# ORPHEUS

# Team Orange

# **Mission Statement**

"Study the internal magnetic field of the Earth's core on a global scale in order to better understand its fine structure dynamics"

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ORTHINGS MISSION

ALPENCH

# WHY ORPHEUS?

• Long term solution for a long-term scientific goal

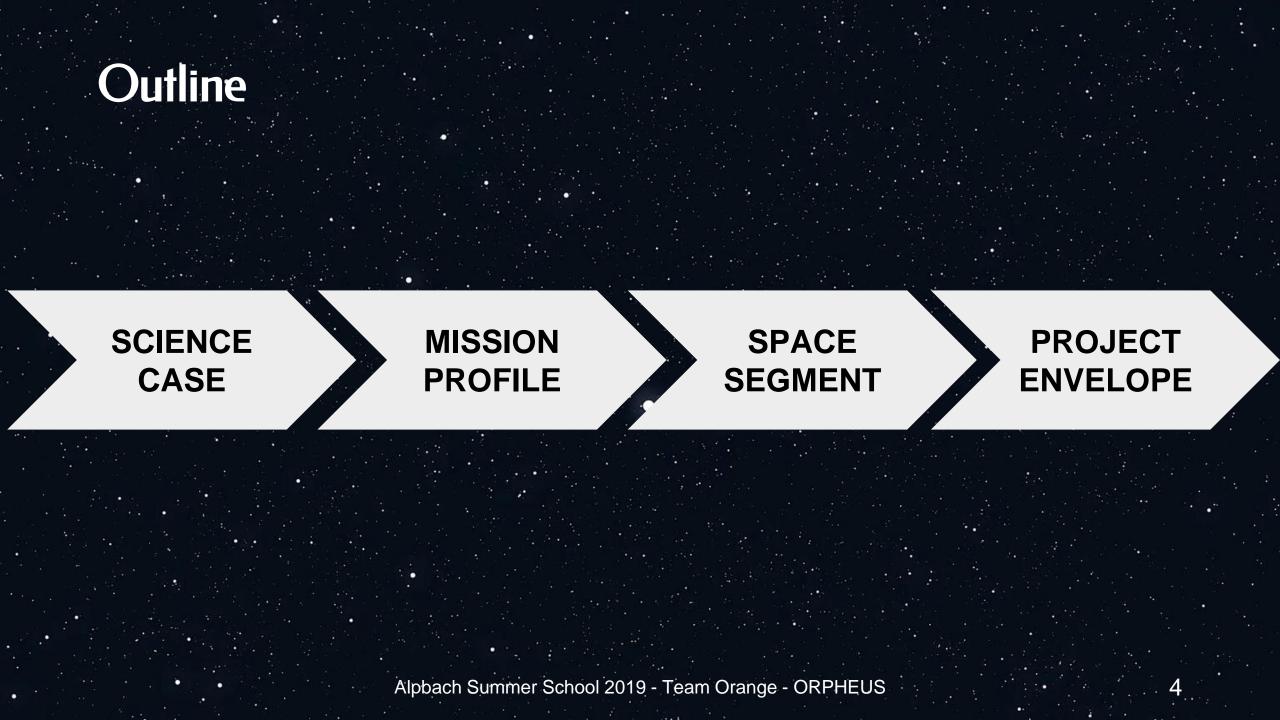
- Continuing the historical record for Earth's magnetic field
- An innovative formation enabling sophisticated measurements in an underexplored atmospheric region

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ORDINELS MISSION

ALPENCH 20





**SCIENCE** 

CASE

<u>SPEAKER:</u> <u>ELENA L. CONTRERAS</u>

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**SPACE** 

**SEGMENT** 

**PROJECT** 

**ENVELOPE** 

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**MISSION** 

**PROFILE** 

## Importance of the Magnetic Field

Measurements of the <u>core dynamics</u> of Earth are imperative to understand the magnetic field of Earth

Constant update of <u>physical models</u>

The magnetic field is <u>vital for life</u> on Earth and <u>modern society</u> depends heavily on its stability

# Earth's Structure

CRUST

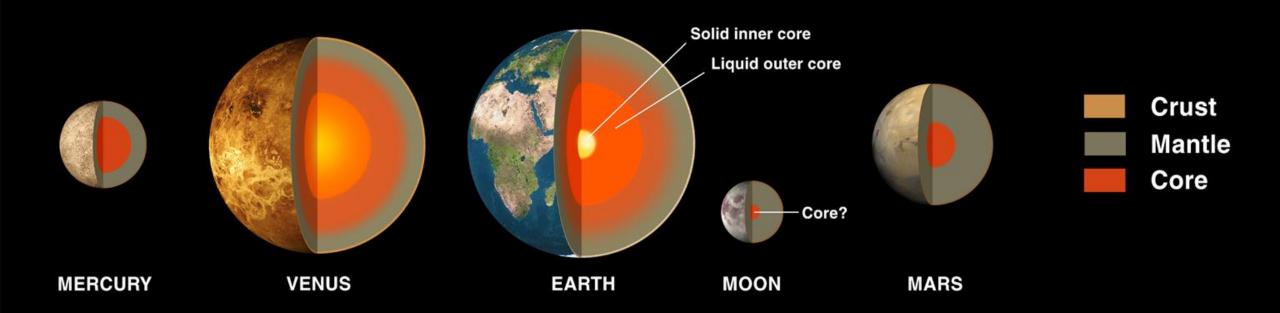


MANTLE

• OUTER CORE

(Structure of the Earth (2019))

## **Geophysics of Other Planets**



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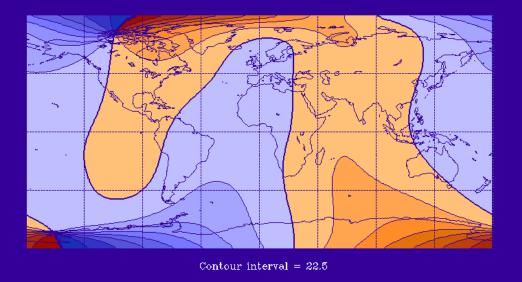


(ETH Zürich. Planetary Science (2019))

## Time Varying Magnetic Field

#### **Magnetic Declination**

1590



**20**0 - 167.5 - 135.0 - 112.5 - 90.0 - 47.5 - 45.0 - 52.5 0 22.5 40.0 67.5 90.0 112.5 135.0 157.5 180.0

#### Time varying representation based on data from 1590-1990

 Earth's magnetic field varies over long timescales

Courtesy of Andrew Jackson

## **Magnetic Field Polarity Reversal**

**Normal Field Reversal Process Reversed Field** A three-dimensional self-consistent computer simulation of a geomagnetic field reversal (Glatzmaier and Roberts. 1995) 10

# SO-1.Provide state of the art measurements of the fine structure and<br/>dynamics of the Earth's core (long time scale)

.SO-2.

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Separate the contributions from the various other magnetic field

sources

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Separate the contributions from the various other magnetic field sources

SO-3.

.SO-2.

Study the ionospheric current system contribution to the total magnetic field

SO-1.

SO-3.

**S()-**4

Provide state of the art measurements of the fine structure and dynamics of the Earth's core (long time scale)

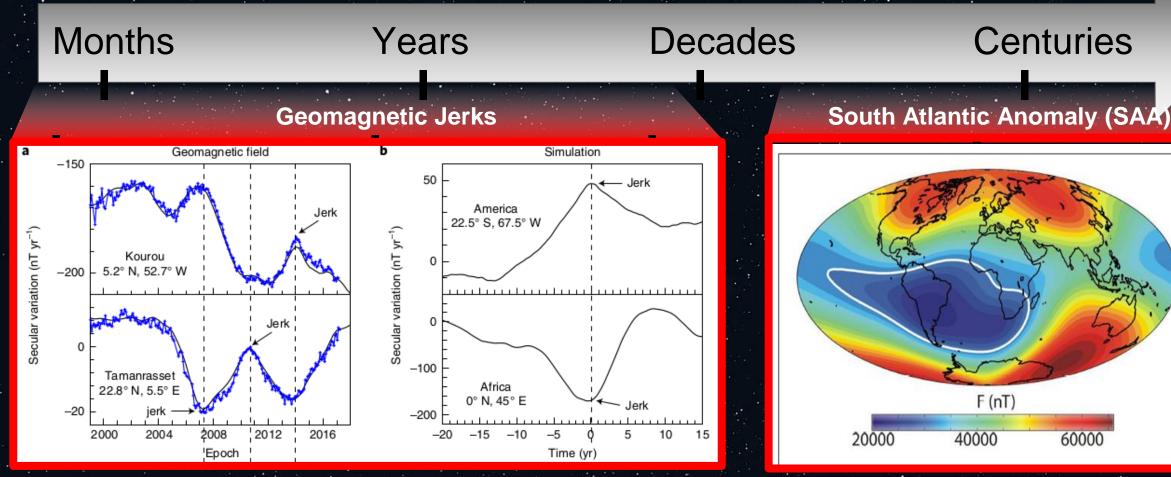
SO-2. Separate the contributions from the various other magnetic field sources

Study the ionospheric current system contribution to the total magnetic field

Characterize and link dynamic events of the geomagnetic field to core processes (shorter time scale) Alpbach Summer School 2019 - Team Orange - ORPHEUS

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### Timescales



Observed and simulated geomagnetic jerks at Earth's surface. Left: observed data. Right: model data

(Aubert. C Finley. 2019)

(F. Javier Pavón-Carrasco and A. De Santis. 2016)

15

Intensity geomagnetic field map at 2015.0

SR-1.

SO-1.

#### Measure all three components of the magnetic field

vector

SR-1.

SR-2.

SO-1.

#### Measure all three components of the magnetic field

17

vector

#### Measure the magnetic field magnitude

SO-1.

#### Measure all three components of the magnetic field

vector

#### Measure the magnetic field magnitude

SR-3.

SR-2.

SR-1.

Earth's magnetic field shall be continuously monitored

SR-4.

SO-2.

#### Remove contributions of non-core internal fields

SO-2.

#### Remove contributions of non-core internal fields

Remove contributions from external fields

SR-5.

SR-4.

#### Remove contributions of non-core internal fields

#### SR-5.

SR-6.

SR-4.

SO-2.

#### Remove contributions from external fields

#### Measurement duration of 25 years

SO-3.

SR-7.

#### Estimate the ionospheric current density

#### SO-3.

 $\cdot$ SO-4.

#### SR-7.

#### Estimate the ionospheric current density

#### SR-8.

#### Global coverage

#### SO-3.

 $\cdot$ SO-4.

#### SR-7.

#### Estimate the ionospheric current density

#### SR-8.

### Global coverage

SR-9.

