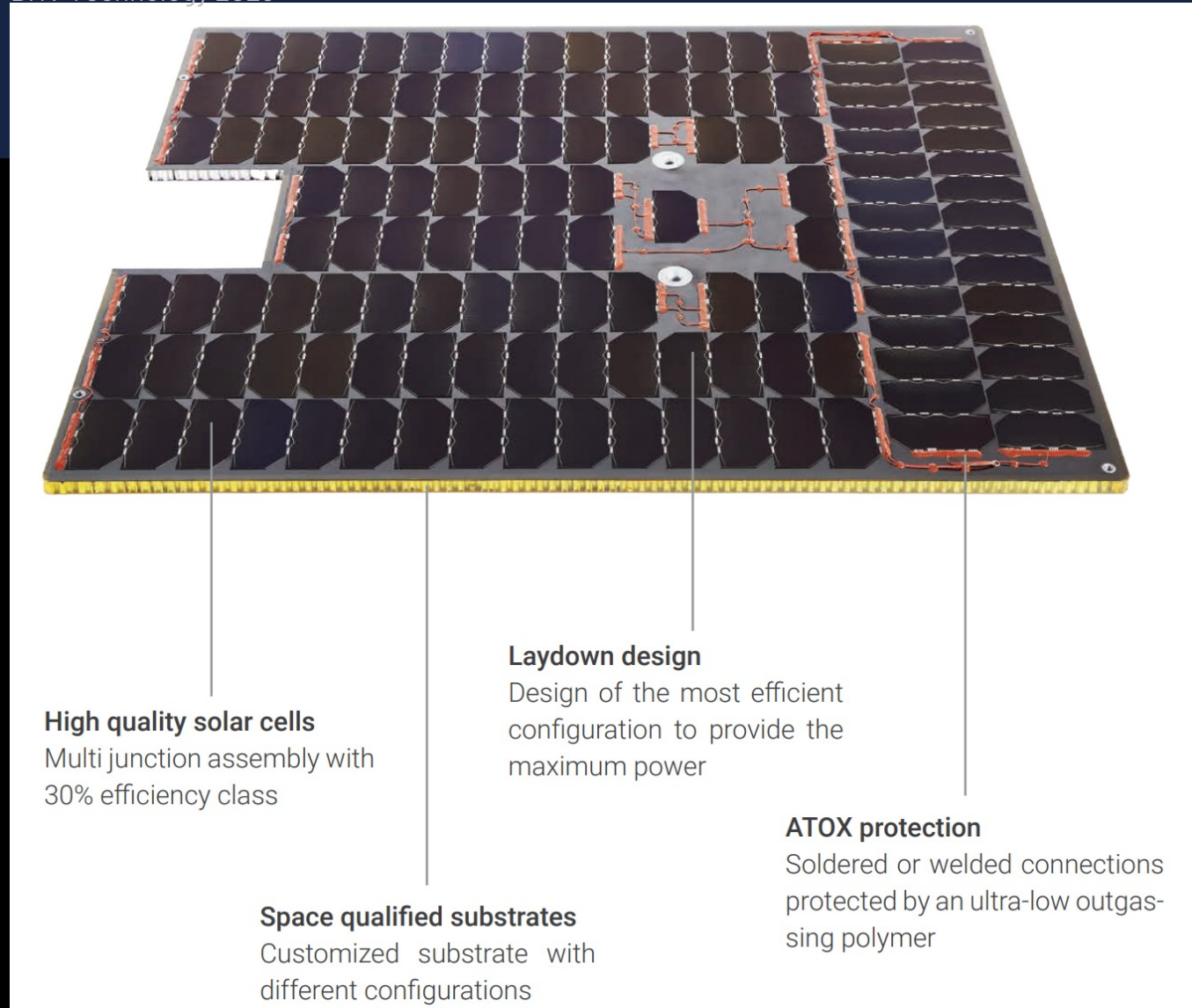


# SOLAR ARRAYS

**Company**

**DHV**

Technology	Multi junction
TRL	9
Efficiency	30 %
Power / Area	300 W / m2
Annual Degradation	2 %
Pointing Error	70 °
Size	20 x 505 x 644 mm
Max. Power Output (BoL)	1950 W
Min. Power Output (EoL)	545 W



**High quality solar cells**  
Multi junction assembly with 30% efficiency class

**Space qualified substrates**  
Customized substrate with different configurations

**Laydown design**  
Design of the most efficient configuration to provide the maximum power

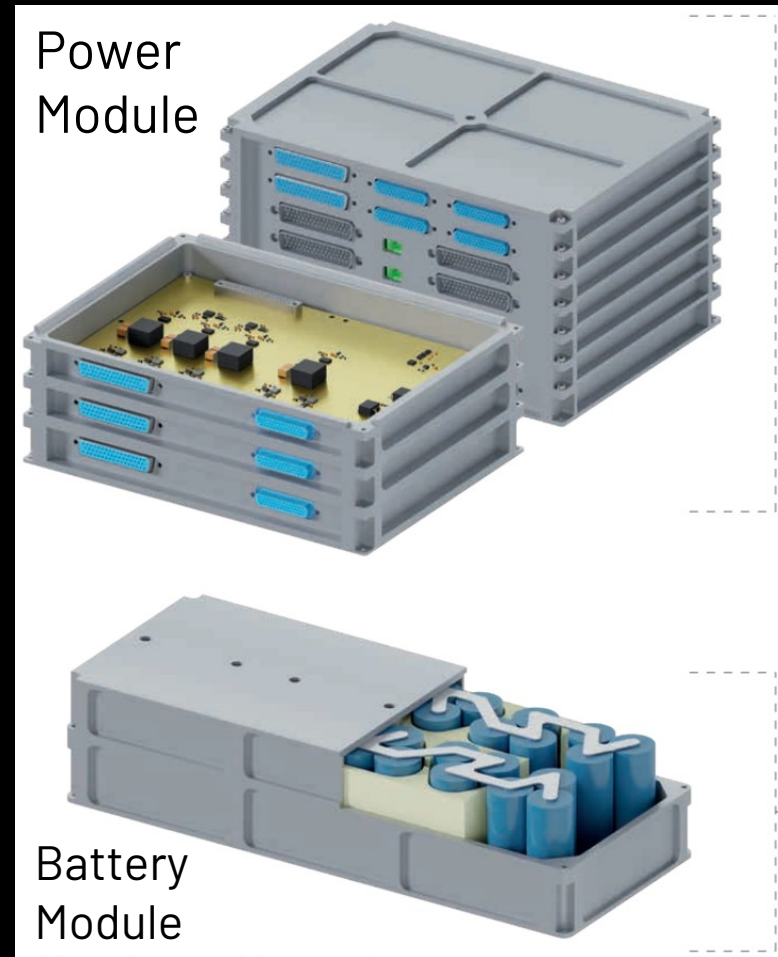
**ATOX protection**  
Soldered or welded connections protected by an ultra-low outgassing polymer