#### Orbital revisit period shall be shorter than 2 weeks

### So why do we need satellites?

**GLOBAL COVERAGE** 

CRUSTAL MAGNETIC FIELDS & ONE OF THE MAJOR CURRENT SYSTEMS 25

ALTITUDE

> 400km

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71% of EARTH'S SURFACE IS WATER

#### Magnetosphere

lonosphere

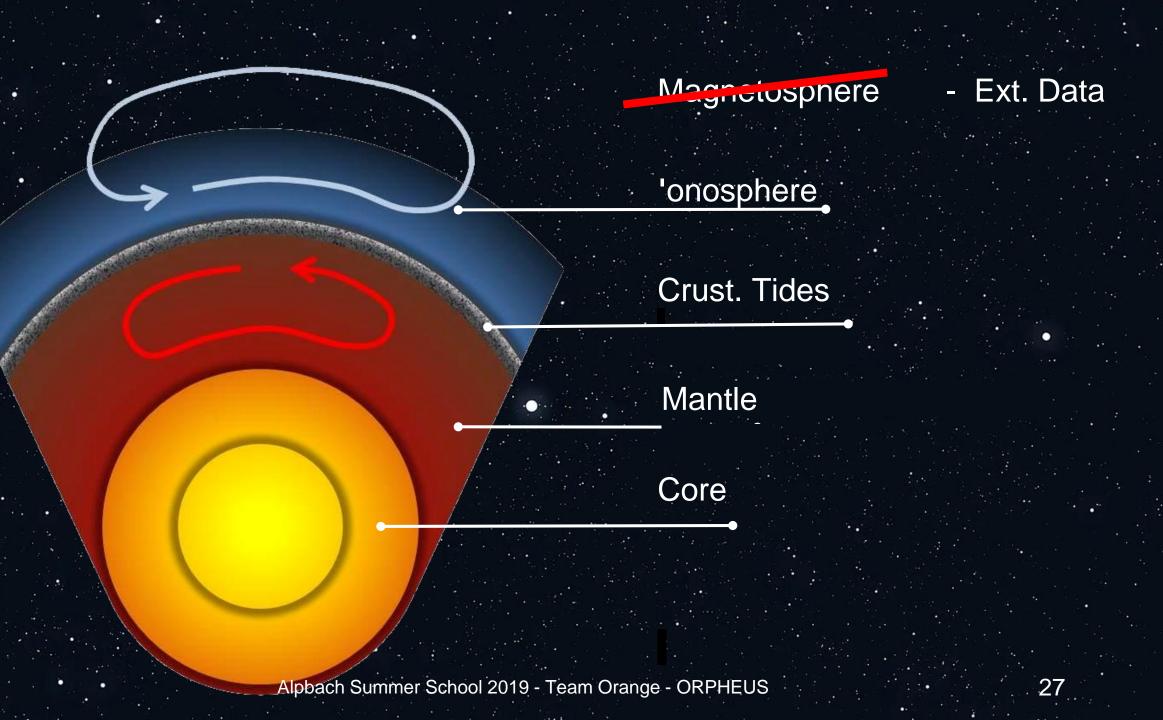
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Crust. Tides

Mantle

Core

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#### Ioncophere - Measure

28

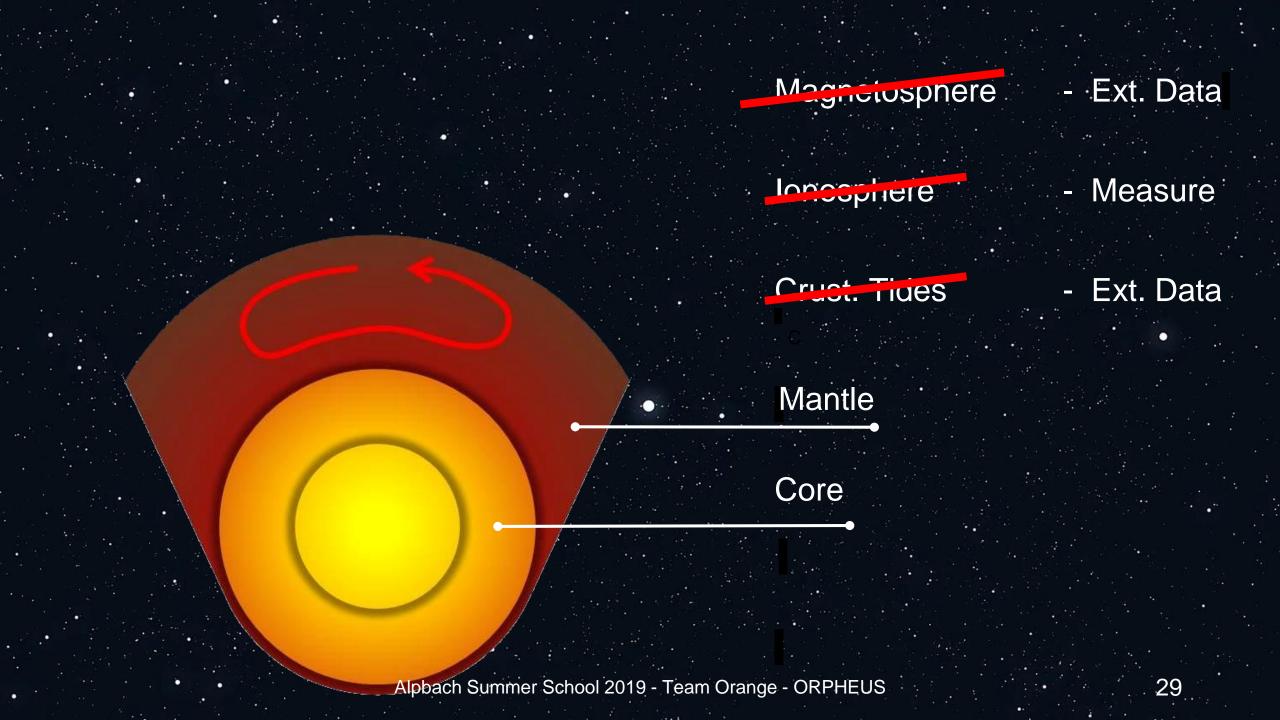
Crust. Tides

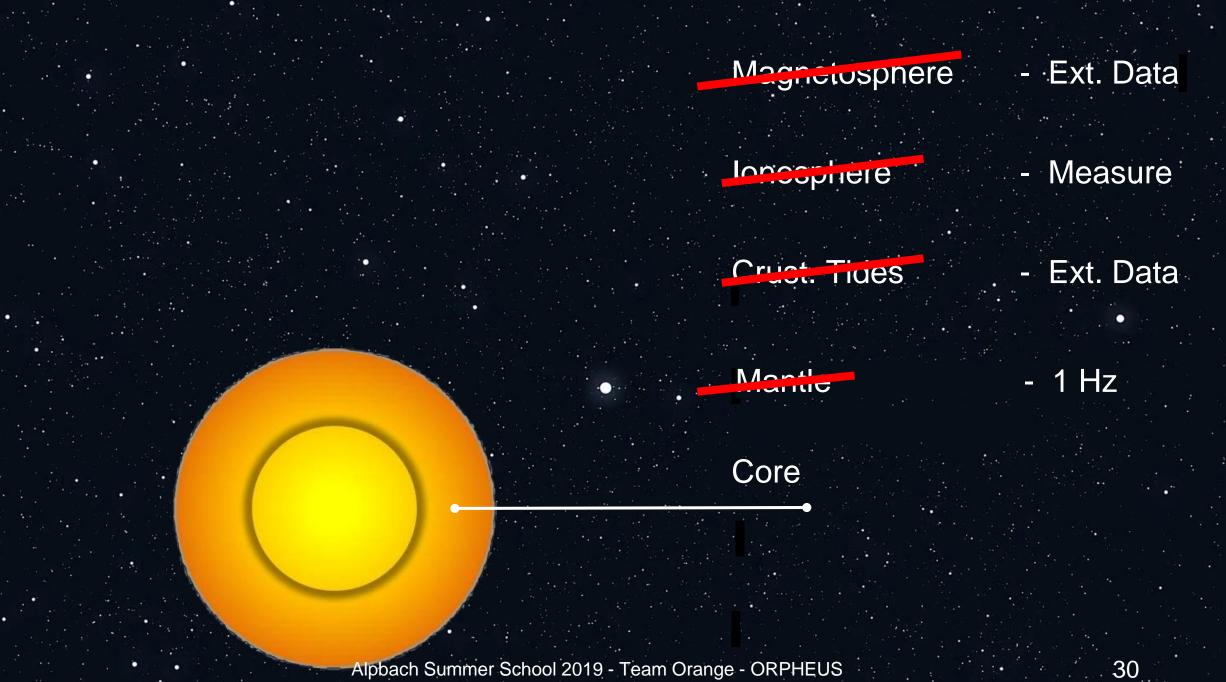
Mantle

Core

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## **Mission Drivers**

Measure all three components of the magnetic field vector

• Measure the magnetic field magnitude

• Determine the magnetic field gradients

• Estimate the ionospheric current density close to the satellite

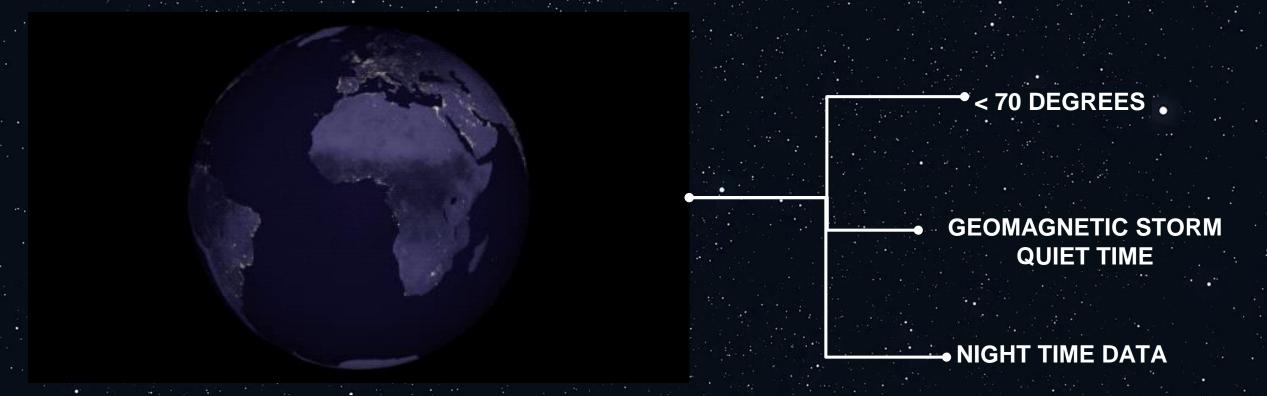
### Past. Present & Future Missions

	POGO scalar only		MAGSAT vector and scalar only				ØRSTED. CHAMP. SAC – C		vector and scalar only	SWARM	3-sat Constellation		SWARM FOLLOW - ON			OKPHEUS MISSIUN	vector and scalar only	
							SWARM	prep	paration		RM FO aration		ORPHEU MISSION reparatio	V				
19	965 1970	1975	1980	1985	1990	1995	2000	200	5 2010	) 20	)15 2	020	2025	203	30 2	203	5	2060

32

## **Historical Missions**

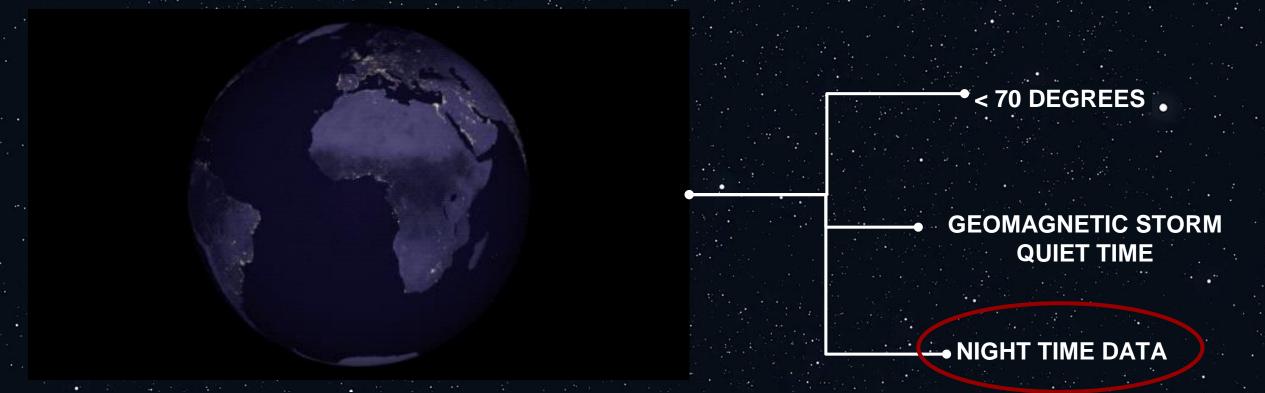
Previous missions have used the following constraints to study the core dynamics:



We intend to estimate ionospheric currents for the purpose of using day time data

## **Historical Missions**

Previous missions have used the following constraints to study the core dynamics:



We intend to estimate ionospheric currents for the purpose of using day time data

## **Spherical Harmonics**

 Earth's magnetic field can be described in terms of a spherical harmonic series expansion.

 Retrieval of the magnetic field data

(University of Texas. Teaching: Spherical Harmonics (2019))

# Satellite Configuration

- SWARM retrieves core field up to order
- ORPHEUS sensitive to core field order
  - $\rightarrow$  Satellite separation of 7 ± 0.5° horizontally
    - $= 860 \pm 60$  km separation

m = 18

m = 25

# Satellite Configuration

Current density at 700 km estimation from SWARM data

 $egin{aligned} & \mathbf{\nabla} imes \mathbf{B} = \mu_0 \mathbf{J} \ & \mathbf{L} &pprox rac{\delta \mathbf{B}}{\mu_0 \mathbf{J}} \end{aligned}$ 

From ionospheric current density:

20 - 100 km

 $\rightarrow$  satellite separation of <u>50 km for good signal-to-noise ratio</u>

## Satellite Configuration

#### TOTAL OF 5 + 1 SATELLITES

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50 KM

7.0 ± 0.5 °

 $(860 \pm 60 \text{ KM})$ 

FORMATION



**SCIENCE** 

CASE



**MISSION** 

PROFILE

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**SPACE** 

**SEGMENT** 

**PROJECT** 

**ENVELOPE** 

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## **Mission Profile**

#### SYSTEM DRIVERS

MISSION REQUIREMENTS

03

01

04

02

ORBIT & LAUNCH SELECTION

MISSION PHASE & TIMELINE

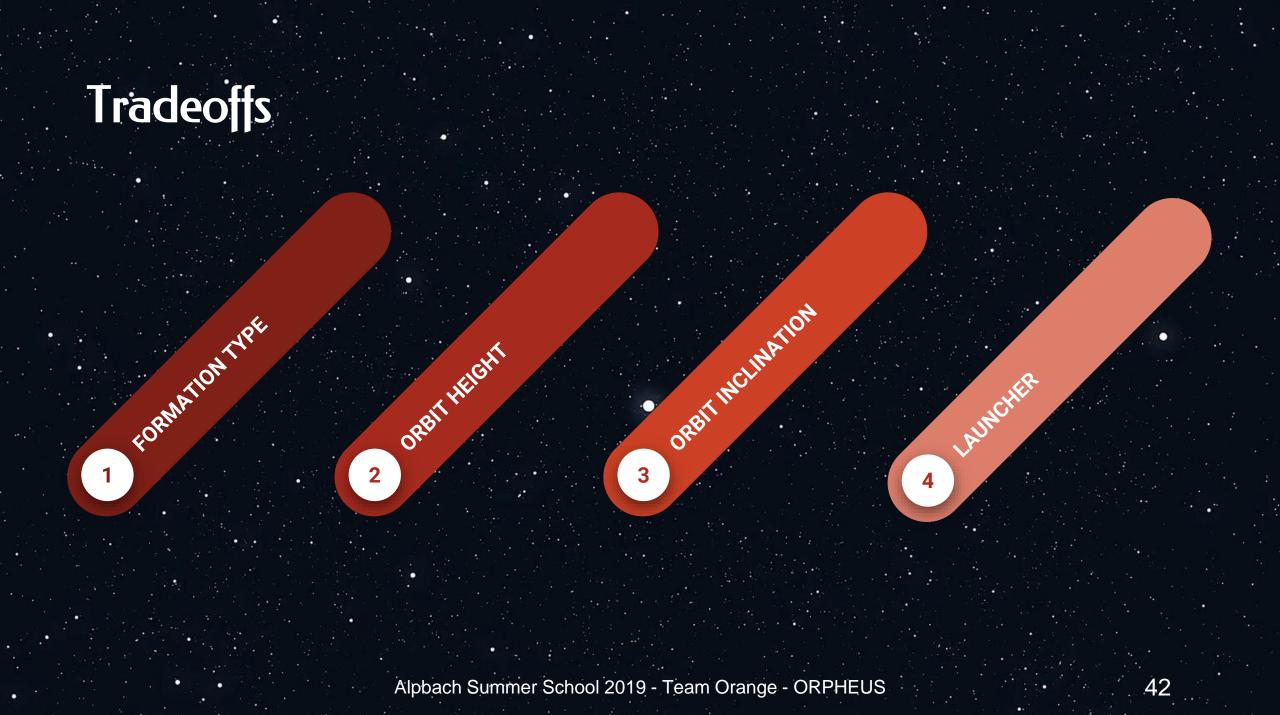
40

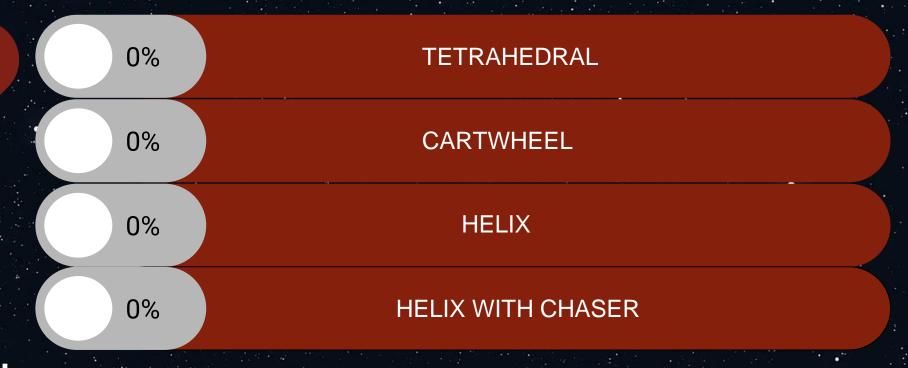
# System Drivers

01	ORBITAL CHARACTERISTICS	<ul> <li>Coverage</li> <li>Lifetime</li> <li>Radiation density</li> </ul>
02	CONSTELLATION CHARACTERISTICS	<ul><li>Quality of measurements</li><li>Formation stability</li></ul>
03	INSTRUMENT HOSTING	<ul> <li>Size</li> <li>Magnetic cleanliness</li> <li>Mechanical &amp; thermal needs</li> </ul>

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#### **Selection Criteria:**

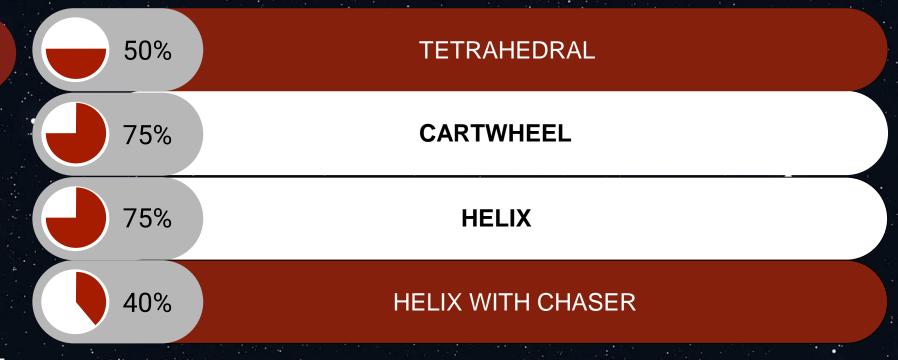
- □ Magnetic field accuracy
- □ Formation stability

FORMATION TYPE

- Active control
- Delta V budget for injection
- Total orbital lifetime

FORMATION TYPE

1



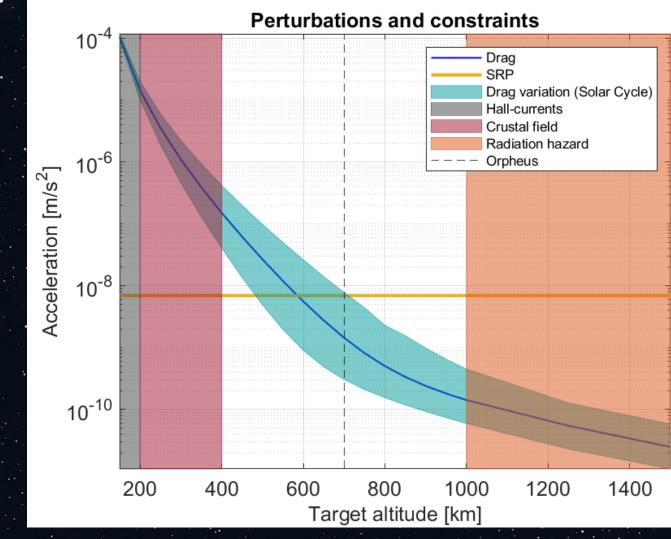
ΔΔ

#### **Selection Criteria:**

- Magnetic field accuracy
   Formation stability
   Active control
   Delta V budget for Injection
- Total orbital lifetime

ORBITHEIGHT

2



45

ORBIT INCLIMATION

3

# □ Global earth coverage

- Ground station coverage
- Variation of local time
- Perturbation strength
- Frequency of launches
- Revisit period





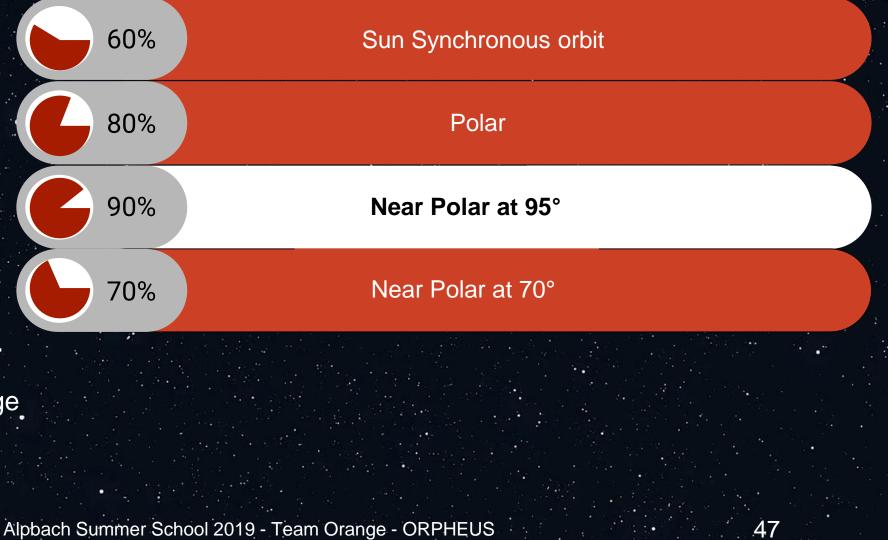
46

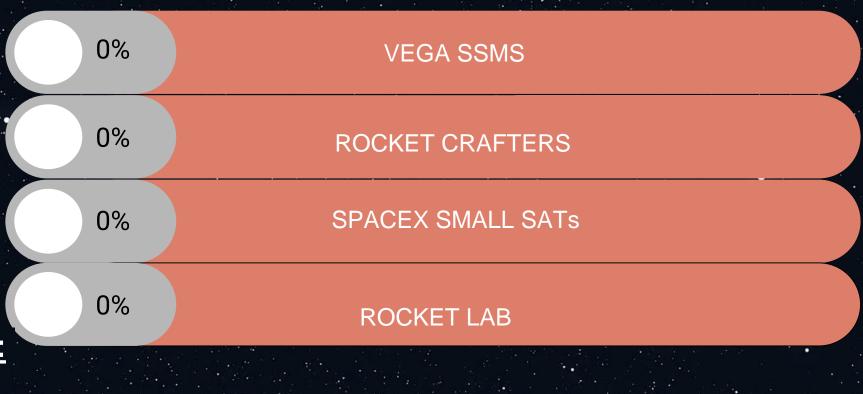
ORBIT INCLINATION

3

### Selection Criteria:

Global earth coverage
 Ground station coverage
 Variation of local time
 Perturbation strength
 Frequency of launches
 Revisit period





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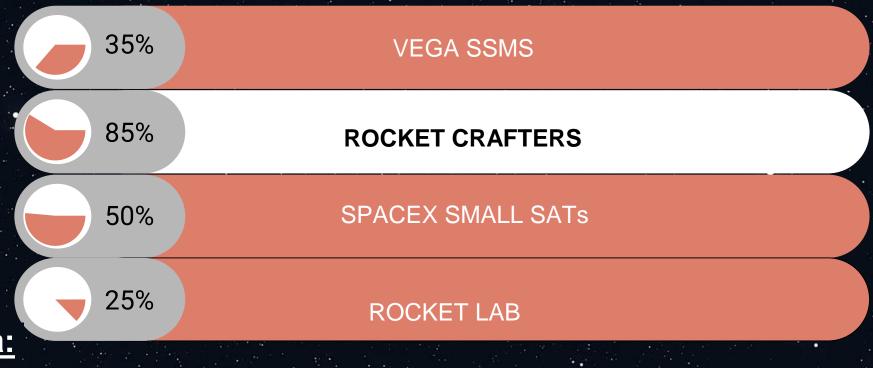
#### **Selection Criteria:**

- □ Orbital inclination
- □ Flexibility
- Reliability
- Cost

LAUNCHER

4

□ Altitude



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- Selection Criteria:
- □ Orbital inclination
- Flexibility
- Reliability
- Cost