# POWER CONDITIONING DISTRIBUTION UNIT



Company	DHV
Battery Charge Regulators Efficiency	> 95 %
Power output efficiency	> 90 %
Output Buses	3.3 V, 5 V, 12 V, 28.8 V (unregulated)
Battery Type	LiFePO4
Capacity	1600 Wh
Features	<ul style="list-style-type: none"><li>• Maximum Power Point Tracker (MPPT)</li><li>• Over-Voltage/-Current Protection</li><li>• Health Checks, Temperature Control</li></ul>
TRL	9



DHV Technology 2023



# MASS BUDGET - OVERVIEW



System/Subsystem	Dry Mass	Mass Margin	Mass w/ Margin
<b>OBC (On-board computer)</b>			
Interfaces	4,0	5%	4,20
Casing	14,0	5%	14,70
OBC	2,0	10%	2,20
<b>Summary mass of subsystem</b>	<b>20,0</b>	<b>20%</b>	<b>25,32</b>
<b>EPS (Power)</b>			
Solar Panels	30,0	10%	33,00
Battery Packs	28,0	5%	29,40
Power Conditioning Distribution Un	2,6	5%	2,73
<b>STR (Structure)</b>			
Backplane (Main Mirror Supporting	646,5	20%	775,82
Secondary Mirror Beam Supporting	78,8	20%	94,50
Launcher Adapter	110,0	5%	115,50
Cables	327,0	5%	343,35
<b>Summary mass of subsystem</b>	<b>1162,3</b>	<b>20%</b>	<b>1394,72</b>
<b>ACS(Attitude Control System)</b>			
Reaction Wheels	35,8	10%	39,41
Gyroscopes	38,0	10%	41,80
Fine Guidance Sensor	233,6	10%	256,96
Sun sensors	0,2	5%	0,21
Star Trackers	8,6	5%	9,03
<b>Summary mass of subsystem</b>	<b>316,2</b>	<b>20%</b>	<b>379,48</b>

System/Subsystem	Dry Mass	Mass Margin	Mass w/ Margin
<b>COM (Communications)</b>			
Hk-data transceiver	15,0	10%	16,50
Payload Transmitters	15,0	10%	16,50
<b>Summary mass of subsystem</b>	<b>30,0</b>	<b>20%</b>	<b>36,00</b>
Thruster	3,7	10%	4,09
Propellant Mass	187,0	12%	173,12
Tank	28,1	5%	29,45
Pressurant Gas	3,0	10%	3,30
Mounting Hardware	46,8	10%	51,43
<b>THE (Thermal)</b>			
Radiators	8,9	10%	9,79
<b>Summary mass of subsystem</b>	<b>8,9</b>	<b>20%</b>	<b>10,68</b>
<b>RADIATION</b>			
Solar shielding	21,48	10%	23,63
<b>Summary mass of subsystem</b>	<b>21,48</b>	<b>20%</b>	<b>25,78</b>
<b>PAYLOAD</b>			
Main Mirror	106,3	10%	116,93
Secondary Mirror	0,7	10%	0,81
Tertiary mirror	0,03	10%	0,03
Fine steering mirror	0,01	10%	0,01
Spectrograph unit	100,0	20%	120,00
Imager unit	70,0	20%	84,00
Mirror Thermal Control Unit	15,8	20%	18,96
<b>Summary mass of subsystem</b>	<b>292,9</b>	<b>20%</b>	<b>351,44</b>
<b>Totals</b>	<b>2180,87</b>	<b>20%</b>	<b>2617,04</b>

# MASS BUDGET – KEY DRIVERS

ELEMENT	MATERIAL	MASS [KG]
BACKPLANE	M55J GRAPHITE EPOXY	776
SECONDARY MIRROR SUPPORT	LIGHTWEIGHT CFRP	95
MIRRORS	BERYLLIUM SUBSTRATE, ALUMINIUM + LF2, COATING	118
PROPELLANT	HYDRAZINE	173

# TECHNOLOGY ROADMAP

Comms, OCS, ADCS, OBC, TCS, EPS



TRL 9

EMCCD (with mitigation)



TRL 8

Electron Multiplying Charge-Couple Device (EMCCD)



TRL 5

TRL 4

TRL 3

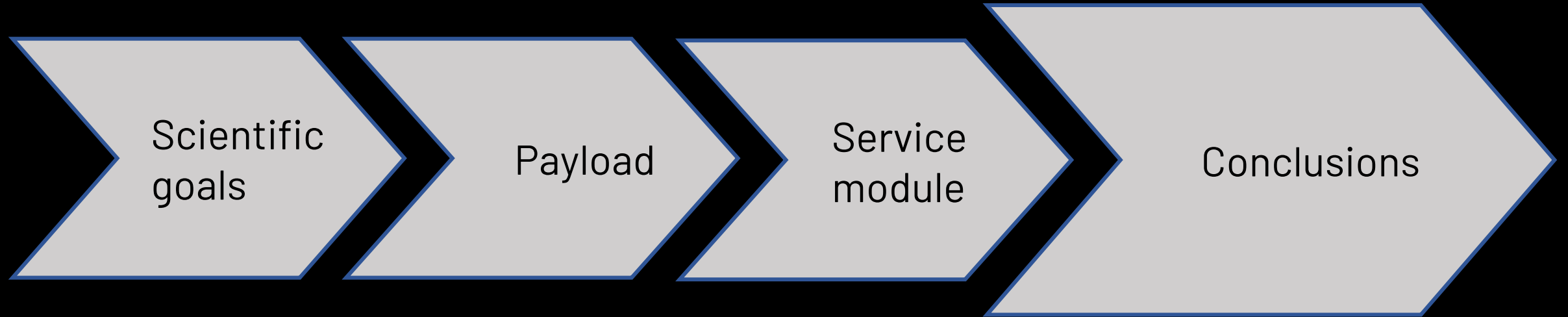
TRL 2

TRL 1

(Robert Jacobsen, NASA)



# CONCLUSIONS



# COST CATEGORY



Subcategory	Category	Cost [M€]
Platform	Industrial cost	450
Telescope & payload instruments	Industrial cost	450
	ESA project cost	170
	Operations	80
	Total	1150
	Total (20 % margin)	1380
	launch	100
Delivery to L2	Payload provided by ESA Member States	-10
	Summary	1420