LAUNCHER

4

□ Altitude

# **ORPHEUS FINALIZED CONFIGURATION**

ORBIT HEIGHT = TOONNI

2

CARIWHEEL HELIX

1

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3

ORBIT INCLINATION - 95 degrees

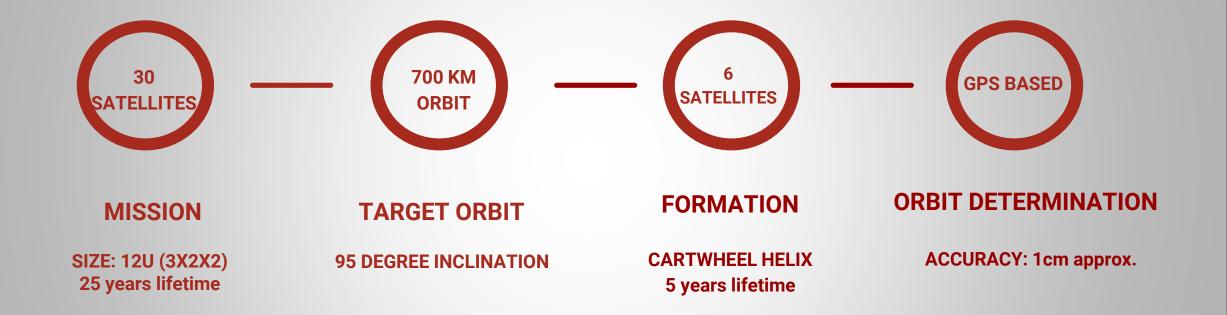
ROCKET CRAFTERS

50

AUNCHER

4

## Mission Profile Overview





#### MISSION PHASES - SINGLE CONSTELLATION

Launch **Orbit insertion** Commissioning **Science operations** . Orbit maintenance **Constellation handover** Disposal **Data analysis** 

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### MISSION PHASES - SINGLE CONSTELLATION

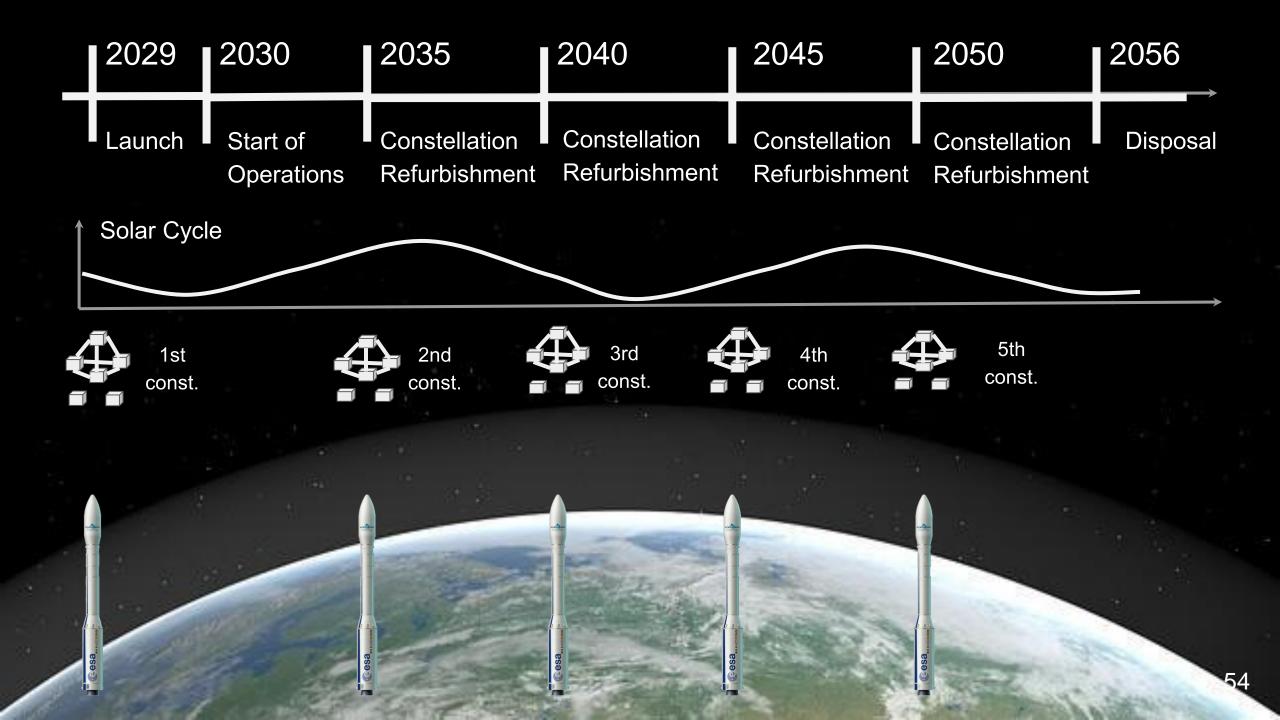
-SATELLITE -DEPLOYMENT COMMISSIONING

FORMATION ACQUISITION

ORBIT INSERTION

DISPOSAL

LAUNCH



### **Spacecraft Deployment and Constellation Formation**

2

3

4

5

6

**6 SATELLITES SINGLE** 

**DEDICATED LAUNCHER** 

REDUCE MEAN ECCENTRICITY OF ALL SATELLITES

TEMPORARY INCLINATION CHANGES TO CREATE RAAN DRIFT

INTRODUCE OFFSETS TO ARGUMENT OF PERIGEE

FINALIZE ORBIT BY PERFORMING PHASE MANEUVERS

ACQUISITION DELTA V = 68 m/s

#### Courtesy of Marcus Hallmann

# Delta V Budget

	Delta V (m/s)	Fuel Mass (grams)
Injection	54-68 (per satellite)	1188.77
Collision Avoidance	~1	~17.48
Formation + station keeping	~40 (per satellite per year)	~3500
Deorbit	~51.33 (per satellite)	~897.35
Total	~320.33 m/s (5 year lifetime)	~5600

# Launch & Transfer

#### **Requirements and Constraints**

41		
1.1.1	No of spacecraft per launch	6 satellites (12U each) -> 72U
	Satellite mass	20kg per satellite
	Satellite dimensions	30 cm x 20 cm x 20 cm per satellite
	Mission orbit	e≈0 a=700km i=95°
	Mission timeline	25 years
	Minimum vibration	31.5 Hz
	Launcher C3	-0.00051234

## Mission Profile Advantages

- Schedule flexibility in handover between deployments
- Incorporation of lessons learned in subsequent deployments
- Cost savings in mass production of identical spacecraft







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## Scientific Payload Identification

STATE OF THE ART MAGNETOMETERS ACCURATE VECTOR & SCALAR MEASUREMENTS OF THE GEOMAGNETIC FIELD	VECTOR FIELD MAGNETOMETERS ACCURATE DETERMINATION OF ATTITUDE & MECHANICAL STABILITY	PAYLOAD SUITABLE MASS & VOLUME FOR A MICRO-/CUBESAT	MAGNETIC CLEANLINESS & PHYSICAL ISOLATION ALLOWS FOR MORE ACCURATE READINGS

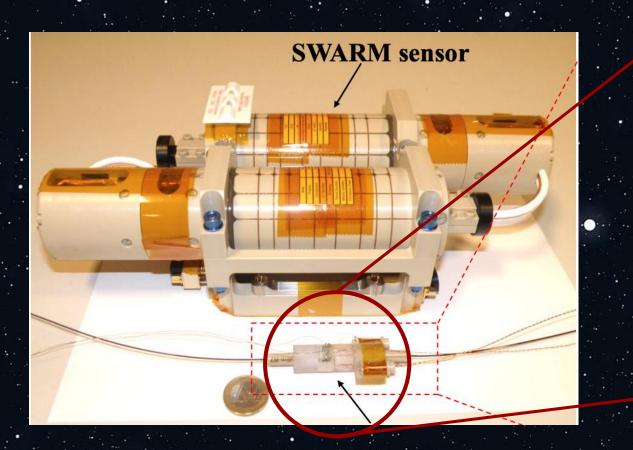
# Payload Requirements

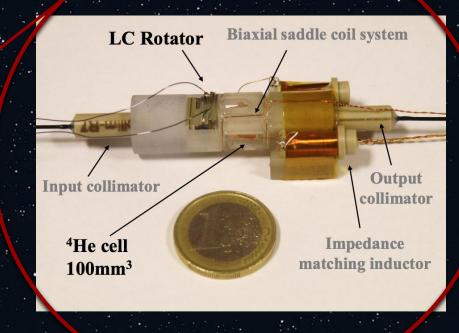
ID number	Short Descriptive Text	Source of Requirement	Object text
INSTR-01	Scalar magnetometers accuracy	SO-02	The scalar magnetometers shall measure the magnitude of magnetic field with an accuracy of 0.3nT.
INSTR-02	Measuring vector field	SO-01	Spacecraft shall use an instrument capable of measuring magnetic vector field.
INSTR-03	Vector magnetometers accuracy	SO-01	The vector magnetometers shall measure all three components of the magnetic field vector with an accuracy of 1 nT.
INSTR-04	Attitude accuracy	SO-01	Spacecraft shall be able to determine its attitude with an accuracy of 0.1 degree to precisely determine magnetic vector field.
INSTR-05	VFM and star trackers alignment	SO-02	Vector field magnetometer and star trackers responsible for determining the VFM attitude shall be aligned to 1 arcsec.
INSTR-06	EM interferences level	SO-01	Electromagnetic interferences measured by magnetometers shall be lower than 1nT.

## Miniaturised Vector Field Magnetometer

#### **TRL 9 Vector Field Magnetometer Performance** Pickup X a) b) Miniaturized fluxgate sensor Feedback Z Ring Core Y,Z Ring Core X,Z Used in conjunction with star trackers (precise attitude measurement) Achievable **1nT** absolute accuracy w/ independent sensor calibration Pickup Z Resolution: 3 pT (24bit) Noise: 10 pT/√Hz Feedback X Feedback Y Pickup

## Miniaturised Scalar Magnetometer

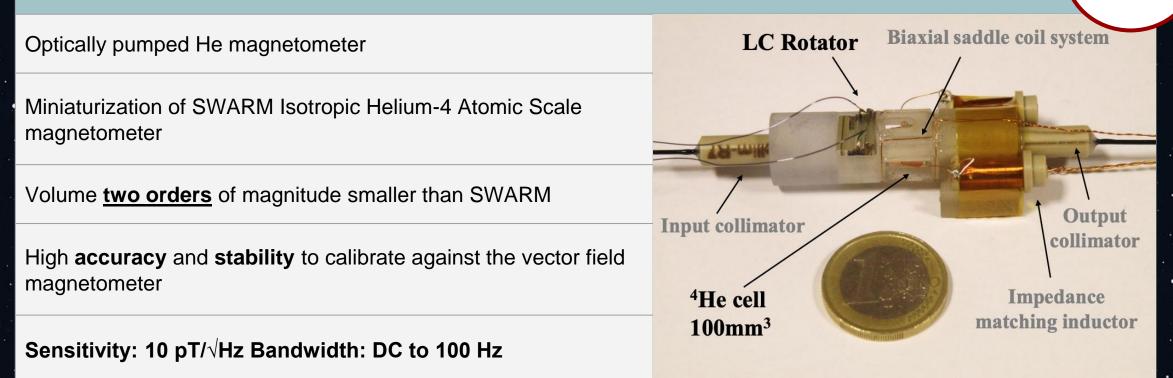




## Miniaturised Scalar Magnetometer

#### **Scalar Magnetometer Performance**

TRL 4



## Magnetometer Boom

#### **Boom Performance**

TRL 6

Vector field magnetometer is placed on a <u>deployable 1.2 meter</u> <u>boom</u> with two star trackers

Scalar magnetometer is placed on a separate 1.2m boom

Magnetic cleanliness programme and data cleaning allow for for further improvements and can compensate for a **shorter boom** 

Expecting to require a **<u>dedicated development programme</u>** to adapt boom for our specific usage needs





## Magnetometer Boom

AstroTube<sup>™</sup> Boom for SmallSat Applications

Oxford Space Systems-Astrotube Boom (1.5m)

050

TRL 9

Roccor- Cubesat ROC Boom (1.5m)

**TRL 7** 

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#### **Star Tracker Performance**

CubeSat compatible star trackers are used for **precise attitude determination** 

The star trackers are arranged with the **boresights at 90°** from each other to improve the accuracy

Accuracy: 6 arcsec Dimensions: 30x30x38mm

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**TRL 9** 

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## Overview of Subsystems

Structure Power

Propulsion

ADCS

Thermal

Communications

Ground Segment

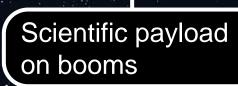
70

### Spacecraft Design & Configuration

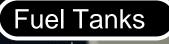
Front



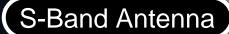




## Bottom



S-Band Transceiver





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Onboard Computer

72



Battery

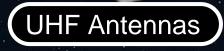
## Middle



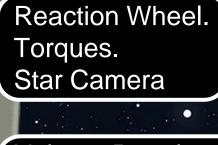
#### Fuel Tanks

#### S-Band Transceiver





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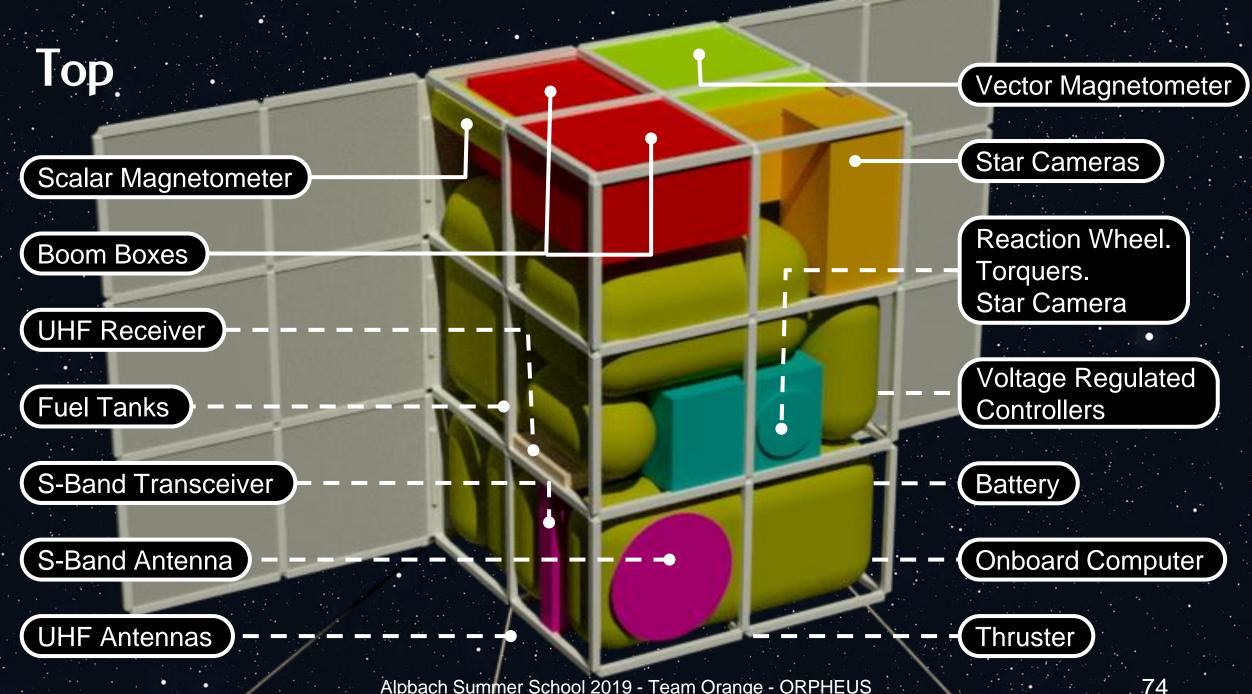
Voltage Regulated Controllers

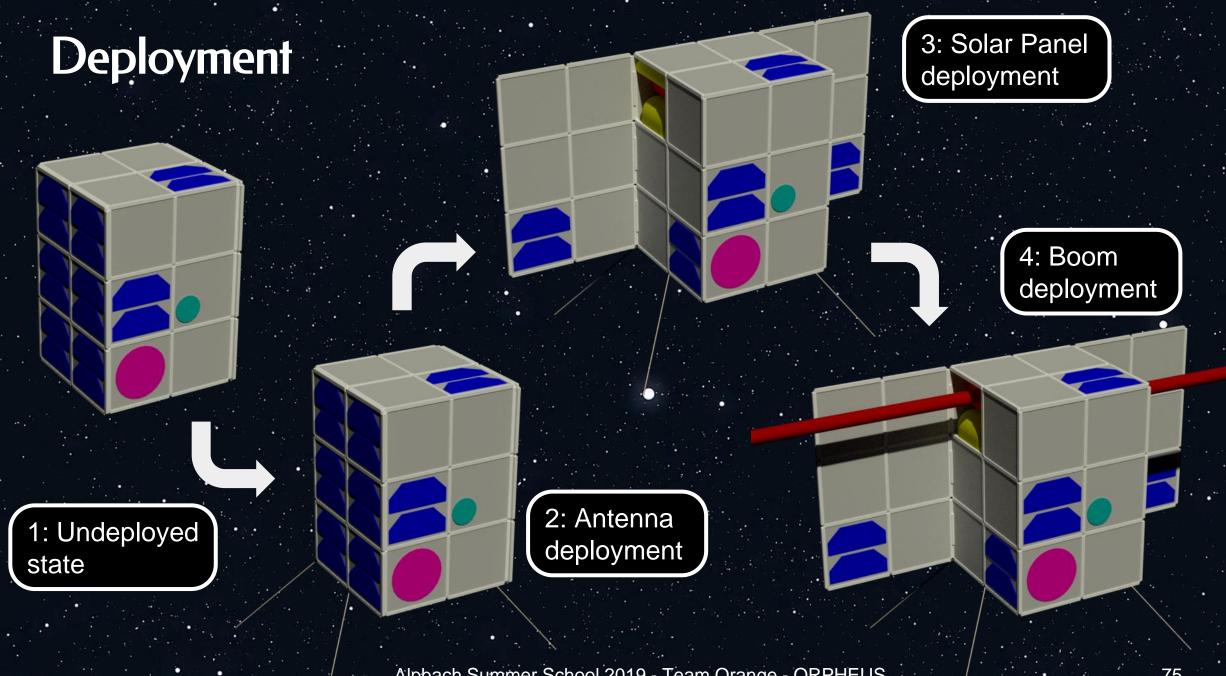


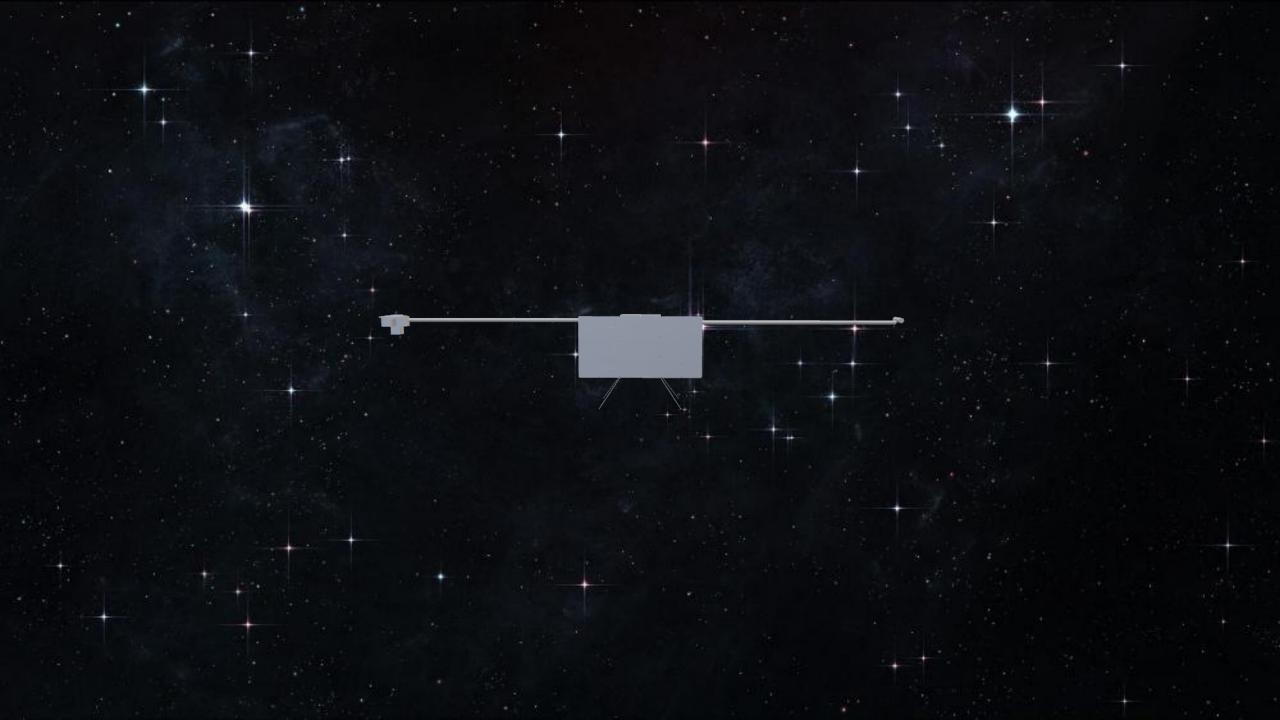


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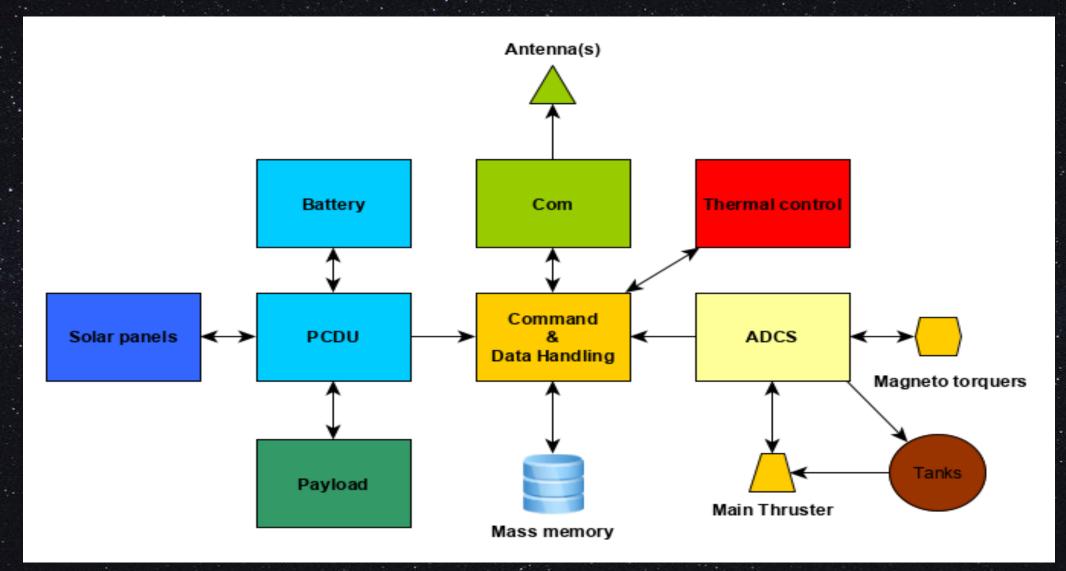








#### System - Overview



#### System - Power Budget

Power Usage [W]									
Operational mode	ADCS	Electrical	P/L (Science)	Propulsion	Telecomms	Thermal Control System	TOTAL	Total w/ margin	
Safe mode	2.03	2.95	0	0	6.05	0.15	11.18	12.66	
Commissioning	3.38	13.95	0	0	6.05	0.15	23.53	26.93	
Science	3.30	13.95	2.8	0	0	0.15	20.20	23.26	
Orbital maintenance	3.30	13.95	0	5	0	0.15	22.40	25.68	
Telecom	1.43	13.95	0	0	6.05	0.15	21.58	24.59	
Science + heating	3.30	2.95	2.8	0	0	10	19.05	21.93	
Margin for subsystem	20%	15%	10%	10%	10%	15%	20%		

#### System - Power



Active attitude control provides a stable sun-facing spacecraft panel (0.18 m<sup>2</sup>) generating 51 W of usable electrical power. Battery size 40 Wh
Power Control Unit
Voltage regulated buses

## System - Propulsion



#### **SELECTED SYSTEM:**

5.6 kg propellant based on ~12.08 kg ceiling for total satellite dry mass

TRL 6

#### Tank volume sizing: ~ 0.69 m<sup>3</sup> ~ 0.32 MPa

Propulsion systems considered:

Resistojet
Cold gas
Electrospray

TRL 9

#### • Drivers:

- Total delta-v
- Magnetic cleanliness
- Total burntime over 5 years
- Propellant mass

#### System-Attitude Determination & Control

**DESIGN DRIVERS** 

Solar Incidence & Power Budget

> **Orbit Maintenance & Formation Keeping**

**Structural Vibrations** (Boom)

**Magnetic Cleanliness** (Science Operations)

### System-Attitude Determination & Control

#### **OPERATIONAL MODES**

#### **Telecommunications**

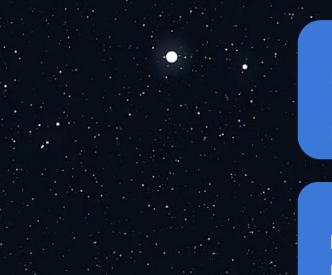
- Ground station communication
- Active transceivers and deactivated magnetometers.