# CRITICAL RISKS- BEFORE MITIGATION

**Risk Index**

Very High
High
Medium
Low
Very Low

**Mitigation:**

- Include Detector on pathfinder mission, i.e., a CubeSat, before EXPOSURE to flight qualify (TRL 8)
- Redundancy and FDIR for OBC
- Increase exposure time if mirror surface degrades

Severity	1					
	2				Single-Event Upset (SEU)	
	3			Collision with micrometeoroid		
	4			Degradation of Mirror Surface Pointing jitter		
	5		Solar-array deployment failure Boom deployment failure	Redirection of funds	Detector (TRL 5)	
		A - Remote	B - Unlikely	C - Likely	D - Highly likely	E - Near certain

**Likelihood**

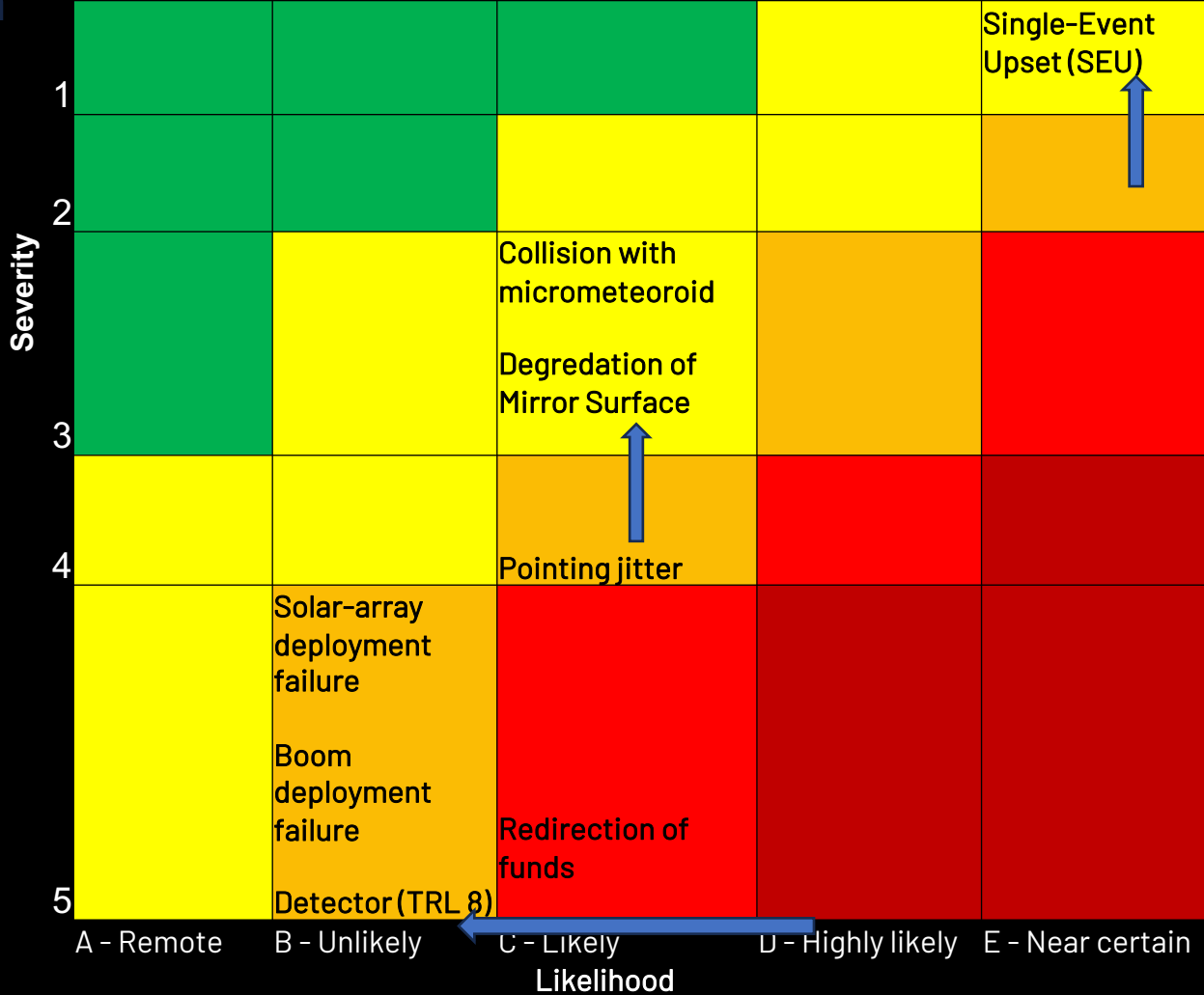


# CRITICAL RISKS – AFTER MITIGATION



**Mitigation:**

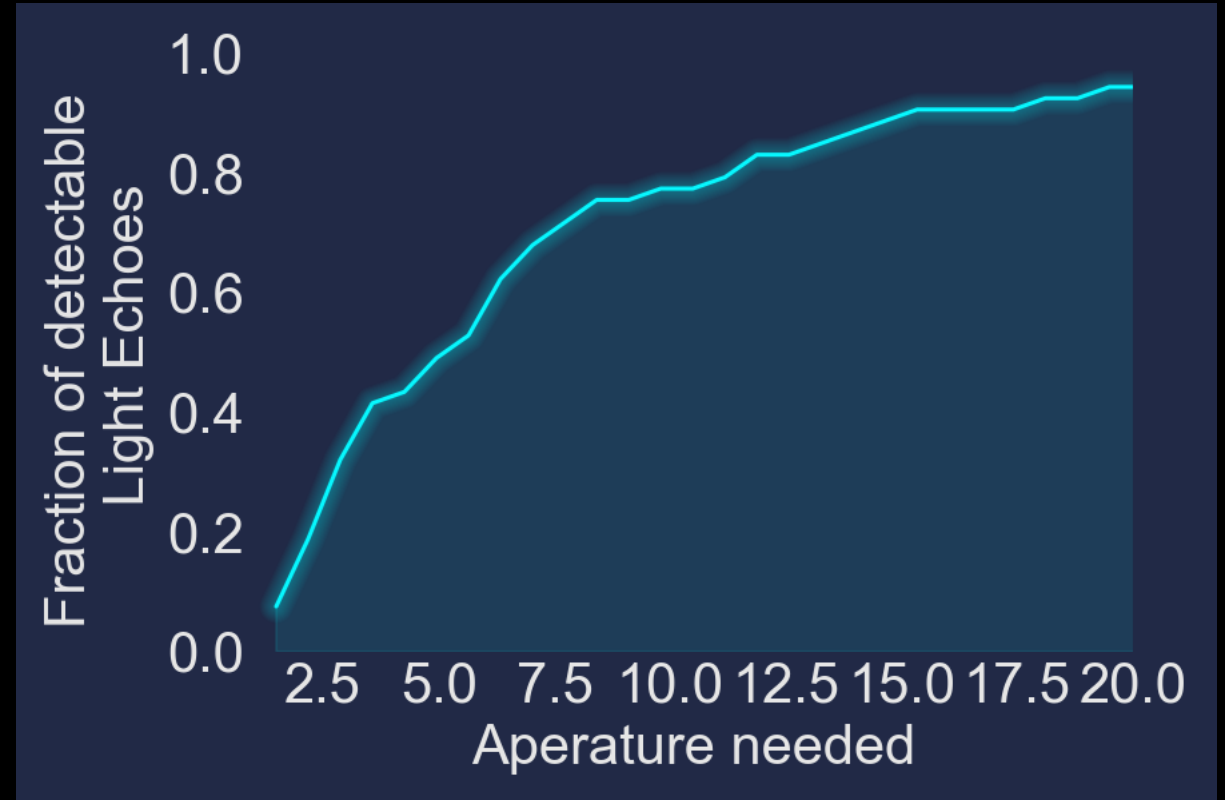
- Include Detector on pathfinder mission, i.e., a CubeSat, before EXPOSURE to flight qualify (TRL 8)
- Redundancy and FDIR for OBC
- Increase exposure time if mirror surface degrades