#### **Science & Heating**

- Notably significant low temperatures
- Additional heating for key subsystems

#### **Orbital Control**

- Orbit acquisition & formation keeping.
- □ Inertial stabilisation for propulsion.



#### Acquisition & Safe

- After deployment/failure of a subsystem.
- □ Coarse Sun pointing.

#### Commissioning

Solar panels & booms deployment with instruments' calibration.

**G** Fine Sun pointing.

#### **Science**

Magnetometer data acquisition.
 Fine attitude knowledge required.

### System-Attitude Determination & Control

#### **ADACS Control Board**

#### TRL 9

#### Attitude determination:

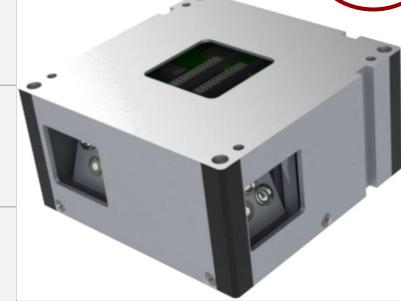
- →2 Sun sensors (Coarse pointing + Safe mode)
- →2 Star trackers (Fine pointing)

#### **3-axis attitude control:**

- →4 Reaction wheels (Redundancy)
- →4 Magnetorquers
- Desaturation of wheels Detumbling & safe mode

#### **Desaturation**

- → ~ 2 days (duration < 1 min)</p>
- → Coordination with science operations for reduced impact



#### MAI-400 Maryland Aerospace. Inc



#### Thermal Control Analysis

 $\left( + \right)$ 

Identify hardware temperature limit

Single node analysis made for worst case scenarios for a 700 km orbit.

Worst Case Scenario Hot conditions ~308K Cold conditions ~242K

Spacecraft will need heating system and extra insulation for critical subsystems

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Nominal operations+20 to +40 °CWorst case envelope0 to +50 °C

### System - Thermal Control

#### **Thermal Control Procedures**

- → Active attitude control ensures a "cold side" of the spacecraft
- → A radiative surface for thermal control is not needed
- → Electrical heaters are included for emergency heating in case of safe-mode
- → Monitoring temperatures of critical components using temperature sensors
- → Extra MultiLayer Insulation (MLI) will be included for the battery, onboard computer, electronics and propellant tank.



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**TRL 8/9** 

### **Communications - Overview**

PRODUCTION

On-board storage (4GB) is not a limiting factor

Scientific payload: 128 bit/s @ 1Hz TT&C: 60bit/s 10Hz for sensors & 1 Hz GPS

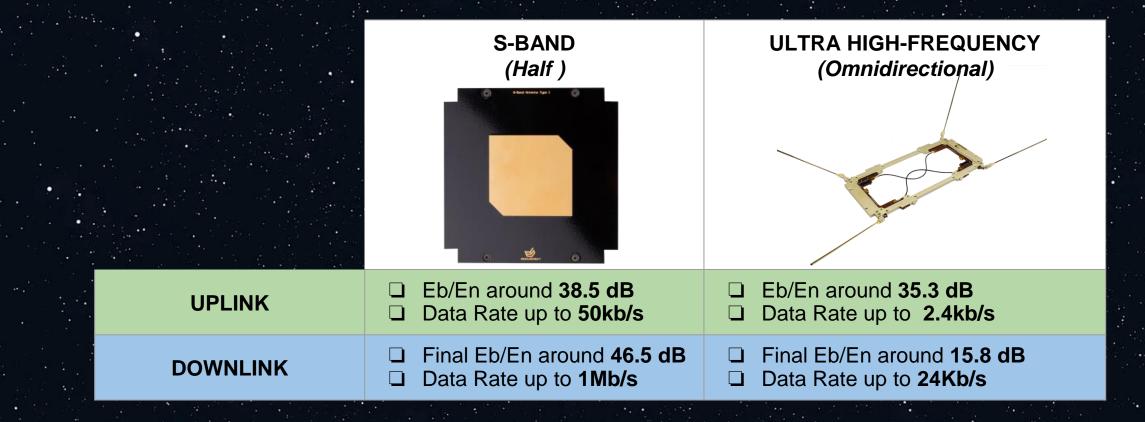
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23 Mbit /satellite/day 140 Mbit /constellation/day

3 years until storage saturation

20% MARGIN

#### **Communications-Link Budget**



#### Communications – Ground Segment



- Specially designed for small sat missions
- □ Control center integrated
- □ 4 Usable passes/Day >12 min
- □ Existing Ground link to <u>KIR\_2</u>

#### CONTROL CENTRE-ESOC

Redu

Cebreros

Kourou

Santa Maria

aüe





Existing Ground link to <u>ESOC</u>
 10 Usable passes/Day >11 min



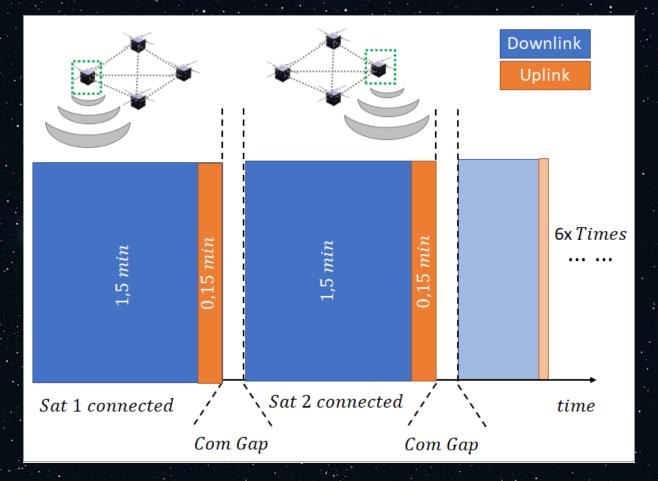
New Norcia

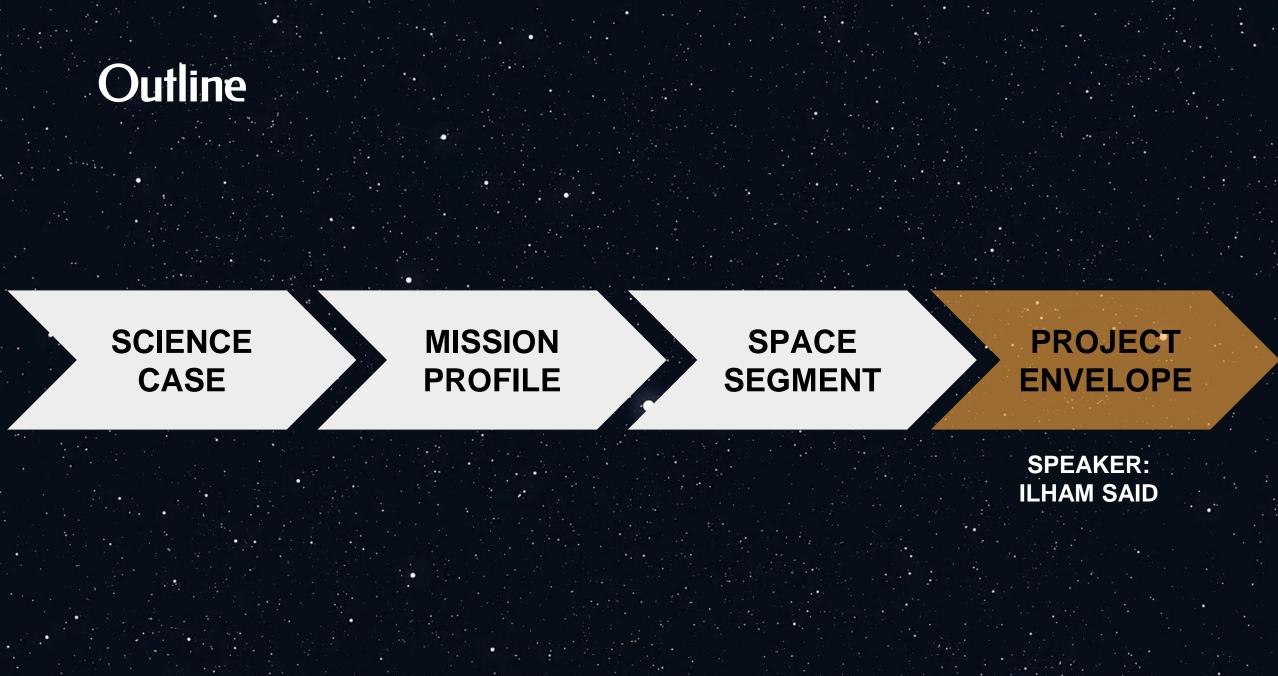
11

#### Communications - Ground station transit timeline

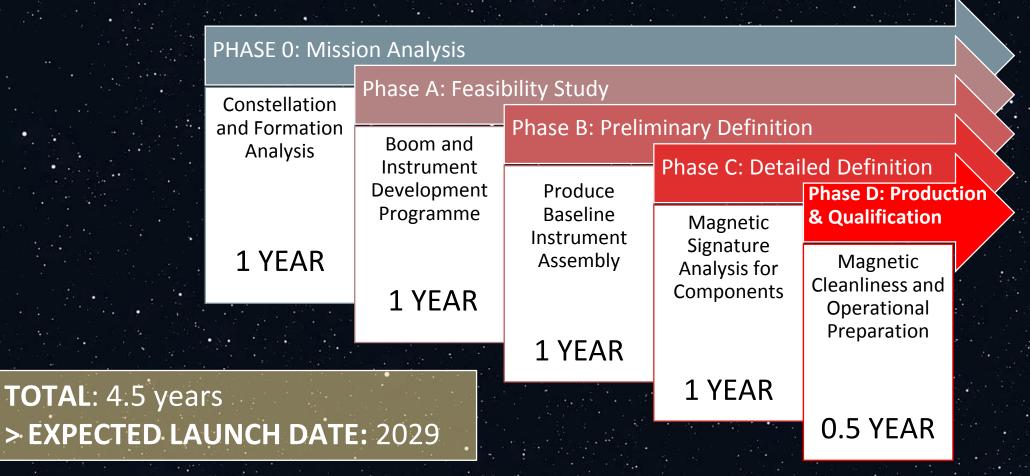
#### Single pass Uplink & Downlink

그는 그는 것 같은 것 같							
S-Band telecom li	Margin						
Revisit time (days)	4						
Downlink S Band (Min)	7.0	20%					
Uplink S Band (Min)	1.7	20%					
Com Gap	0.03	20%					
Time/const (min)	8.8	20%					





### **Development Schedule for Orpheus**



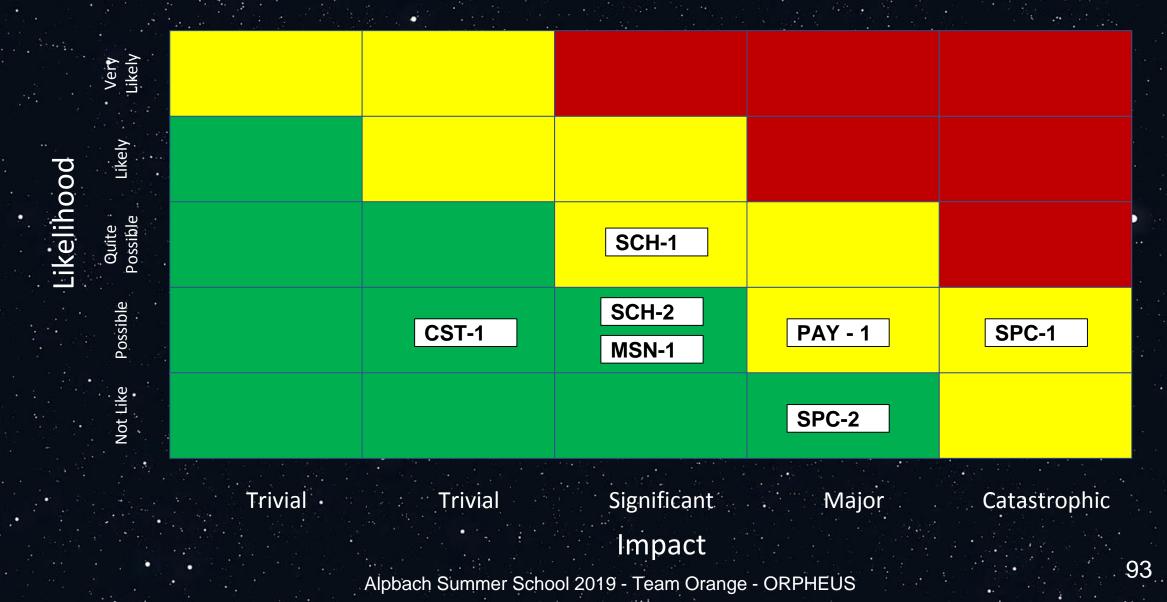
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### Risk Assessment

Risk	Call Sign	Mission Risk	Likelihood	Impact	Mitigation
Schedule	SCH-1	Failure to keep launch schedule due to launcher unavailability	3	3	Allow for sufficient launcher development time and overlap between constellation deployments
Schedule	SCH-2	Low TRL boom concept causes delay of schedule	2	3	Allow for sufficient development time when selecting boom provider
Payload	PAY-1	Structural / deployment failure of boom	2	4	Constellation approach allows for degraded performance when one or two satellites are compromised
Spacecraft	SPC-1	Being unable to communicate with spacecraft	2	5	Test campaign and backup communication scheme
Spacecraft	SPC-2	Failure to meet Cubesat launch standards results in delayed launch and additional dispenser and component costs	1	4	Conformance to Cubesat standards is to be paramount at every stage of the design process
Cost	CST-1	COTS parts no longer available for later satellite generations	2	2	Adherence to Cubesat standards will increase likelihood of compatibility with functionally equivalent successor components
Mission	MSN-1	Faster than anticipated degradation of formation degrades data and shortens spacecraft lifetime due to propellant consumption	2	3	Allow for schedule flexibility for early replacement of constellation

#### **Risk Assessment**



### **Cost Breakdown Estimation**

#### COST BREAKDOWN ESTIMATION

SPACECRAFT ELEMENTS	Price per Spacecraft	Recuring Spacecraft	NRE +5 RS/C			
	Million€					
Structure	0,60	0,48	3,00			
Thermal Control	0,28	0,22	1,40			
ADCS	1,20	0,96	6,00			
Electrical Power supply	0,50	0,40	2,50	_		
Propulsion(Reaction Control)	0,15	0,12	0,75	TOTALS		
Telemetry Tracking and Command	0,60	0,48	3,00	al کا		
CD&H	0,75	0,60	3,75			
Payload	1,20	0,96	6,00			
IA&T	0,20	0,16	1,00			
Engineering, Software	0,50	0,40	2,50			
TOTAL	5,98	4,78	29,90	149,5		
Total S/C including payload+20% Margin				179,4		
Program management	3 person for 5 years 1.5 for 20 years					
Mission Operation (incl GS)	1 h GS per day+1 person					
Science Operation(incl archiving)	3 persons for 5 and 1.5 for 20 years					
Launch	5.4/launch-for 200 Kg total launch mass					
Contigency (20%)						
TOTAL COST				281,3		

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Contribution per ESA Member State Citizen <u>0.54 €</u>

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### **Descoping Options**

 Reducing number of satellites per formation by 1 (still be able to gather data with the help of existing models)

 Reduce handover between launches / accept gaps in data to extend mission duration without compromising long-term goals

 Cut down total mission duration would save on cost (manufacturing and operation)

• Cut down to one boom with a single magnetometer in vector / scalar mode

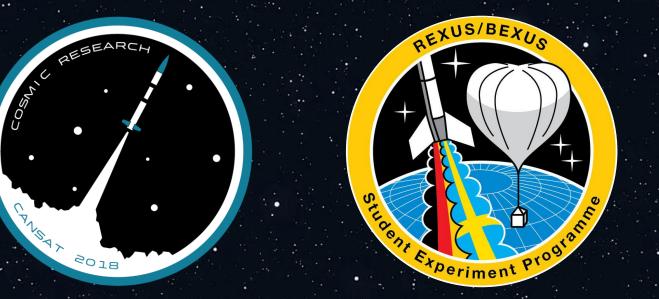
### Outreach - Education!

Magnetometers are a common payload on student payloads ESA Education programs (CanSat. REXUS/BEXUS. Student Cubesats).

□ Students can compare data with ORPHEUS

□ ORPHEUS experts can serve as mentors

□ Promising students can intern with the ORPHEUS science and mission operations team



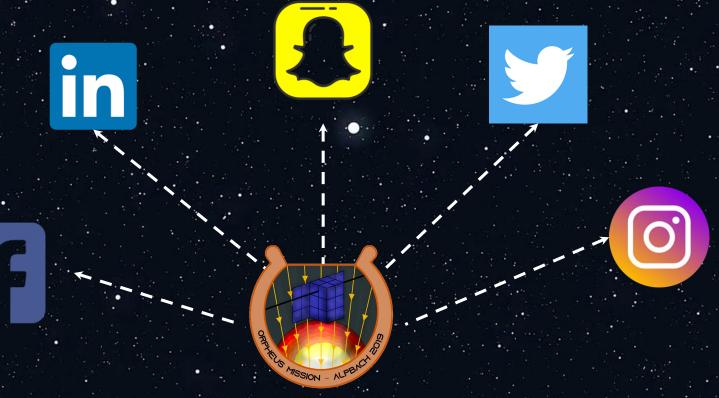


### Outreach - Education!

Orienteering competitions organized on university campuses with prominent geomagnetic research to shine a light on recent developments

### Outreach - Social Media!

Social media presence for general public with emphasis on citizen science (we all live in the magnetic field. let's go out and measure it!)





- Scientists have told us: "We want Swarm to continue!"
- Engineers responded: "We can make it cheaper, longer and better!"
- Orpheus is the solution:
- "Magnetic measurements in an orbit formation designed to focus on the Earth's core. This will enable for sophisticated measurements in an all-new region. with a mission configuration to deliver continuous world class data for years to come" - ORPHEUS 2019

### **Special Thanks!**

To our tutors. lecturers. organizers. and the whole summer school for this fabulous learning opportunity and for all the new knowledge. techniques. and friendships we've gained along the way.

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# **BACKUP-SLIDES**

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### Backup Slide-Past Missions

Ørsted:

- 1 satellite
- mapping of internal and external field

#### Swarm:

- 3 satellites
- mapping of internal and external field

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#### Backup Slide – Power Subsystem

- Deployable solar panels provide 18U of surface area. which translates into ~0.134 m<sup>2</sup> of solar panel area (40mm x 80mm panels)
- Assuming a solar constant of 1361 W/m<sup>2</sup> and a high-TRL multijunction solar panel efficiency of 30 %. we obtain 52.25 W
- Considering a degradation of 0.5 % per year. the available power at the end-of-life is 51 W
- Source: https://sst-soa.arc.nasa.gov/03-power

#### Backup Slide – Power Subsystem

 Eclipses up to 35 minutes on our 99 minute period will require a battery of size 40 Wh to account for a worst-case power consumption of 32 W. assuming a 50% depth-of-discharge as best practice.

Source: https://sst-soa.arc.nasa.gov/03-power

### Backup Slide - Power ADCS subsystem

Perturbations .	Туре	Maximum torque	with 20% margin
Drag	Variable*	<sup>2</sup> 5.46 x 10 <sup>-8</sup> Nm	6.55 x 10 <sup>-8</sup> Nm
Gravity gradient	Cyclic*	<sup>1</sup> 3.45 x 10 <sup>-6</sup> Nm	4.14 x 10 <sup>-6</sup> Nm
Solar pressure*	Secular*	<sup>2</sup> 4.37 x 10 <sup>-8</sup> Nm	5.36 x 10 <sup>-8</sup> Nm

\*Sun pointing (x-axis)

<sup>1</sup> Assuming uniform density for 12U cuboid + 2 booms + optical payload. I = [2.163. 0.195. 0.120] kg/m<sup>2</sup> <sup>2</sup> Assuming center of gravity to center of pressure (SRP & Drag) distance  $\sim$  10 cm

			· · ·		
			0.3 m		
(		•			
				1.2 m	
		0.2	m		

### Backup Slide - Power ADCS subsystem

Momentum storage	11.076 mNms
Maximum Torque	0.635 mNm
Magnetic dipole moment	0.108 Am <sup>2</sup> (0.15 @72% Duty Cycle)

\*Integrated ADACS Stock #: MAI-400-31130200 Maryland Aerospace. Inc.



ΔT between desaturations\*

Desat. time

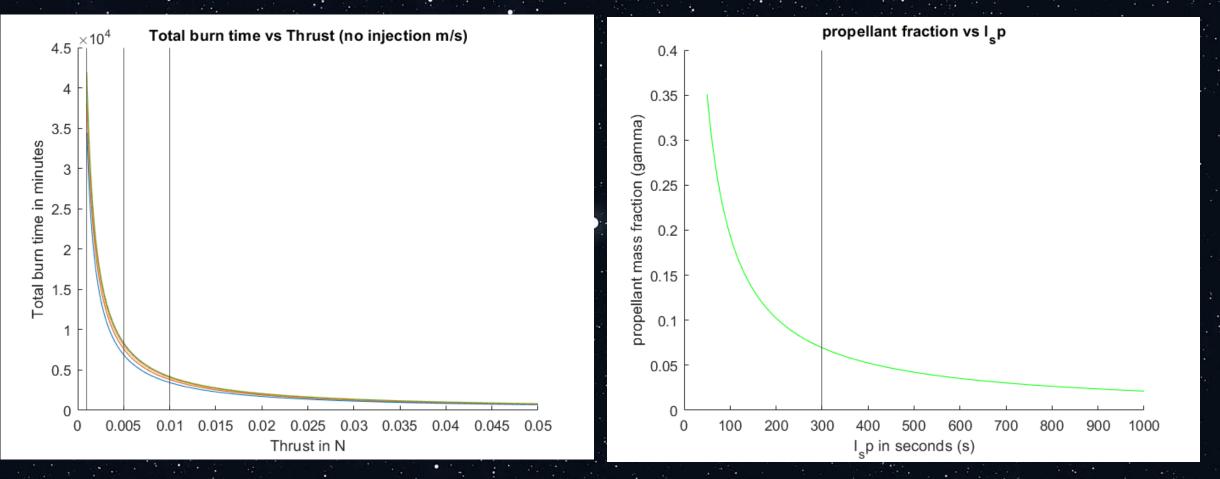
~ 10 sec

~ 40 hours

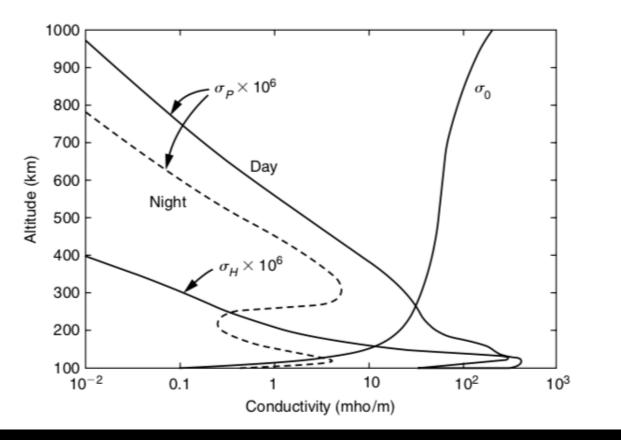
(24h target for operations)

\* Including allocated 4 mNms momentum storage for GG compensation and slew manoeuvering.