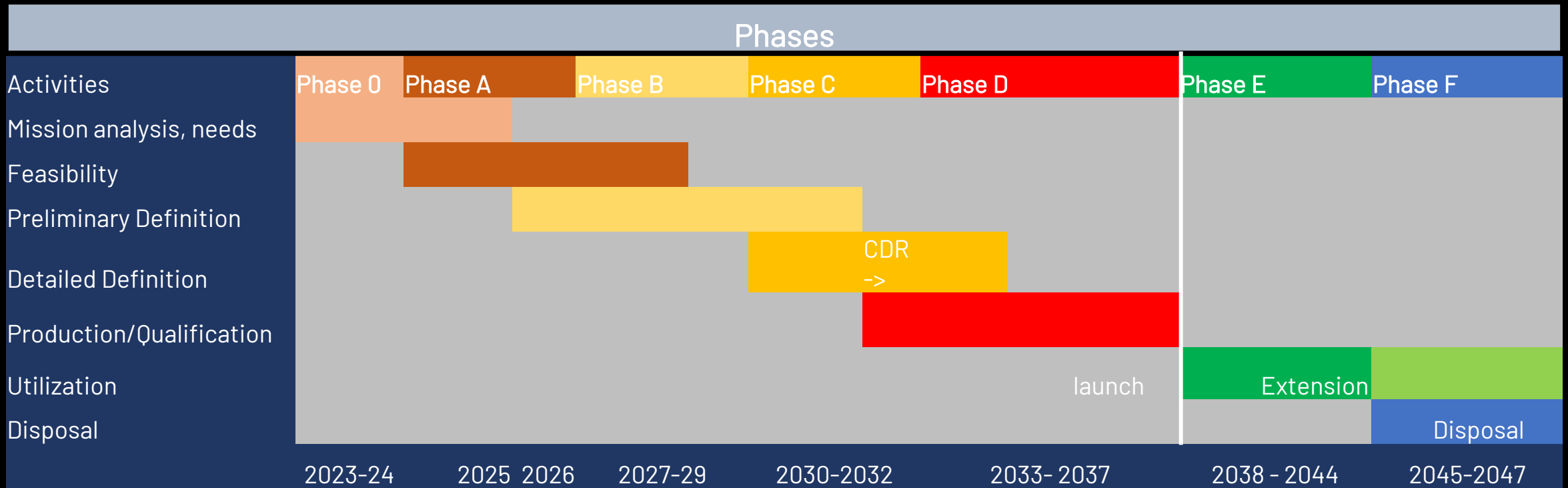
# DE-SCOPING POSSIBILITIES



- **Shrink telescope diameter** -> lose light echo detection and possibly exospheres of habitable planets
- **Take out imager** -> lose UV-flux and light echo detection
- **Take out spectrograph** -> lose detection of escaping and extended atmospheres

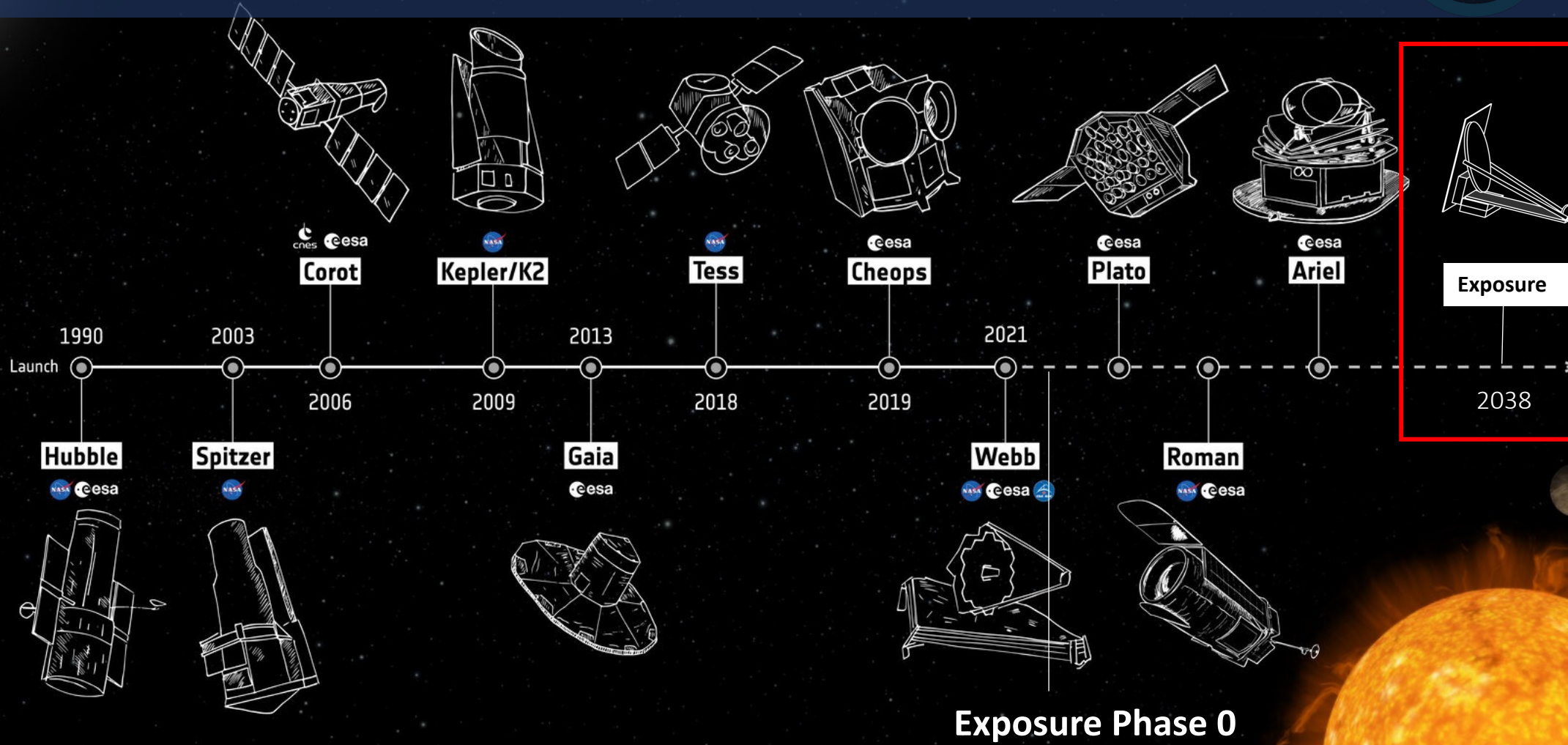
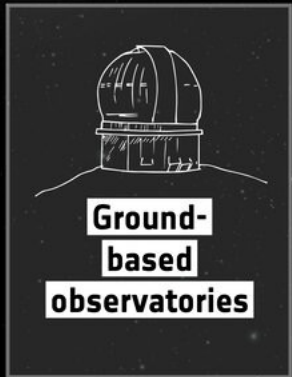


# DEVELOPMENT SCHEDULE





# ESA MISSION TIMELINE



# PUBLIC OUTREACH



- Instagram
- Facebook
- Course of funfact sessions at schools
  - Variations in exoplanets
  - Dangers of UV-radiation from the Sun



# FINAL REMARKS

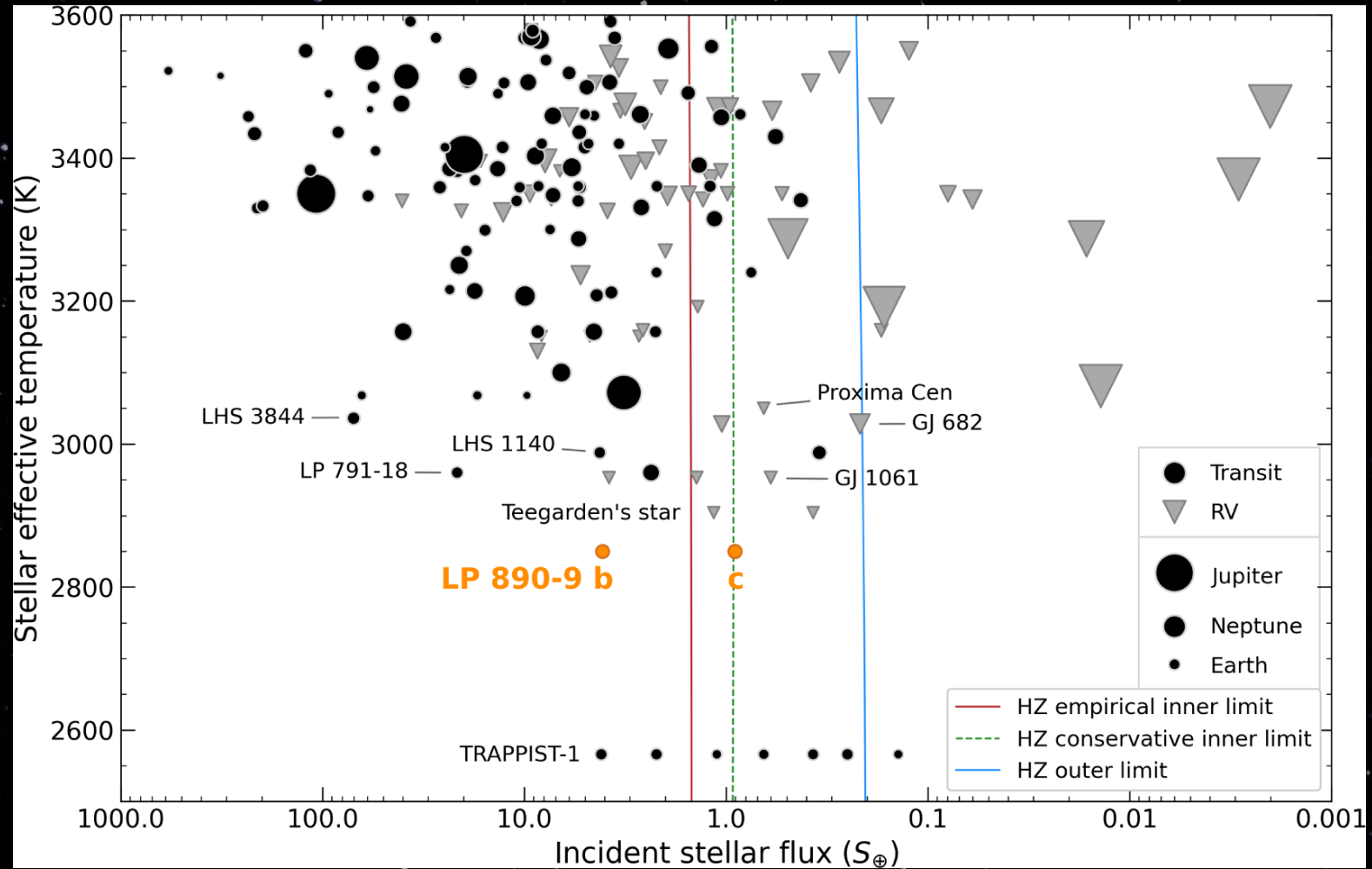


Summer School Alpbach 2023 – Team Blue





# BACKUP: Habitability Zone (HZ)



- Clouds can have perfect temperature for life
- Temperature below clouds/atmospheres → life
- Exomoons in CHZ
- Atmospheres potentially harbor life f.e. upper-to-middle cloud layers of Venus (0–60°C; pH~0): might suitable for thermo- or psychro-acidophilic microorganisms