#### Backup slide - propulsion



### Backup Slide - Hall Current



- Pedersen and Birkeland currents are still contributing however
- Hall currents are of minor importance

With courtesy of M. Kelly

### Backup Slide - Spherical Harmonics 1

- Spherical harmonics are a series of special functions defined on the surface of a sphere
- Solution to the Laplace equation:

 $\nabla \times \mathbf{B} = 0 \rightarrow \mathbf{B} = -\nabla U$ , where U is the scalar potential.

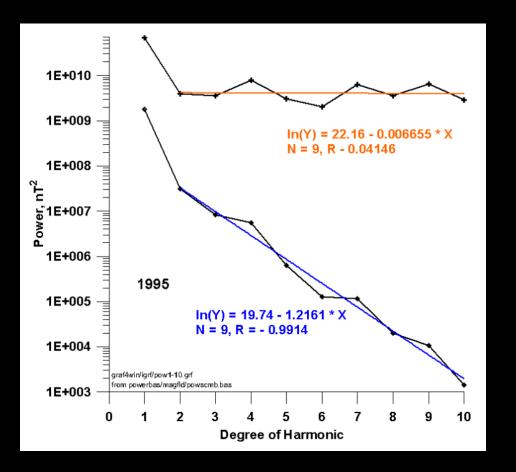
 $\nabla \cdot \boldsymbol{B} = 0 \rightarrow \nabla^2 U = 0$  (Laplace equation)

There are two types of solutions:

- potential  $U_i$  due to sources internal to the Earth  $(r < R_E)$ 

- potential  $U_e$  due to sources external to the Earth  $(r > R_E)$ such that  $U = U_i + U_e$ .

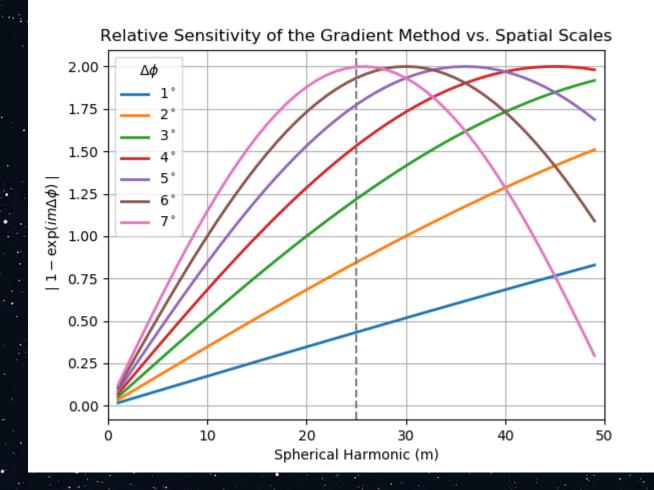
#### Backup Slide - Spherical Harmonics 2



 Power of Earth's 1995 field as a function of harmonic degree. shown for surface field (blue) and CMB field (orange)

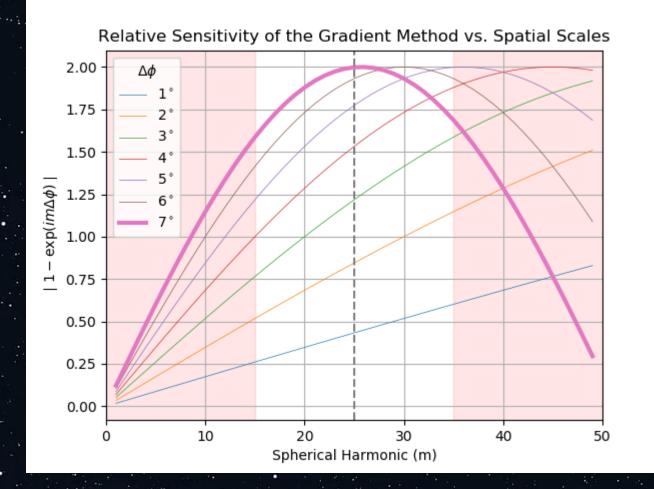
With courtesy of C. Harrison

### Backup Slides - Spherical Harmonics 3



Measure core dynamics to degree m = 25 $\rightarrow$  Satellite separation of 7° horizontally

### **Backup Slides - Spherical Harmonics 4**



Measure core dynamics to degree m = 25 → Satellite separation of 7° horizontally = 860 km (time separation ~2 min)

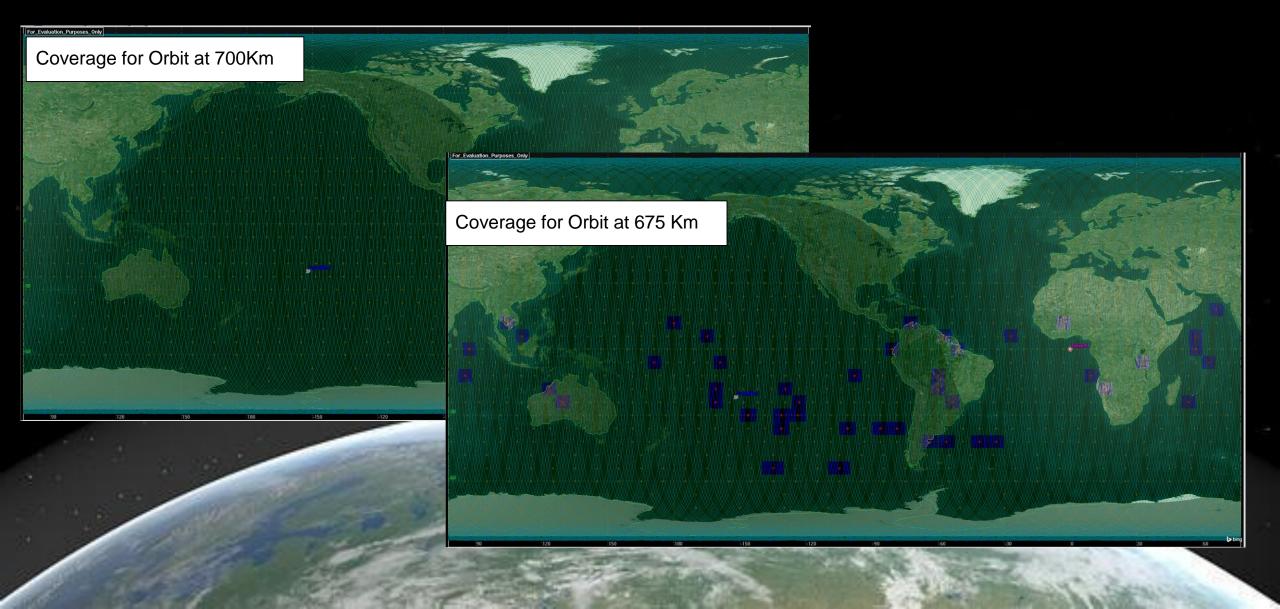
#### Backup Slide – Mass Budget

Subsystem	Element	Units	Mass (kg)	Total Mass (kg)	Margin	Final Mass (kg)	
	Magnetorquer	4	0,176				
	On-board computer	1	0,010				
ADCS	Reaction wheel	4	0,192	0,464	5%	0,487	
	Star tracker	2	0,080				
-	Sun sensor	2	0,006				
	Battery	1	0,335				
Electrical	PCU	1	0,074	1,545	5%	1,622	
Liccultar	Solar cell	40	0,792	1,545	570	1,022	
	Voltage regulated buses	NA	0,344	•			
	Boom	2	2,000				
Payload	Scalar Magnetometer	1	0,300	2,675	20%	3,210	
	VFM	1	0,375	•			
	Propellant	1	1,244				
Propulsion	Tank	1	0,456	7,500	15%	8,395	
-	Thruster	1	5,600				
Structure	Frame	1	2,000	2,000	20%	2,400	
	S-BAND Antenna	1	0,064		10%		
Telecommunications	S-BAND Transceiver	1	0,075	0,253		0,278	
relecommunications -	UHF Antenna	1	0,090	0,235	10%	0,278	
	UHF Transceiver	1	0,024				
	Heater	1	0,005				
Thermal	MLI	NA	0,153				
Inermai	Temperature sensors	2	0,080	0,257	10%	0,283	
	Thermistors	2	0,024				
Total Mass	16,80 kg						
System margin	System margin $1120\%$ $20\%$ $2140$ T						
Final mass	Final mass Alphach Summer School 2019 - Team Orange -						
	, 0	<u>ORP</u>	HEUS				

### Backup Slide - TRL

Subsystem	Element	TRL				
	Magnetorquer	7				
	On-board computer	9				
ADCS	Reaction wheel	9				
	Star tracker	9				
	Sun sensor	9				
	Battery	9				
Eletrical	PCU	7				
Licuical	Solar cell	9				
	Voltage regulated buses	9				
	Boom	6				
Payload	Scalar Magnetometer	4				
	VFM	9				
	Propellant	9				
Propulsion	Tank	6				
	Thruster	9				
Structure	Frame	7				
	S-BAND Antenna	9				
Telecommunications	S-BAND Transciever	7				
relecommunications	UHF Antenna	7				
	UHF Transceiver	7				
	Heater	9				
Thormal	MLI	7				
Thermal Temperature sensors						
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ORPHEUS						

### Orbit -Backup Slide



### Communications - Backup Slide

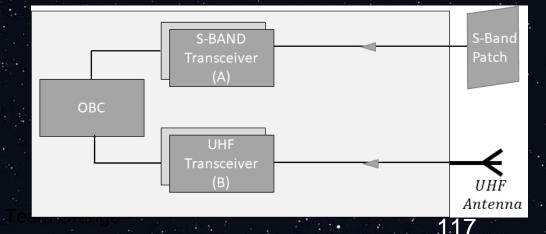
20%

#### **Data Production**

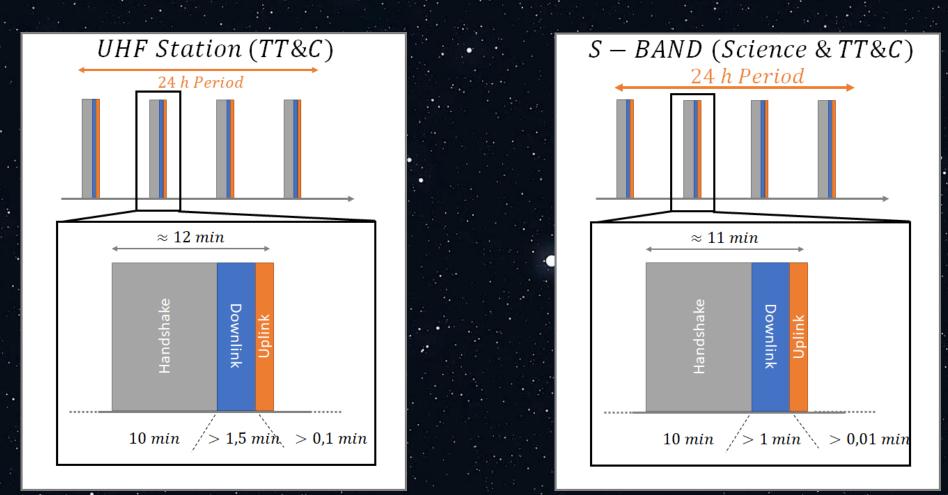
Days to Mem Sat

Data Generation		Packet Size	Mes Rate (Hz)	
Scalar Mag. 1 (bit/	s)	32	1	
Mag. 2 (bit/s)		96	1	
TT&C (bit/s)		60	1	
Total (bit/s)		225.6	20% Margin	
	•			
Data Output		Margin		
Mbit/Day/Sat 23		20%		
Mbit/Day Const	117			

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### Communications (Backup Configuration)



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### Spacecraft Operational Modes- Back-up

- Measurements require magnetic cleanliness; not possible during:
  - orbit maintenance (monthly)
  - $\circ$  reaction wheel desaturation (weekly)

 All measurement downtimes will be coordinated with the PI and scientific planning team to minimize impact on long-term science goals
 e.g. reaction wheel desaturation over poles