(Merino, N., (2019), Living at the extremes: Extremophiles and the limits of life in a planetary context, *Frontiers in Microbiology*, 10)

# Backup Slides - Requirements

Req-ID	Traceability Source		Description
M-01	O-1	Measurement	Measure the Lyman alpha transit of exoplanets orbiting in the habitable zone
M-02	O-2	Measurement	Measure the Lyman alpha radial velocity profile of exoplanets of planets with expanding hydrogen atmospheres
M-03	O-3	Measurement	Measuring variations in the UV emission of stars with planets in the habitable zone
M-05	M-01,M-02	Measurement	The spacecraft shall perform spectroscopy around the Lyman-alpha line while exoplanets transit
M-06	M-01	Measurement	The mission shall do spectroscopy around the lyman alpha line on stars with planets which are supposed to have evaporating atmospheres
M-07	M-03	Measurement	The mission shall perform photometry on M-Type stars
S-01	O-4	Scientific	Separate flare emissions from the star from the reflected emission of the flares on an exoplanet
S-02	M-02	Scientific	The observed exoplanets shall be larger than 1.5 earth radii for case 1.1
S-03	M-01,M-02	Scientific	The observed exoplanets shall orbit around M, K or G-type stars for case 1.1 and 1.2
S-04	M-01	Scientific	The observed exoplanets shall be smaller than 2.5 earth radii for case 1.2
S-05	M-01	Scientific	The insulation flux received by the observed planet for case 1.2 shall be between 0.2 and 2 time the solar flux
S-06	M-02	Scientific	The radial velocity of the Lyman alpha line shall be resolved to be able to distinguish the red wing and blue wing
S-07	S-01	Scientific	The distance between stars and planets for observed targets shall be larger than 0.006 AU
S-08	S-01	Scientific	The mission shall observe more than 30 % of M-dwarfs with the within the observed region for case 3
S-09	S-01,M-02,M-03,M-04	Scientific	The observed targets shall be closer than 100 psc
S-10	M-02	Scientific	The minimum number of targets that shall be observed is 150 planets for case 1.1



# Backup Slides - Requirements

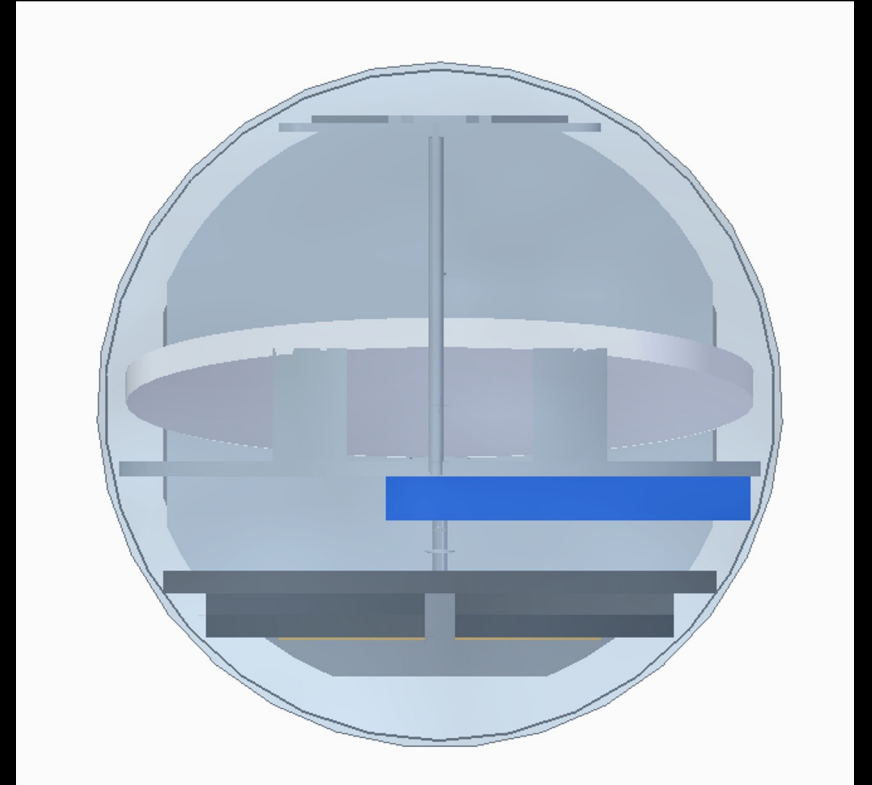
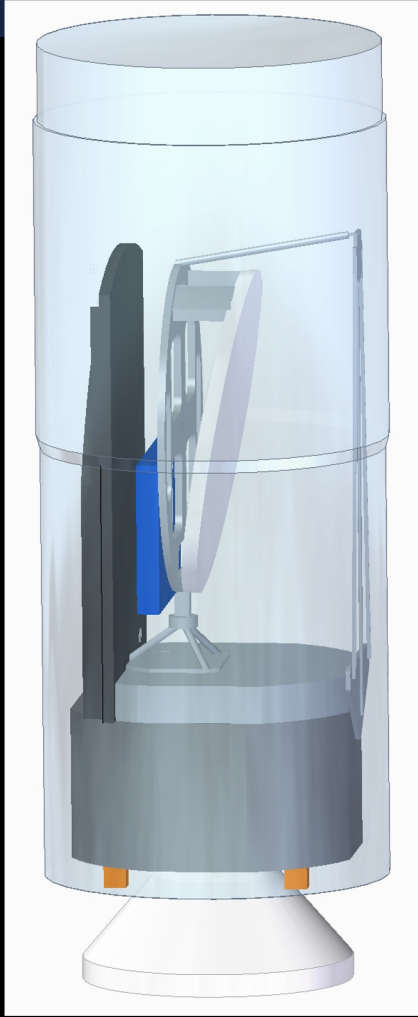
Req-ID	Traceability Source		Description
S-10	M-02	Scientific ▼	The minimum number of targets that shall be observed is 150 planets for case 1.1
S-11	M-01	Scientific ▼	The minimum number of targets that shall be observed is 20 planets for case 1.2
S-12	M-03	Scientific ▼	The minimum number of targets that shall be observed is 150 stars for case 2
S-13	S-01	Scientific ▼	The minimum number of targets that shall be observed is 27 planets for case 3
S-14	M-02	Scientific ▼	There shall be no geocoronal absorption visible in the observed spectra
S-15		Scientific ▼	3 transits shall be observed for each target
S-16		Scientific ▼	Each star shall be observed for 15 days (for it to be likely to observe a flare)
S-17		Scientific ▼	The spectral resolution for case 1.1 shall be larger than 20000
S-18		Scientific ▼	The spectral resolution for case 1.2 shall be larger than 1000
S-19		Scientific ▼	The field of view shall be at least 25"
S-20		Scientific ▼	The f/ratio of the telescope shall be 1/ (100 +/-0.1)
S-21		Scientific ▼	The pointing accuracy shall be 0.005arcsec
S-22	S-15	Scientific ▼	The stars shall be observed for more than one rotation of the star
S-23		Scientific ▼	The minimum uninterrupted observation time shall be 16 hours
S-24		Scientific ▼	The S/C shall be able to store and transfer 240Gb of data pr. day.
S-25	M-03,S-01	Scientific ▼	The UV-flux-variations observable shall be smaller than 0.1 %



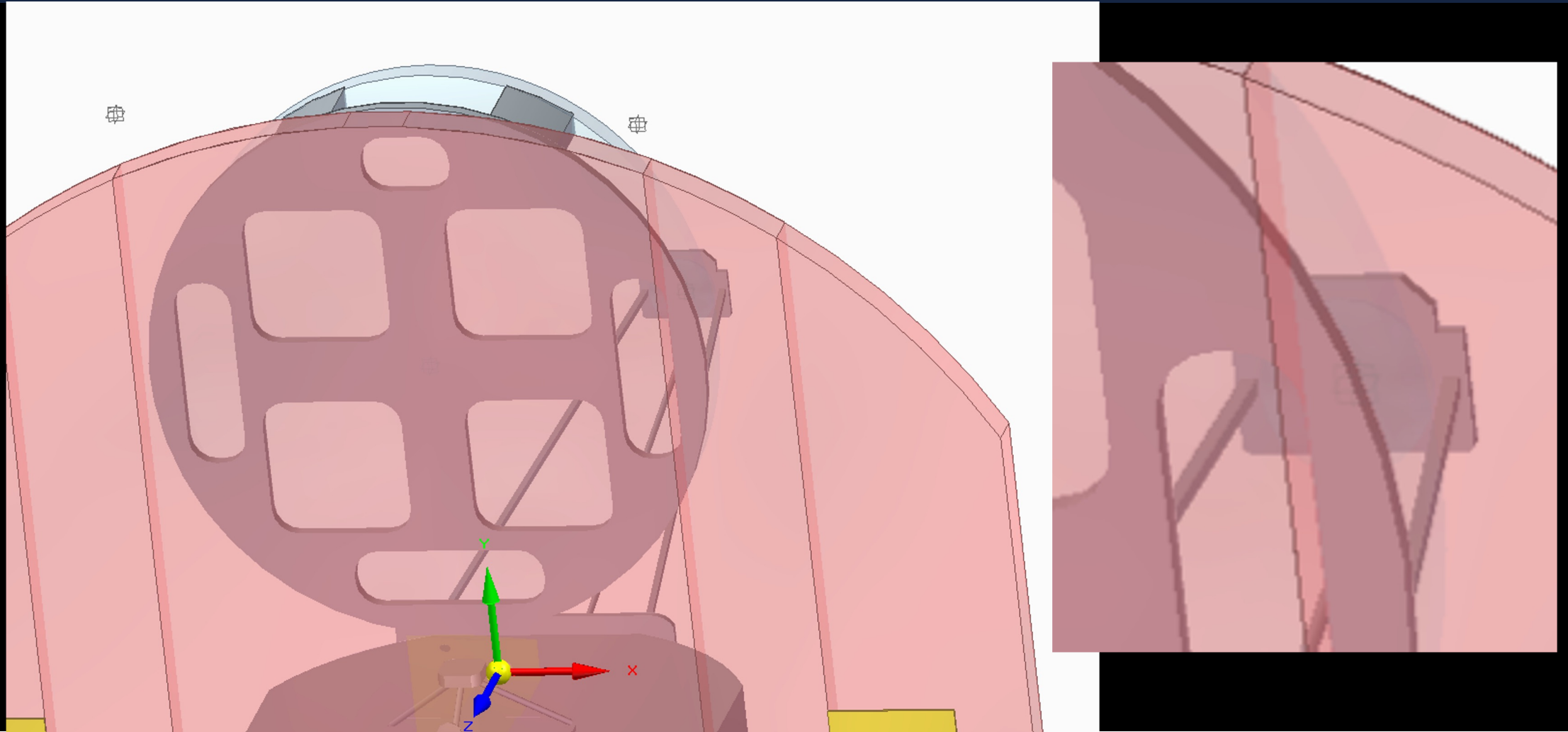
# Backup Slides - Requirements

Req-ID	Traceability Source		Description
I-03		Instrument	The detector for the imager shall have a Minimum Quantum Efficiency of 0.4 between 150+/-1 nm -300+/-1 nm
I-04		Instrument	The detector for the imager shall have a Minimum Quantum Efficiency of 0.2 between 115+/-1 nm -400+/-1 nm
I-05		Instrument	The detector for the spectrograph shall have a Minimum Quantum Efficiency of 0.6 between 115+/-0.5 nm and 130+/-0.5 nm
I-06		Instrument	The spectrograph detector shall have 2048x10 pixels
I-07		Instrument	The spectrograph pixel size shall be 7µm +- 0.01 µm
I-08		Instrument	The imager detector shall have a minimum of 2048x10 pixels
I-09		Instrument	The imager pixel size shall be 7µm +- 0.01 µm
I-10		Instrument	The throughput shall be larger than 4 %
I-11	C-03,	Instrument	The smallest integration time shall be 1s
I-12	S-21, T-03, T-05, T-06, T-09	Instrument	The roughness of the reflecting surfaces shall be smaller than 0.1
I-13	S-21, T-03, T-05, T-06, T-09	Instrument	The wavefront error shall be smaller than 7 nm
I-14	T-02, T-06, T-12, T-13	Instrument	The primary mirror shall have a Diameter of 4.2

# BACKUP: SPACECRAFT IN LAUNCH CONFIGURATION

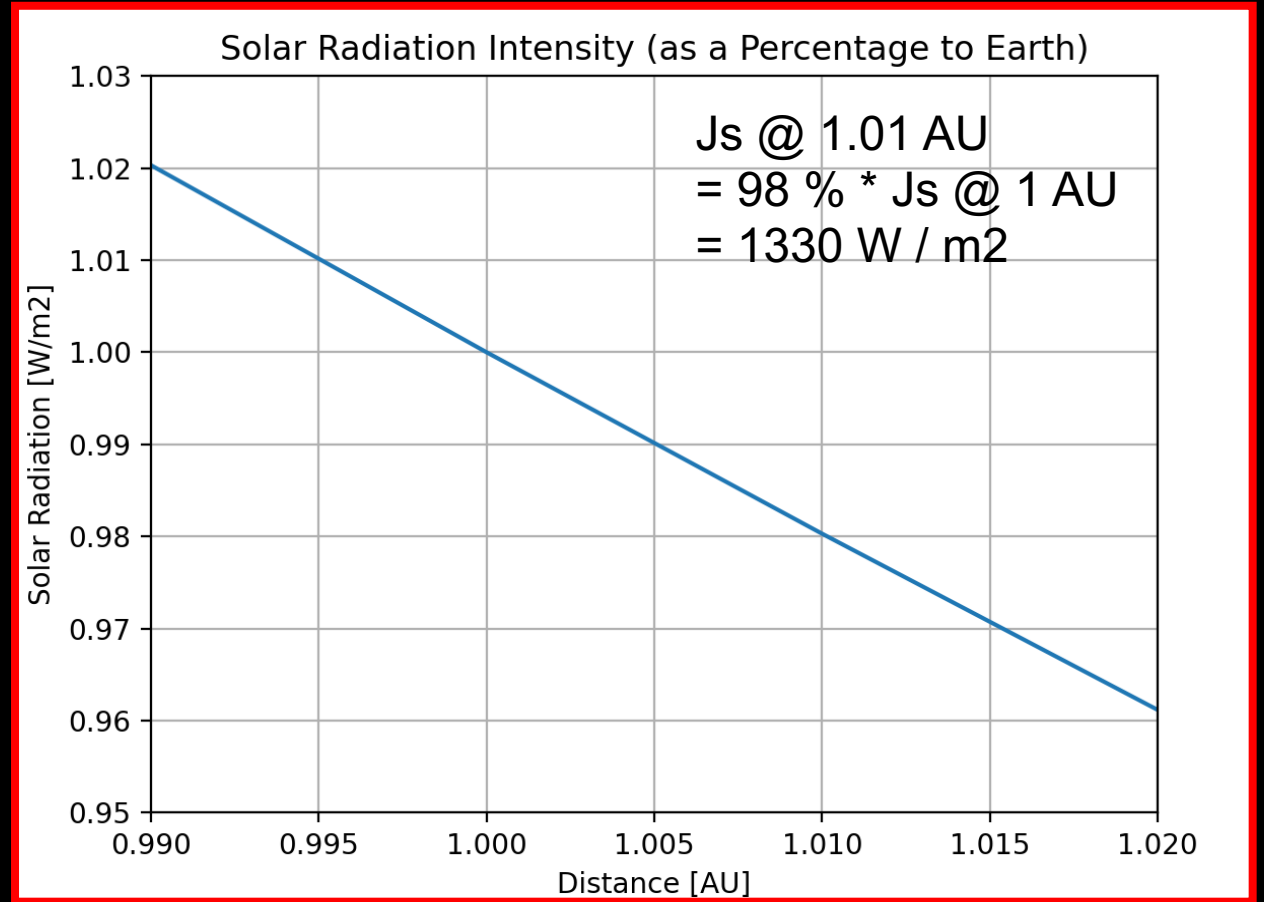
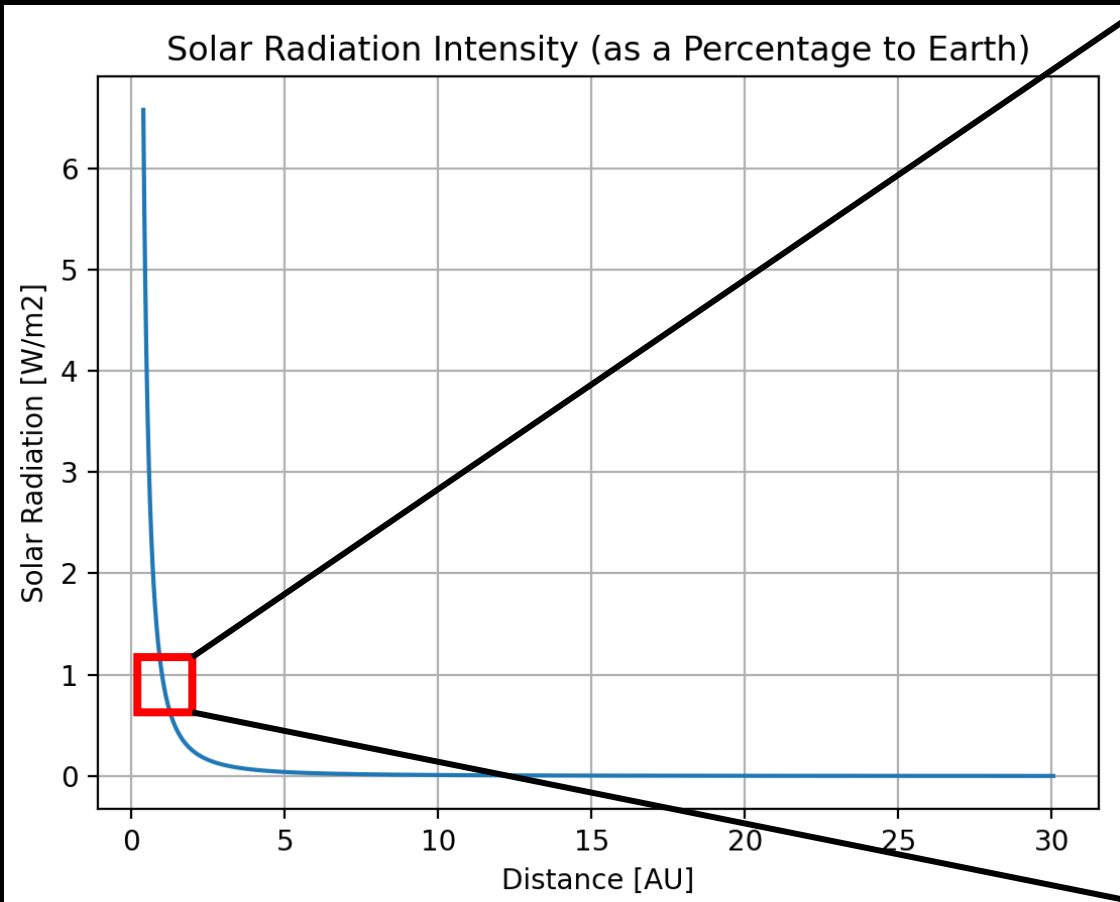


# BACKUP: 70 ° from Ecliptic





# BACKUP: SOLAR RADIATION FLUX



# BACKUP: 50 M/S MANEUVER DURATION

